

NANOINDENTING THE CHELYABINSK METEORITE TO LEARN ABOUT IMPACT DEFLECTION EFFECTS IN ASTEROIDS

Studying Chelyabinsk for impact deflection

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ABSTRACT

The Chelyabinsk meteorite is a highly shocked, low porosity, ordinary chondrite, probably similar to S- or Q-type asteroids. Therefore, nanoindentation experiments on this meteorite allow us to obtain key data to understand the physical properties of near-Earth asteroids (NEAs). Tests at different length scales provide information about the local mechanical properties of the minerals forming this meteorite: reduced Young's modulus, hardness, elastic recovery, and fracture toughness. Those tests are also useful to understand the potential to deflect threatening asteroids using a kinetic projectile. We found that the differences in mechanical properties between regions of the meteorite, which increase or reduce the efficiency of impacts, are not a result of compositional differences. A low mean particle size, attributed to repetitive shock, can increase hardness, while low porosity promotes a higher momentum multiplication. Momentum multiplication is the ratio between the change in momentum of a target due to an impact, and the momentum of the projectile, and therefore higher values imply more efficient impacts. In the Chelyabinsk meteorite the properties of the light-colored lithology materials facilitate obtaining higher momentum multiplication values, compared to the other regions described for this meteorite. Also, we found a low value of fracture toughness in the shock-melt veins of Chelyabinsk, which would promote the ejection of material after an impact and therefore increase the momentum multiplication. These results are relevant in the context of a future mission to test asteroid deflection, currently being studied by ESA and NASA: the Asteroid Impact and Deflection Assessment (AIDA) mission.

TEXT

1. INTRODUCTION

Although the probability of an asteroid causing a catastrophic impact is statistically small (Atkinson et al. 2000), the public concern about impact hazard increased in 2013, when a small asteroid overflew the Russian region of Chelyabinsk, producing a large airburst accompanied by the fall of thousands of meteorite specimens, with a total mass of ~1000 kg (Nazarov et al. 2013; Brown et al. 2013). The asteroid diameter, estimated as ~18 meters (Brown et al. 2013), was relatively small compared with km-sized bodies capable of produce a mass extinction, but still the shockwave released caused significant damage to buildings, and ~1500 people were injured. The Chelyabinsk event shows that even if a highly destructive impact has a very low probability of occurrence, airburst effects can still be dangerous (see e.g. (Wasson 2003)). Such events can be expected on a once on a decade-to-century scale (see e.g. (Atkinson et al. 2000)) due to collisions with objects coming directly from the main asteroid belt, or also to disruptive processes occurred in near-Earth space (Trigo-Rodríguez et al. 2007).

There is an ongoing discussion concerning the best strategy to deal with any potential threatening asteroid (Morrison 2010). Some techniques, such as using a gravity tractor, require years or even decades to be effective (Lu & Love 2005). Kinetic impact strategies, which imply using a projectile to slightly change the orbit of a near-Earth Asteroid (NEA), are technologically more advanced and require a much shorter time scale (Ahrens & Harris 1992). Due to the controversy related to the use of nuclear weapons, non-explosive projectiles are preferred (Koenig & Chyba 2007). Between 2005 and 2007 the European Space Agency (ESA) proposed the Don Quijote mission (Carnelli et al. 2006), with the aim to test the feasibility of using a kinetic projectile to deflect an asteroid, and also to properly observe and analyze the consequences on the target asteroid. The mission was not adopted, but aspects of it are incorporated in the newly proposed Asteroid Impact and Deflection Assessment (AIDA) mission. AIDA is

a collaboration between ESA and NASA to develop two complementary spacecrafts: the Asteroid Impact Mission (AIM), by ESA, and the Double Asteroid Redirect Test (DART), by NASA (Michel et al. 2015a, 2016). The two missions are planned to travel to the binary NEA (65803) Didymos, composed of a primary 800 m asteroid and a 150 m satellite. The latter will be impacted by the 300 kg DART spacecraft, while AIM will characterize the system before and after such event (Michel et al. 2015a).

The success of AIDA and similar concepts highly depends on the knowledge of the physical (i.e., mechanical) properties of the NEA to be deflected. A proper characterization of these objects is therefore required to avoid, or minimize, unexpected outcomes such as a multiple fragmentation (Holsapple & Ryan 2002). Here we present a laboratory approach using meteorite specimens in order to quantify, in controlled small scale experiments, mechanical parameters that might be used to predict the effects caused by a projectile on the surface of an asteroid. Nanoindentation is selected here as an almost non-destructive technique, compared to impact tests. The Chelyabinsk meteorite was selected as a good example of the different materials that form small NEAs such as (65803) Didymos. The results obtained here using quasi-static conditions from a Chelyabinsk specimen, are interpreted in view of their correlation with dynamic mechanical parameters that play a role during an impact between the asteroid and an external body.

2. RATIONALE FOR SAMPLE SELECTION AND METHODS

Chelyabinsk has been classified as an LL5 and LL6 ordinary chondrite (OC) breccia with an S4 shock stage and exhibiting different lithologies (Nazarov et al. 2013; Kohout et al. 2014; Bischoff et al. 2013). The most abundant is the light-colored lithology,

constituting ~65% of the meteorite. It shows a typical equilibrated chondritic texture, exhibits an intermediate shock state, and contains recrystallized chondrules that are deformed or broken, plus very thin inter-granular metal and troilite veins (Nazarov et al. 2013; Kohout et al. 2014; Galimov et al. 2013; Righter et al. 2015). Some LL6 fragments with rare chondrule relicts and highly recrystallized LL5/6 or LL6 regions, with and without shock veins, have been classified as part of the light-colored lithology, but also as different lithologies (Bischoff et al. 2013). The following most common lithology is the dark-colored (or shock-darkened) one, in which only a small fraction of the original equilibrated chondritic texture remains. This lithology contains a much larger amount of inter- and intra-granular thin melt veins of opaque material (metal and troilite) due to shock mobilization (Kohout et al. 2014; Bischoff et al. 2013; Galimov et al. 2013; Righter et al. 2015; Reddy et al. 2014). The light- and dark-colored lithologies are rarely found in the same specimen (Kohout et al. 2014). A third lithology is often encountered together with any of the other two: a fine-grained dark impact-melt lithology containing finely dispersed metal and troilite droplets, variable abundances of mineral and lithic clasts, but no high-pressure phases (Kohout et al. 2014; Bischoff et al. 2013; Galimov et al. 2013; Righter et al. 2015). Due to their similar color, the dark-colored and the impact-melt lithologies have often been considered together as a single lithology (Nazarov et al. 2013; Reddy et al. 2014). The three lithologies have roughly similar compositions (Kohout et al. 2014; Galimov et al. 2013), with olivine (~Fa₂₈), strongly affected by shock, being 2 to 4 times more abundant than pyroxene (~Fs₂₃) (Nazarov et al. 2013; Galimov et al. 2013). Although both minerals are mostly homogeneous in composition (Nazarov et al. 2013; Kohout et al. 2014; Galimov et al. 2013; Righter et al. 2015), pyroxene is mainly orthopyroxene, in a proportion superior to 2:1 over clinopyroxene (Nazarov et al. 2013; Galimov et al. 2013). Small anhedral

plagioclase (~Ab86) grains also show the consequences of shock, and indeed in the dark-colored lithology the isotropy of plagioclase is complete (Nazarov et al. 2013; Galimov et al. 2013). Opaque minerals consist of 6-7 wt% of troilite and 2-4 wt% of metal phase, the last being mostly kamacite (~5 wt% of Ni) and taenite (~35 wt% of Ni) (Nazarov et al. 2013; Galimov et al. 2013; Popova et al. 2013). Chromite, phosphate (apatite) and ilmenite, among others, are accessory minerals (Galimov et al. 2013).

As an LL-type ordinary chondrite, Chelyabinsk can be easily connected with most NEAs, usually associated with S- or Q-class asteroids (Reddy et al. 2014; Gaffey 1976; Binzel et al. 2001; Vernazza et al. 2008). In fact, it has been suggested that ~2/3 of NEAs show a better match with LL chondrites than with the other ordinary chondrite types (Vernazza et al. 2008). It has been found that S- and Q- asteroids probably form as rubble piles (Holsapple 2001) and therefore should show a considerable degree of shock and brecciation, similarly to what has been seen in the Chelyabinsk meteorite (Nazarov et al. 2013; Bischoff et al. 2013). Didymos, in particular, has been classified as an Apollo-class asteroid related to the S-complex, and it also has been spectroscopically connected to L/LL-type meteorites (Dunn et al. 2013). As a binary system, the formation of Didymos as a rubble pile is logical (Walsh & Richardson 2006). As well as being a good analogue for Didymos, Chelyabinsk also allows testing the effects of shock on the mechanical properties of OC-forming minerals. Such chondrites originated from the catastrophic disruption of moderately large asteroids, whose fragments formed families with complex collisional histories (Michel et al. 2001, 2015b; Bottke Jr. et al. 2015).

We analyzed one of the specimens of Chelyabinsk, the polished thin section shown in Figure 1 (see Appendix). The microstructure of the sample was studied by optical microscopy and scanning electron microscopy (SEM), and the chemical composition of

the different regions was determined by Energy Dispersive X-ray spectroscopy (EDS). The mechanical properties of the specimen were evaluated by nanoindentation (see Appendix) . To extract the local nanomechanical properties of the sample, indentations with a maximum applied force of 20 mN were performed on the different mineral phases comprised in this OC. The results served to test the consistency between our measurements and the ones obtained from previous studies, and as a reference for mechanical properties assessed using higher applied forces. We were also interested in the average mechanical properties of the material composing the Chelyabinsk meteorite as a whole, as they are important to model the eventual response of NEAs to impacts. Thus we also performed larger indentations, with an applied force of 500 mN. With such configuration we get rid of the indentation size effect, which is a progressive increase of hardness observed for low indentation forces and ascribed to a variety of factors (see, e.g. Gerberich et al. 2002; Nix & Gao 1998). Thus, the results obtained from indentations performed using 500 mN should be more representative of the real behavior. An array of 16 large indentations was performed on each of the lithologies and regions identified on our selected specimen (Figure 1): the light-colored and the impact-melt lithologies, and a thick black shock-melt vein consisting of a fine grained silicate matrix with abundant metal and troilite inside (Nazarov et al. 2013; Bischoff et al. 2013). In order to calculate their mean mechanical properties we averaged the results of the large indentations performed on the silicates of each region (olivine, pyroxene and plagioclase), as these minerals account for ~90% of the material in Chelyabinsk (Galimov et al. 2013). We believe that this combination of several indentations can provide a good estimation of the average mechanical properties of each region. Although higher applied loads could result in even more representative values, larger forces cannot be applied with our indenter. The hardness (H) and reduced Young's

modulus (E_r) values were determined using the conventional method (Oliver & Pharr 1992; see Appendix). The elastic recovery was evaluated as the ratio between the elastic and the total (plastic + elastic) energies during nanoindentation, W_{el}/W_{tot} (see Appendix).

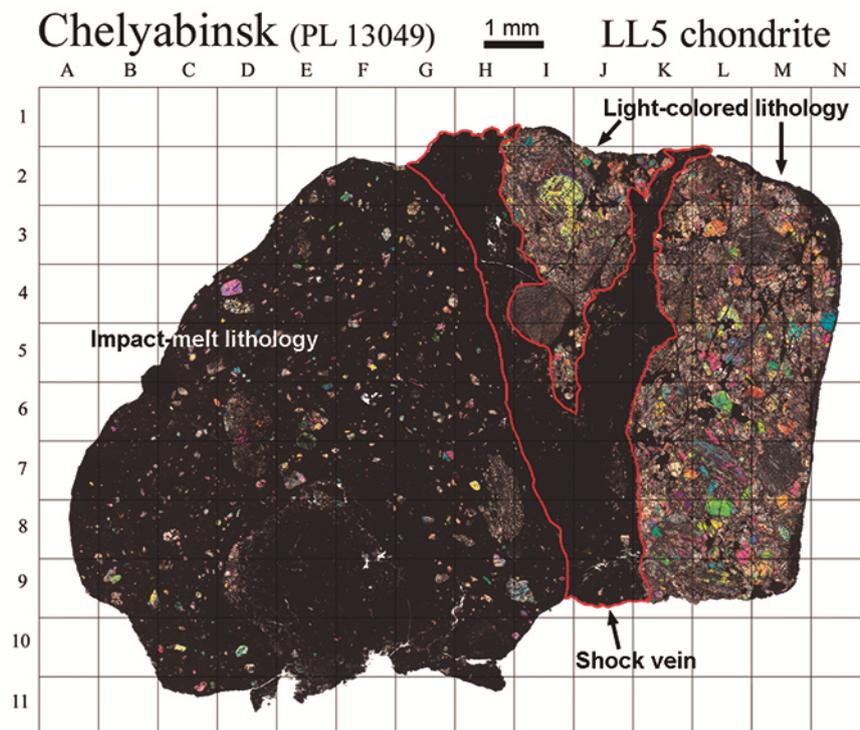


Figure 1. Transmitted light optical mosaic from the Chelyabinsk PL 13049 thin section. The three different regions studied here are highlighted. The size of each square of the grid is 1 mm^2 .

3. RESULTS

A summary of calculated mechanical properties of the main minerals obtained after local indentations (maximum load of 20 mN) is presented in Table 1, and representative

load-displacement curves acquired from each mineral phase are shown in Figure 2. According to our results, olivine shows higher hardness and reduced Young's modulus than pyroxene, while both properties notably decrease for plagioclase. It has to be taken into account that the indented olivine included inter-granular metal and troilite veins, which probably increase the variability in mechanical properties. Lower average values of both mechanical parameters are obtained for troilite. For metal grains, composed of kamacite and taenite in variable proportion, rather low hardness and relatively low reduced Young's modulus were found. Taenite has a lower reduced Young's modulus and similar (but slightly lower) hardness than kamacite. The last mineral phase indented was chromite, which shows high values of both H and E_r . Concerning the elastic recovery, the highest mean values are achieved in regions where pyroxene and plagioclase are mixed. Chromite also shows a high elastic recovery, whereas troilite and metal inclusions show much lower values.

Table 1

Average mechanical properties of Chelyabinsk minerals

Mineral phase	E_r (GPa)	H (GPa)	W_{el}/W_{tot}
Olivine	136 ± 5	13.6 ± 0.9	0.551 ± 0.023
Pyroxene	122 ± 11	11.9 ± 2.2	0.59 ± 0.03
Pyroxene + Plagioclase	71 ± 5	9.6 ± 1.0	0.720 ± 0.025
Troilite	71 ± 8	5.1 ± 0.8	0.45 ± 0.05
Taenite	82 ± 6	3.05 ± 0.29	0.232 ± 0.007
Kamacite	127 ± 16	3.58 ± 0.24	0.20 ± 0.06

Chromite

 131 ± 3 15.9 ± 1.3 0.666 ± 0.016

Notes. Reduced Young's modulus (E_r), hardness (H) and elastic recovery (W_{el}/W_{tot}) of the same mineral phases from where the curves at Figure 2 were obtained. Each was calculated averaging the results obtained from several small indentations (up to 20 mN).

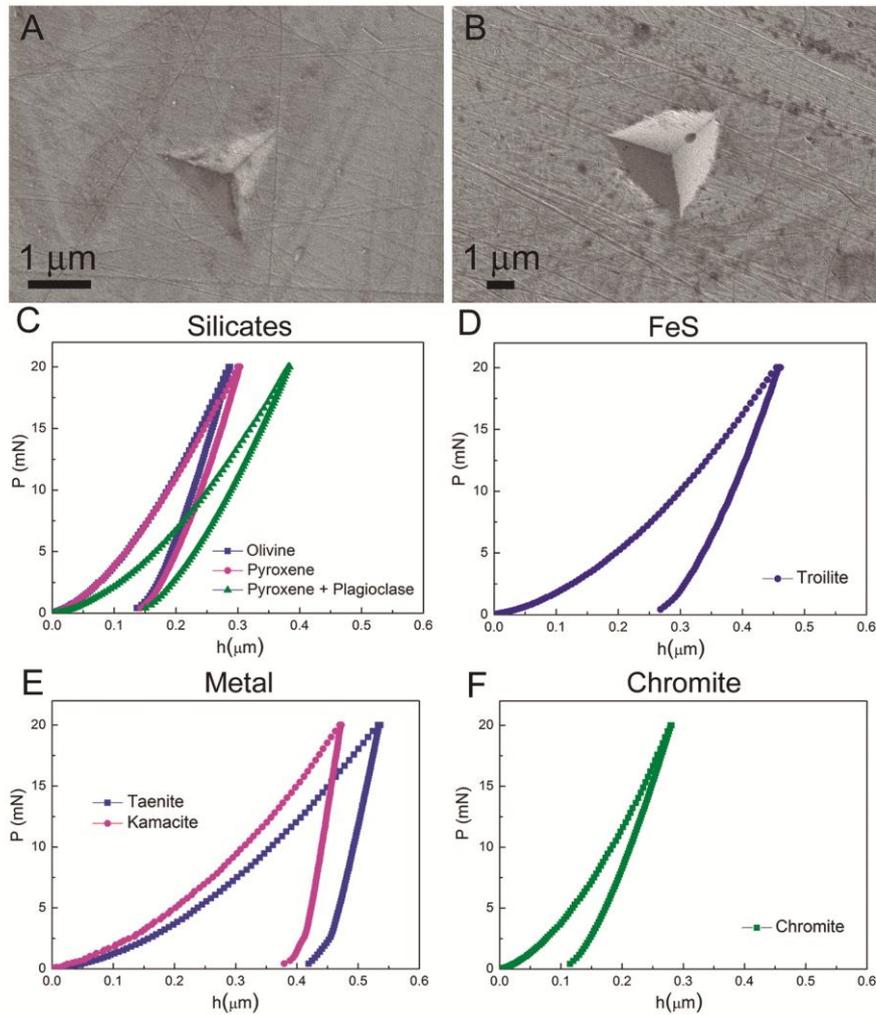


Figure 2. Low-load indentations (up to 20 mN) performed on the Chelyabinsk meteorite. At the top BSE images show indentations on: olivine (A), and metal (B, taenite or kamacite). Indentation curves for different mineral phases are shown below. Silicates (C): Olivine (F₀₇₅, plus tiny troilite veins), Pyroxene (En₈₀), and Pyroxene +

Plagioclase (pyroxene plus a small amount of Ab₉₀ plagioclase). Troilite (**D**): contains ~54% of atomic S. Metal grains (**E**): Taenite (~35% of atomic Ni) and Kamacite (~5% of atomic Ni, plus a small amount of troilite). Finally chromite curve is shown (**F**). The selected curves are considered representative of the mean mechanical properties of each mineral phase (shown in Table 1).

The results obtained after performing larger indentations (maximum load of 500 mN) to obtain combined mechanical properties are summarized in Table 2, while representative indentation curves for each lithology or region are shown in Figure 3. The impact-melt lithology shows the highest hardness, while the highest values of reduced Young's modulus correspond to the black shock-melt vein. The three regions show very similar values of elastic recovery.

Table 2

Average mechanical properties of Chelyabinsk regions

Region	E_r (GPa)	H (GPa)	W_{el}/W_{tot}
Light-colored lithology	69 ± 8	9.7 ± 2.1	0.65 ± 0.07
Impact-melt lithology	71 ± 8	12.2 ± 2.2	0.69 ± 0.05
Shock-melt vein	77 ± 8	11.8 ± 2.3	0.679 ± 0.027

Notes. Reduced Young's modulus (E_r), hardness (H) and elastic recovery (W_{el}/W_{tot}) were calculated averaging the results obtained from several large indentations (up to 500 mN).

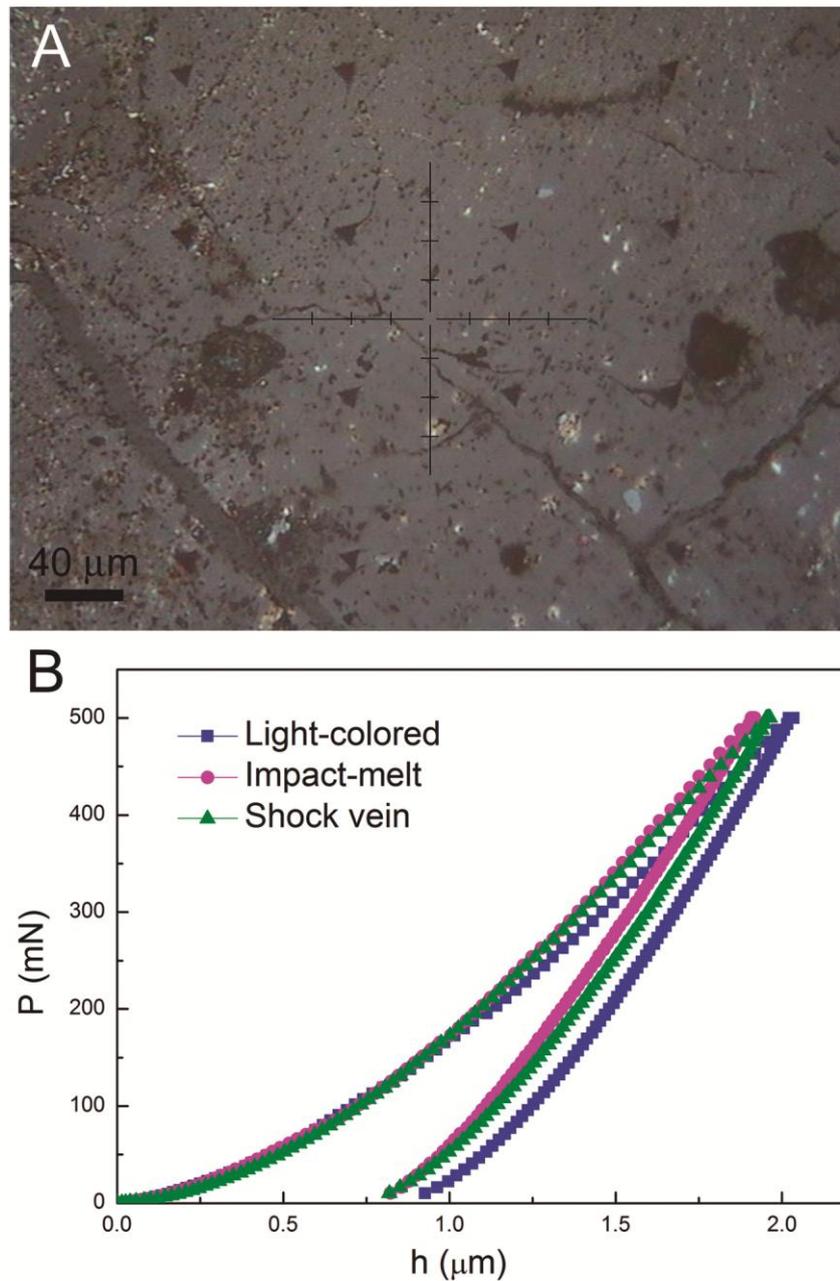


Figure 3. High-load indentations (up to 500 mN) on the Chelyabinks lithologies. At the top, an optical image (A) shows an array of 16 high-load indentations on the shock-melt vein. Below (B), three indentation curves representative of the mean mechanical properties obtained from the three regions analyzed with the high-load indentations (see data in Table 3).

Remarkably, small cracks are sometimes formed at the edges of the indentations performed using 500 mN load. These cracks are the result of localized fracture, and are almost only observed in the shock-melt vein, as can be seen in Figure 4. The formation and length of these indentation cracks can be correlated with fracture toughness of the indented regions, with longer cracks being indicative of materials more prone to fragmentation (see Appendix). The length of the most significant cracks we could find on our indentations is between 7 and 13 μm , from the center of the indentation to the end of the crack. Thus, with an average c (length from the center of the indentation to the end of the crack) of $\sim 10 \mu\text{m}$, we obtain a fracture toughness of $0.62 \pm 0.12 \text{ MPa}\cdot\text{m}^{1/2}$.

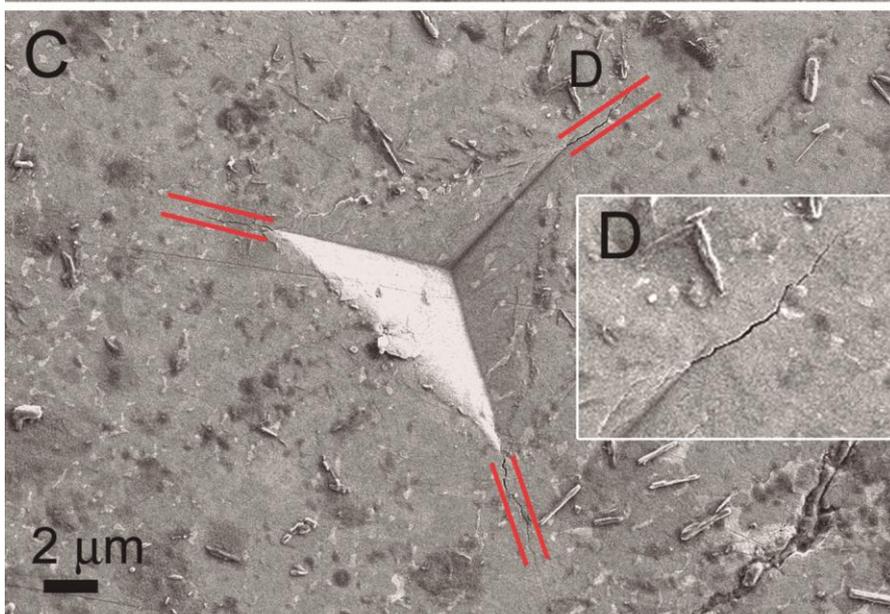
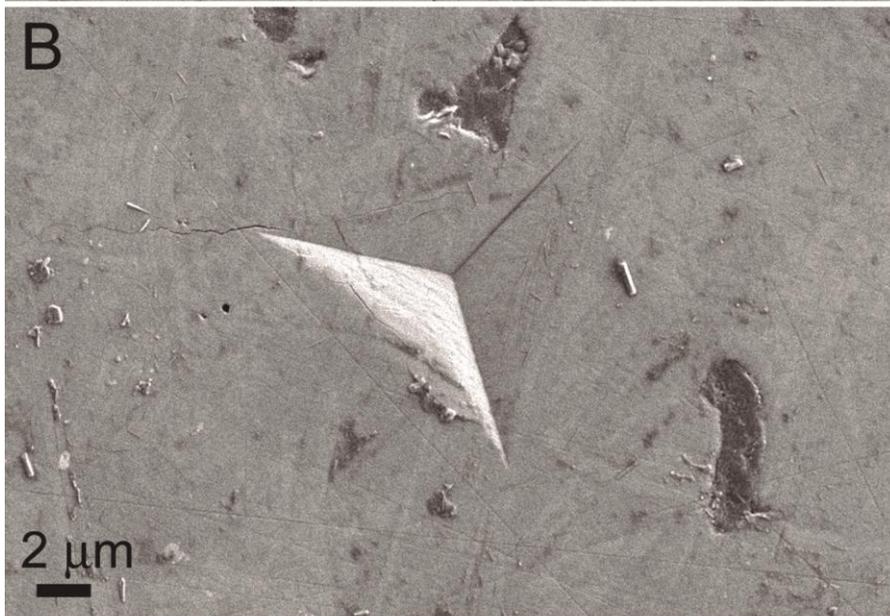


Figure 4. Fractures after high-load (up to 500 mN) indentations on the three regions analyzed on this Chelyabinsk specimen. Three BSE images show high-load indentations on: **(A)** -> the light-colored lithology; **(B)** -> the impact-melt lithology; **(C)** -> the shock-melt vein. Fractures can only be clearly identified in the shock-melt vein (surrounded by parallel red lines in **(C)**, and shown into more detail in **(D)**).

The reduced Young's modulus can be easily related with the Young's modulus (E) if one knows the Poisson's ratio (ν) of the indented material (see Appendix). Using the Poisson's ratio defined for ordinary chondrites in previous studies (see e.g. Yomogida & Matsui 1983), we see that the reduced Young's modulus values obtained here for the Chelyabinsk meteorite are consistent with the Young's modulus measured previously for ordinary chondrites, between 10 and 140 GPa (see e.g. Yomogida & Matsui 1983; Flynn 2005). As expected, our results indicate that Chelyabinsk has a much higher average reduced Young's modulus than the reported for carbonaceous chondrites, typically around 20 GPa (Britt et al. 2015). Concerning hardness, there is a dearth of previous studies where nanoindentation is applied to meteorites, hence preventing us from an extensive comparison with the literature. However, our measurements of hardness in kamacite and taenite seem to be as expected (Brusnitsyna et al. 2016), thus making our results reliable. Similarly to reduced Young's modulus, the nanoindentation hardness of carbonaceous chondrites must be orders of magnitude lower than that of Chelyabinsk (see Appendix), mainly due to the high porosity of carbonaceous chondrites, which can exceed 30% (Macke et al. 2011; Consolmagno et al. 2008; Pellicer et al. 2012). Ordinary chondrites typically show porosities around 5 to 10% (Consolmagno et al. 2008), and indeed the values reported for Chelyabinsk range between 2 and 11%, with an average value of ~6%. However, the porosity is almost

identical all around Chelyabinsk (Kohout et al. 2014), and therefore changes in porosity cannot be considered the cause for the measured differences in mechanical properties between the different investigated regions.

4. DISCUSSION

As an LL5-6 ordinary chondrite breccia, Chelyabinsk is representative of the properties exhibited by the surface of a heterogeneous asteroid with distinguishable lithologies, at least at mm-size scale as exemplified by our sample. The two lithologies and the shock-melt vein studied in this work are rather similar in chemical composition, mineralogy and porosity (Kohout et al. 2014; Galimov et al. 2013). Therefore, the differences in mechanical properties between them have to be ascribed to other factors. Also it does not seem that the hardness values obtained are strongly affected by the value of applied indentation load (Nix & Gao 1998), since the values of H at 20 mN and 500 mN are not significantly different. This could mean that these loads are already high enough to avoid the apparent hardening phenomena related to the indentation size effects.

However, our results reveal that when jumping from the in-situ properties of mineral phases (Table 1, Figure 2) to the overall behavior of lithologies or shock-melt veins (Table 2, Figure 3), there is an evident decrease in reduced Young's modulus. Porosity can explain this decrease with applied indentation load (Asmani et al. 2001). As the load is increased, the influence of porosity is exacerbated since the probability to encounter voids in the sample increases. For the same reason, hardness obtained with larger load indentations would probably be smaller than what is shown here (Palchik & Hatzor 2004). The presence of various minor phases and their interaction with the matrix, mainly the melt veins filled with both troilite and metal, possibly alter the resulting mechanical properties of the lithologies, too. In this sense, indentations

performed using 500 mN are more representative of the real behavior than those carried out using 20 mN, and are therefore better to calculate average bulk properties of the lithologies, as expected. In turn, repetitive shock in the impact-melt lithology and the shock-melt vein can induce a refinement of the mean particle size while also increasing the amount of structural defects (e.g., dislocations), hence leading to mechanical hardening with respect to the light-colored lithology, due to the Hall-Petch strengthening relationship (Gil Sevillano et al. 1980). Finally, it should be noted that the extrapolation of material properties to length scales much larger than the micrometer-sized regions sampled by the 500 mN indentations should be made with caution. It is well known that the strength of bulk materials consisting of multiple phases with dissimilar properties can be very non-linear (see e.g., Tullis et al. 1991, Durham et al. 2009). The strain rate flow law for an aggregate does not necessarily follow a simple power law, as for monomineralic aggregates. However, here, the mechanically hard phases (olivine, pyroxene) constitute the majority phases in the meteorite and, therefore, the polyphase aggregate can be considered, in a first approximation, as a load-bearing strong framework with a relatively small volume proportion of weaker phases (troilite, taenite, kamacite). In such a case (i.e., with no obvious connection between the weak inclusions), deformation of the aggregate involves strain of the stronger matrix, while not much concentration of strain can occur in the inclusions. This is possibly why the hardness of the different lithologies (with values ranging between 9.7 GPa and 12.2 GPa, Table 2) is close to that of the mechanically hard phases (Table 1), since a load-bearing framework of the stronger material is formed. Otherwise the behavior of the aggregate bulk material would be similar to that of the weak phase alone, as reported by Tullis et al. 1991.

In order to take into account these results in impact scenarios, the key parameter to be considered is the β parameter, known as the momentum multiplication factor (see Appendix). If $\beta > 1$ there are fragments ejected after the impact and the impact itself becomes more efficient, due to the “momentum multiplication effect” (Hoerth et al. 2015). Several models have tried to account for the effects of collision and impact on the momentum multiplication in brittle porous materials (Hoerth et al. 2015; Benz & Jutzi 2006), and in all these models material parameters play a key role. Momentum multiplication can be assessed considering its dependence on porosity and strength of the target (Hoerth et al. 2015), besides from the influence of impact velocities and densities (or masses) of the colliding objects, their size, and other properties (see e.g. Schultz 1993; Holsapple & Housen 2012; Jutzi & Michel 2014; among many others). In solids with low porosity, like the Chelyabinsk meteorite, momentum multiplication is more pronounced since material ejection is then more directional and no impact energy is dissipated in the form of pore compaction (Hoerth et al. 2015). Then, the role of strength can be considered using the strength-dominated cratering model (see Appendix). According to this model, a lower porosity increases the momentum multiplication, and hence an impact on an asteroid like Chelyabinsk would be more effective on deflecting its trajectory than on NEAs with higher porosity. Also, for a given porosity meteorites with lower hardness (and therefore lower strength) will lead to larger values of β and higher efficiency of the impact, since the formation of the impact crater is promoted (see Appendix). Different definitions of strength (compressive strength, yield strength, ultimate strength, among others) can be related to each other on predictable ways, and therefore can be used to estimate the β parameter (Holsapple 2009). We chose the compressive yield stress (σ_C), because it can be related with hardness through the expression $H = C\sigma_C$ (see Appendix). Following the typical

definition of the constraint factor C (with a value around 3 for metals), the values of σ_C that can be inferred from or measurements of H would be much higher than the compressive strengths found in some previous studies (Buddhue 1945; Kimberley & Ramesh 2011), which are of the order of ten to a few hundred MPa. Although σ_C and compressive strength are not the same, they should be fairly similar. The apparent discrepancy with previous works comes from the specific correlation between H and σ_C , which varies for different materials, scales, and techniques (see Appendix). Indeed, C can attain values much higher than 3 for ceramics and brittle materials (Zhang et al. 2011). Since C is difficult to be determined, we calculated the ratio between β^{-1} from the light and the impact-melt lithologies, instead of their β parameter (see Appendix). Assuming a scaling parameter (μ) between 0.4 and 0.55, we see that the β^{-1} of the light-colored lithology of Chelyabinsk is between 5 and 20% higher than for the impact-melt lithology. For Chelyabinsk-like asteroids this implies that an impact in objects enriched in light-colored lithology would be more efficient than in others where the impact-melt lithology predominates. The light-colored lithology can be easily distinguished spectroscopically from the dark-colored and the impact-melt lithologies, due to the remarkable darkening of the latter two (Reddy et al. 2014; Popova et al. 2013). This would allow an asteroid impact mission to select the most suitable region to be targeted in order to obtain a more efficient deflection.

We have also seen that although the reduced Young's modulus of the two lithologies and the shock-melt vein are rather similar, the prominent formation of cracks after indentations on the latter is indicative of the ease to create fractures within these veins. It is likely that the amount of mass ejected after an impact depends on fracture toughness, with low fracture toughness values promoting larger ejecta mass, and therefore higher momentum multiplication (Walker & Chocron 2015). Not many studies

calculate the fracture toughness of meteorites, but our result of $0.62 \pm 0.12 \text{ MPa}\cdot\text{m}^{1/2}$ is clearly lower than the $2 \text{ MPa}\cdot\text{m}^{1/2}$ estimated by some other authors (Walker & Chocron 2015). This is indicative of the ease to break these shock-melt veins, which would promote the fracturation and consequent ejection of surface materials, promoting the momentum multiplication.

In order to use these results in the frame of an impact deflection mission such as the AIDA, it is necessary to understand the connection between quasistatic indentations and dynamic indentations or impacts. Although it has been observed that for most brittle materials dynamic hardness values are larger than quasistatic values, typically by 10-25% (Anton & Subhash 2000; Wheeler 2009), important mechanical behavior properties still hold for the case of dynamic tests (see Appendix). The difference in scale between a cm-size sample and an asteroid can also have an important effect on the effective mechanical properties. Although several studies have already considered how hardness, strength, and momentum multiplication are affected by variations in size (Schultz 1993; Holsapple & Housen 2012; Jutzi & Michel 2014), a deeper understanding provided by real scale experiments is required for a proper extrapolation of properties from meteorites to asteroids.

5. CONCLUSIONS

Using the nanoindentation technique, we have studied the mechanical properties of one thin sample of the Chelyabinsk meteorite. Since this meteorite is an LL5-6 ordinary chondrite breccia, it is a good proxy for the surface materials of heterogeneous NEAs with distinguishable lithologies and some of the most abundant chondritic materials impacting the Earth. Its study provides constraints to improve our understanding of the mechanical response of such bodies. We summarize our conclusions as following:

1. The value of applied indentation load (20 mN or 500 mN) does not significantly affect Hardness, but reduced Young's modulus decreases notably when moving from the in-situ properties of mineral phases to the overall behavior of lithologies or shock-melt veins. Porosity can explain this decrease, since with larger loads the probability to encounter voids in the sample increases.
2. The differences in mechanical properties between the two lithologies and the shock-melt vein studied in this work cannot be attributed to variations in chemical composition, but the presence of minor phases and melt veins can possibly affect them. Also, lower mean particle size produced by repetitive shock implies increasing the structural defects and therefore the mechanical hardening, which increases the Hardness of the impact-melt lithology and the shock-melt.
3. Indentations produce cracks in the shock-melt vein, providing as a result a low fracture toughness value, which is indicative of the ease to create fractures within these veins. Low fracture toughness can promote the ejection of surface materials after an impact, therefore increasing momentum multiplication. The shock-melt veins are, therefore, one important structural weakness of Chelyabinsk-like asteroids.
4. Since for a given porosity lower hardness implies larger momentum multiplication, asteroids dominated by light-colored lithology would be easier to deflect than asteroids mainly composed of the impact-melt lithology. As they can be easily distinguished spectroscopically, an asteroid impact mission would be able to select the region where an impact would be more efficient.
5. Our results represent a first step in the use of nanoindentation as a technique to get additional insight on the mechanical properties of chondritic bodies, and

support AIDA and other future asteroid deflection missions to palliate unexpected impact hazard on human beings.

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APPENDIX

TECHNICAL SPECIFICATIONS

In this study we used a thin ($\sim 30 \mu\text{m}$) section (by the name of PL 13049) from one specimen of the Chelyabinsk meteorite, kindly provided by Addi Bischoff. The section was polished to mirror-like appearance using diamond paste.

Two high-resolution mosaics of the section were created from separate 50X images taken with a Zeiss Scope Axio petrographic microscope. They were composed of reflected and transmitted light images (Figure 1), respectively. Those mosaics allowed us to establish target features and regions to be analyzed with scanning electron

microscopy (SEM), Energy Dispersive X-ray spectroscopy (EDS) and nanoindentation. A 1 mm² grid was superimposed to locate and naming the different features under study through the specimen (Figure 1).

SEM allowed us to study the microstructure of the specimen, while the chemical composition of the different regions was determined with EDS. SEM images were taken on a Zeiss Merlin field emission (FE) SEM at 1.20 kV. The same instrument allowed acquiring EDS patterns at 15 kV.

CALCULATIONS USING NANOINDENTATION DATA

Nanoindentation consists in applying a controlled load into a sample through the use of a hard indenter. The indenter pushes the surface while increasing load, up to a specific maximum load. The maximum depth achieved is then measured. When the indenter is unloaded, the sample surface pushes back due to elasticity. The obtained load-displacement curves provide information about the deformation mechanisms (elastic, and plastic), and the elastic recovery, through the loading and unloading curves, respectively. The hardness (H) and reduced Young's modulus (E_r) values are determined from these curves using the method of (Oliver & Pharr 1992). From the initial unloading slope, the contact stiffness, S , are determined as:

$$S = \frac{dP}{dh} \quad (\text{A1})$$

where P and h denote the applied force and the penetration depth during nanoindentation, respectively. The reduced Young's modulus is evaluated based on its relationship with the contact area, A , and the contact stiffness:

$$S = \beta' \frac{2}{\sqrt{\pi}} E_r \sqrt{A} \quad (\text{A2})$$

Here, β' is a constant that depends on the geometry of the indenter ($\beta' = 1.034$ for a Berkovich indenter according to Fischer-Cripps 2004). E_r is defined as follows:

$$\frac{1}{E_r} = \frac{1-\nu^2}{E} + \frac{1-\nu_i^2}{E_i} \quad (\text{A3})$$

The reduced Young's modulus takes into account the elastic displacements that occur in both the specimen, with Young's modulus E and Poisson's ratio ν , and the diamond indenter, with elastic constants E_i and ν_i . Note that for diamond, $E_i = 1140$ GPa and $\nu_i = 0.07$. Hardness can be calculated using the following expression:

$$H = \frac{P_{Max}}{A} \quad (\text{A4})$$

where P_{Max} is the maximum force applied during nanoindentation. The elastic recovery is evaluated as the ratio between the elastic and the total (plastic + elastic) energies during nanoindentation, W_{el}/W_{tot} . These energies are calculated from the nanoindentation experiments as the areas between the unloading curve and the displacement axis (W_{el}) and between the loading curve and x-axis (W_{tot}) (Fischer-Cripps 2004).

The nanoindenter used here was a UMIS equipment from Fischer-Cripps Laboratories, operated in the load control mode and using a Berkovich pyramidal-shaped diamond tip. The maximum load applied was of 20 mN for low-load local indentations, and up to 500 mN for larger indentations (the maximum available force for our indenter). All indentations reached depths between 0.2 and 2 μm (see Figures 2 and 3), but the compressive stresses caused by the indenter are not limited to the size of the indent. Therefore, the depth affected and measured by the nanoindentations can go down to ~ 5 μm for 20 mN loads, and to 20 μm for the 500 mN loads, not deep enough to be disturbed by the properties of the glass supporting the sample (Fischer-Cripps 2004).

From the load displacement curves we obtained the hardness (H), reduced Young's modulus (E_r) and elastic recovery (W_{el}/W_{tot}) of the sample with the method described above (Oliver & Pharr 1992). The thermal drift during nanoindentation was kept below 0.05 nm s^{-1} . Proper corrections for the contact area (calibrated with a fused quartz specimen), initial penetration depth and instrument compliance were applied.

Whenever cracks are formed after indentations, the formation and length of these cracks can be correlated with fracture toughness of the indented regions. For a Berkovich indentation impression, the fracture toughness, K , can be given as (Fischer-Cripps 2004):

$$K = k \left(\frac{E}{H} \right)^n \frac{P}{c^{3/2}} \quad (\text{A5})$$

where k is an empirical constant close to 0.016, P is the applied indentation force, $n = 1/2$ and c is the length from the center of the indentation to the end of the crack. Longer cracks thus result in lower fracture toughness values.

THE MOMENTUM MULTIPLICATION FACTOR (β)

The momentum multiplication factor (β), necessary to connect the obtained mechanical properties with impact scenarios, is defined as the momentum change divided by the momentum input:

$$\beta = \frac{\Delta p_t}{M_p v_p} = \frac{M_p v_p + M_e v_e}{M_p v_p} = 1 + \frac{M_e v_e}{M_p v_p} \quad (\text{A6})$$

Here Δp_t is the momentum change of the target due to the impact. M_p and v_p are the total mass and average velocity of the projectile, whereas M_e and v_e are the total mass and

average velocity (on anti-impact direction) of the ejected material. If $\beta > 1$ there are fragments ejected after the impact, and the efficiency of an impact would be larger, since the attained Δv would be also more significant. This effect is called “momentum multiplication” and has been studied by several authors (Hoerth et al. 2015).

The role of strength on impacts can be considered using the strength-dominated cratering model, which provides the following scaling relation (Holsapple & Housen 2012):

$$\beta - 1 \approx \left(v_p \sqrt{\frac{\rho_t}{Y}} \right)^{3\mu-1} \left(\frac{\rho_t}{\rho_p} \right)^{1-3\nu} \quad (\text{A7})$$

Where v_p is the projectile velocity, ρ_t and ρ_p are the densities of the target and the projectile, Y is some measure of the strength of the target, μ is a scaling parameter close to 0.55 for non-porous materials and between 1/3 and 0.4 for highly porous materials, and ν is a constant close to 0.4 for most target materials (Holsapple & Housen 2007). Many definitions of strength can be used for a geological material, but they all can be related on predictable ways (Holsapple 2009). We choose the compressive yield stress (σ_C) because it can be related to hardness following the typical expression $H = C\sigma_C$ in constrained materials (free from porosity), and because this relationship between hardness and compressive yield stress holds for the case of dynamic tests (Subhash et al. 1999). C is the constraint factor and attains values close to 1.6 for rocks and 3 for metals, although values as high as 180 have been found for ceramics and other brittle materials (Zhang et al. 2011). Due to the size of nanoindentations, porosity, interactions between different phases, and cracks, are not perceived on these measurements. Altogether, those factors can imply a much higher H , and therefore σ_C , than expected for those materials at larger scales (Pellicer et al. 2012; Palchik & Hatzor 2004).

Considering those points, we lack the means to calculate from our measurements of H a value of σ_C representative of a real asteroid, and therefore we cannot find β . However, we know that H and σ_C can be related through a specific constrain factor C . Therefore, for two similar materials A and B (and assuming the same C , density, and μ), we can calculate their ratio of β -1:

$$\frac{(\beta-1)_A}{(\beta-1)_B} \approx \left(\frac{Y_B}{Y_A}\right)^{\frac{3\mu-1}{2}} = \left(\frac{\sigma_{CB}}{\sigma_{CA}}\right)^{\frac{3\mu-1}{2}} = \left(\frac{H_B}{H_A}\right)^{\frac{3\mu-1}{2}} \quad (\text{A8})$$

That ratio provides us an idea of which of those materials would grant a higher momentum multiplication and a more efficient impact.

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