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# Segregation scheme of Indium in AlGaInAs nanowire shells

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Quaternary alloys enable the independent optimization of different semiconductor properties, such as the separate tuning of the bandgap and the lattice constant. Nanowire core-shell structures should allow a larger range of compositional tuning as strain can be accommodated in a more effective manner than in thin films. Still, the faceted structure of the nanowire may lead to local segregation effects. Here, we explore the incorporation of In in AlGaAs shells up to 25 % . In particular, we show the effect of In incorporation on the energy shift of the AlGaInAs single-photon emitters present in the shell. We observe a redshift up to 300 meV as a function of the group-III site fraction of In. We correlate the shift with segregation at the nanoscale. We find evidence of the segregation of the group-III elements at different positions in the nanowire, not observed before. We propose a model that takes into account the strain distribution in the nanowire shell and the adatom diffusion on the nanowire facets to explain the observations. This work provides novel insights on the segregation phenomena necessary to engineer the composition of multinary alloys.

**Introduction** Quaternary alloys enable to independently tune the semiconductor lattice constant and bandgap by careful composition engineering. However, the controlled deposition of quaternary alloys with randomly distributed composition can be challenging. On one hand, miscibility gaps and diffusion may segregate the different species that compose the alloy, and limit the possible compositions [1, 2]. On the other hand, strain may build up when quaternary semiconductors epitaxially grow on substrates of different lattice constant. To avoid plastic relaxation [3, 4], the theoretical range of available compositions is significantly narrowed in real applications. Due to their reduced diameter, semiconductor nanowires (NWs) provide a suitable growth platform to minimize plastic relaxation, [5–10] permitting lattice-mismatched material combinations not achievable by planar schemes [11–14].

In this work, we explore the range in which In can be incorporated in AlGaAs shells of core-shell GaAs-AlGaAs NWs. The goal is to understand how much the emission energy of single-photon emitters, spontaneously forming in the shell, can be redshifted [15–17]. In particular, we want to dissociate the bandgap engineering from additional segregation effects that may occur due to the presence of In. Consequently, we characterize the optical emission and correlate it with the incorporation of In using several techniques at different length scales. We find that group-III elements in the shell segregate to different positions and generate regions of different bandgaps. In

particular, we find consistent evidence of the In segregation in novel wedge-shaped In-rich features. We correlate these observations with the 3-fold polarity of the NW cross section and the migration of the adatoms on the NW sidewalls. In addition to the role of the crystal directions of the NW facets [18–20], we consider strain to cause the peculiar In segregation; the presence and distribution of strain are further analyzed by finite-element simulations as well as Raman and photoluminescence (PL) spectroscopy.

**The material system: core-shell NWs** The GaAs-AlGa(In)As core-shell NWs are grown in a high-mobility molecular-beam-epitaxy system (MBE, DCA P600). We use Si(111) substrates and self-catalyzed growth [23, 24] at a substrate temperature of 640°C. When the NWs are about 10 $\mu$ m long, we stop the axial growth by interrupting the Ga supply and consuming the catalyst. We then lower the substrate temperature to 460°C to promote radial growth of an AlGa(In)As shell on the NW sidewalls [25, 26]. We start with an Al<sub>33</sub>Ga<sub>67</sub>As shell and then incorporate In without modifying the Ga or Al rates. For this, we varied the In partial pressure from 4.5 X 10<sup>-9</sup> Torr to 2.5 X 10<sup>-7</sup> Torr, while the Al and Ga pressures were kept constant. Unless stated, the samples are identified by the group-III site fraction of In measured through various techniques and expressed as percentage in the notation Al<sub>x</sub>Ga<sub>y</sub>In<sub>1-x-y</sub>As (Al, Ga, and In sum to 100%) : 0%, 1%, 2%, 3%, 4%, 15%, and 25%. An outer GaAs shell of 5 nm prevents the oxidation of the inner AlGa(In)As.

Previous reports show that the composition of an AlGaAs shell deposited around a GaAs NW exhibits

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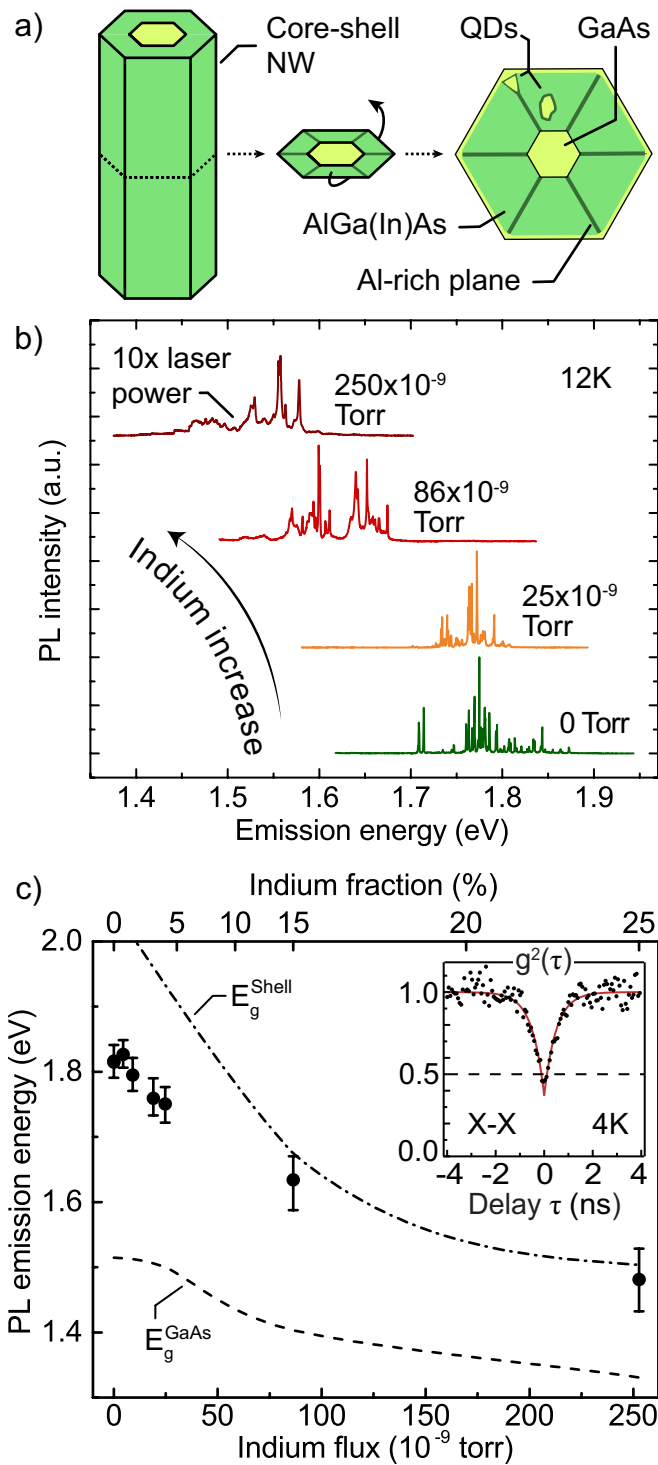


FIG. 1: (a) Sketch of a core-shell NW and shell QDs. (b) Micro-PL spectra of shell QDs grown at increasing In pressures. (c) Ensemble medians of the PL QD emission energy vs the In MBE pressure (bottom abscissa) or shell fraction (top abscissa) given by APT, XRF, STEM EDS and growth-rate calibrations.  $1^{st}$  and  $3^{rd}$  quartile as error bars. The dash and the dash-dot lines are the strained core and shell bandgaps, respectively [21, 22]. Inset: single-photon  $g^2$  of a QD exciton (3% In).

74 nanoscale fluctuations [15, 16, 27–30]. This phenomenon  
 75 depends on the different surface mobility of Ga and Al  
 76 during growth and the presence of surface-potential gra-  
 77 dients ( $\mu$ ) on the NW sidewalls. The formation of Al-rich  
 78 planes at the ridge between two NW facets was reported  
 79 [19]. Furthermore, Ga-rich clusters can form quantum-  
 80 dots (QDs) that behave as single-photon emitters in the  
 81 NW shell [15, 16, 29]. A sketch of a core-shell GaAs-  
 82 AlGaAs NW and its cross section is presented in fig-  
 83 ure 1a, illustrating the distribution of Al-rich planes and  
 84 QDs made from Ga-rich nanoclusters.

85 **Light-emission properties** First we address the opti-  
 86 cal functionality of the core-shell NWs and embedded  
 87 QDs. We used a single-frequency optically pumped semi-  
 88 conductor laser with a wavelength of 532 nm and power  
 89 of 100 W/cm<sup>2</sup> focused in a spot of less than 1  $\mu$ m in di-  
 90 ameter to measure NW samples at 12 K using a helium  
 91 cryostat. The PL signal is collected into a spectrome-  
 92 ter and dispersed by a 300 l/mm grating onto a Peltier-  
 93 cooled CCD. The QDs in the AlGaAs shell emit bright  
 94 and narrow PL lines (linewidth below 100  $\mu$ eV) [15] be-  
 95 tween 1.7 eV and 1.9 eV [27]. The green spectrum at the  
 96 bottom of figure 1b illustrates the PL emission of these  
 97 structures; the sharp peaks are attributed to the pres-  
 98 ence of QD single-photon emitters [15, 31]. Figure 1b  
 99 also contains PL spectra of the shell-QD emission for  
 100 increasing In fraction, revealing several sharp peaks at  
 101 different energies. The emission-energy range redshifts  
 102 with increasing In fraction in the shell. One can also  
 103 qualitatively determine that the QD emission linewidth  
 104 broadens in samples of higher In fraction. The sample  
 105 with the highest In fraction showed a decreased PL in-  
 106 tensity, compensated for by increasing the laser power  
 107 (see figure 1b).

108 For a statistical analysis on large ensembles of NWs  
 109 and QDs, figure 1c shows the median emission energy  
 110 of the QDs as a function of the measured In pressure  
 111 in the MBE chamber (bottom axis) or as a function of  
 112 the In fraction (top axis) measured in the shell by X-ray  
 113 fluorescence (XRF) combined with atom probe tomogra-  
 114 phy (APT) for 1% to 4% of In and scanning transmis-  
 115 sion electron microscopy X-ray energy-dispersive spec-  
 116 troscopy (STEM EDS) for 15% and 25% In (supplemen-  
 117 tary information). We measured 25 NWs from each sam-  
 118 ple. The QD emission lines were identified by an auto-  
 119 matic routine [27]. The dash-dot line indicates the ex-  
 120 pected shell bandgap as a function of shell composition  
 121 [21]. The dashed line corresponds to the GaAs bandgap.  
 122 Both lines are corrected for the simulated strain [22] that  
 123 arises from the core-shell lattice mismatch, as further dis-  
 124 cussed in the manuscript.

125 The inset in figure 1c shows a Hanbury-Brown-Twiss  
 126 autocorrelation measurement ( $g^2(\tau)$ ) of the exciton line  
 127 of a QD from a sample with 3% In. The power-dependent  
 128 PL showing the exciton nature of the emission line is  
 129 reported in the supporting information. The sample is  
 130 measured at 4.2 K in an Attodry 700 closed-cycle cryostat  
 131 and is excited in continuous wave by a HeNe laser at

132 632.8 nm through an objective with NA = 0.81, which  
 133 also collects the QD signal from the cryostat. A 1200  
 134 1/mm grating is used to select the QD line of interest  
 135 that is sent to a 50:50 beam splitter. The two paths  
 136 of the beam splitter are coupled to single-mode optical  
 137 fibers that send the signal to two single-photon avalanche  
 138 diodes, one for each path. The dip in the autocorrelation  
 139 function in the inset of figure 1c is below 0.5, which is the  
 140 signature of the single-photon emission. Several factors  
 141 increase the count at zero delay, including background  
 142 counts from QD lines spectrally close to the chosen one.

143 The effect of the shell-composition engineering on the  
 144 QD emission energy is significant. From figure 1c, the  
 145 median QD emission redshift is visible as a function of  
 146 the increasing In incorporation, up to about 300 meV.  
 147 In a random-distribution alloy, the addition of In to Al-  
 148 GaAs lowers the shell bandgap and the QD emission en-  
 149 ergy. Since these QDs form as lower-bandgap nanoscale  
 150 clusters in the AlGa(In)As matrix, their emission is ex-  
 151 pected to be redshifted with respect to the shell-matrix  
 152 bandgap (dash-dotted line in figure 1c). While this is true  
 153 in the samples with a low In fraction (up to 4%), the me-  
 154 dian QD-emission shifts less significantly from the shell  
 155 bandgap in the samples with 15% and 25% of In. Con-  
 156 comitantly, the overall QD emission-energy range broad-  
 157 ens in the shells with high In fractions (figure S1 in the  
 158 supplementary information and error bars in figure 1c).  
 159 Together with the linewidth broadening and brightness  
 160 quenching previously commented (figure 1b), these obser-  
 161 vations suggest an increasing alloy disorder as a function  
 162 of the In incorporation.

168 **Compositional analysis at the nanoscale** So far,  
 165 we have reported on the average In incorporation (fig-  
 166 ure 1c). In NWs the composition can vary significantly  
 167 at the nanometer scale due to their faceted nature, as  
 168 different facets exhibit different sticking coefficients [30].  
 169 Here we provide measurements with spatial resolution  
 170 down to the atomic scale: APT and STEM-based elec-  
 171 tron energy loss spectroscopy (EELS) and EDS.

172 We performed laser-assisted APT measurements on  
 173 the sample with 2% In. The specimen is cooled to 80 K  
 174 and irradiated with UV laser light (343 nm wavelength)  
 175 in 2-nJ pulses; the detection rate is 0.0025 events/pulse.  
 176 The evaporated NW volume is a cylinder with diameter  
 177 of 64 nm and length of 90 nm; in figures 2a - c, it is shown  
 178 as a 2D projection on a plane perpendicular to the NW  
 179 growth axis. Figures 2a - c show respectively the In, Al  
 180 and Ga fractions. It is possible to distinguish the GaAs  
 181 core by the absence of In and Al. In the shell, radial seg-  
 182 regation of Al along the ridges between two facets of the  
 183 hexagonal NW core is visible as the three Al-rich stripes  
 184 in the reconstructed NW volume, in agreement with pre-  
 185 vious works [15, 16, 18, 19, 29, 30]. In addition to the  
 186 Al-rich planes, APT reveals that the shell distribution  
 187 of the other group-III atoms is not perfectly random. A  
 188 slightly higher In fraction (up to 3%) is visible in prox-  
 189 imity of two of the Al rich planes indicated as 1 and 3  
 190 in figure 2a. Farther from these positions, the average In

191 fraction decreases to about 1.6%. In figure 2b and c we  
 192 can also observe that the shell in proximity of the Al-rich  
 193 planes is slightly depleted in Al and enriched in Ga.

194 To study the segregation of Al, Ga, and In in more  
 195 extended regions, we studied the NW cross-sections by  
 196 high angular annular dark field (HAADF) STEM and  
 197 EELS in an aberration-corrected (AC) TEM microscope  
 198 (FEI Titan) operated at 300 keV. We used the sample  
 199 with 15% In. We prepared NW cross sections to directly  
 200 map the shell composition: the NWs were spread on a  
 201 Si substrate and the cross sections were FIB-cut per-  
 202 pendicularly to the NW growth direction ( $\langle 111 \rangle$ ). The  
 203 cross-section HAADF micrographs are acquired along the  
 204  $\langle 111 \rangle$  zone axis, with the NW sidewalls corresponding  
 205 to the  $\{110\}$  family (figure S4 in the supporting informa-  
 206 tion). Figure 2d reports the HAADF micrograph of one  
 207 NW cross section. The HAADF contrast depends on the  
 208 sample thickness and composition: the higher the atomic  
 209 number (Z) of the species in the sample, the brighter  
 210 the HAADF signal. Through high-precision FIB cut, the  
 211 NW cross sections have negligible thickness variations:  
 212 the contrast in figure 2d depends on the average Z at  
 213 different positions across the sample. In this figure, the  
 214 hexagonal shape of the NW cross section is clearly visible  
 215 and the core is distinguishable from the shell. The dark  
 216 stripes that form in the shell along the ridges between two  
 217 NW facets ( $\{112\}$  crystalline directions) are due to local  
 218 enrichment with a light element, such as Al, as already  
 219 observed by APT and in agreement with the literature  
 220 [15, 18–20]. It is possible to observe that the thickness of  
 221 the Al-rich planes follows a 3-fold symmetry as observed  
 222 by Zheng et al. [18]. We obtain a thickness of 3 nm and  
 223 1.7 nm for the thick and thin planes, respectively (more  
 224 details in figure S4 in the supporting information). Sim-  
 225 ilarly, the HAADF contrast shows that the thick planes  
 226 are richer in Al than the thin ones. Comparing our obser-  
 227 vations with the literature [18], we assign the thin planes  
 228 to the A polarity and the thick ones to the B polarity.  
 229 Accordingly, we labelled the two polarities in figure 2d as  
 230 A (orange) and B (blue).

231 With the exception of the Al-rich planes, the HAADF  
 232 micrograph in figure 2d shows little compositional con-  
 233 trast in the shell. By carefully inspecting figure 2d  
 234 and figure S4 in the supporting information, a slightly  
 235 brighter contrast is visible in proximity of the Al-rich  
 236 planes: it corresponds to a local increase in the fraction of  
 237 elements with higher Z, such as Ga and In. To clarify this  
 238 observation, in figures 2e - h we report the EELS maps of  
 239 the upper half of the NW cross section shown in figure 2d:  
 240 the selected region, indicated by the dashed rectangle in  
 241 figure 2d, includes an A-polar and two B-polar ridges as  
 242 well as part of the  $\{110\}$  NW sidewalls. The Ga, In, Al,  
 243 and As distributions are shown in figure 2e, f, g, and h  
 244 respectively. Principal component analysis (PCA) was  
 245 used to enhance the signal-to-noise ratio in these maps  
 246 [32, 33]. The NW core is clearly distinguished from the  
 247 shell by the absence of Al and In and a thin Ga-rich  
 248 layer, corresponding to the GaAs capping, surrounds the

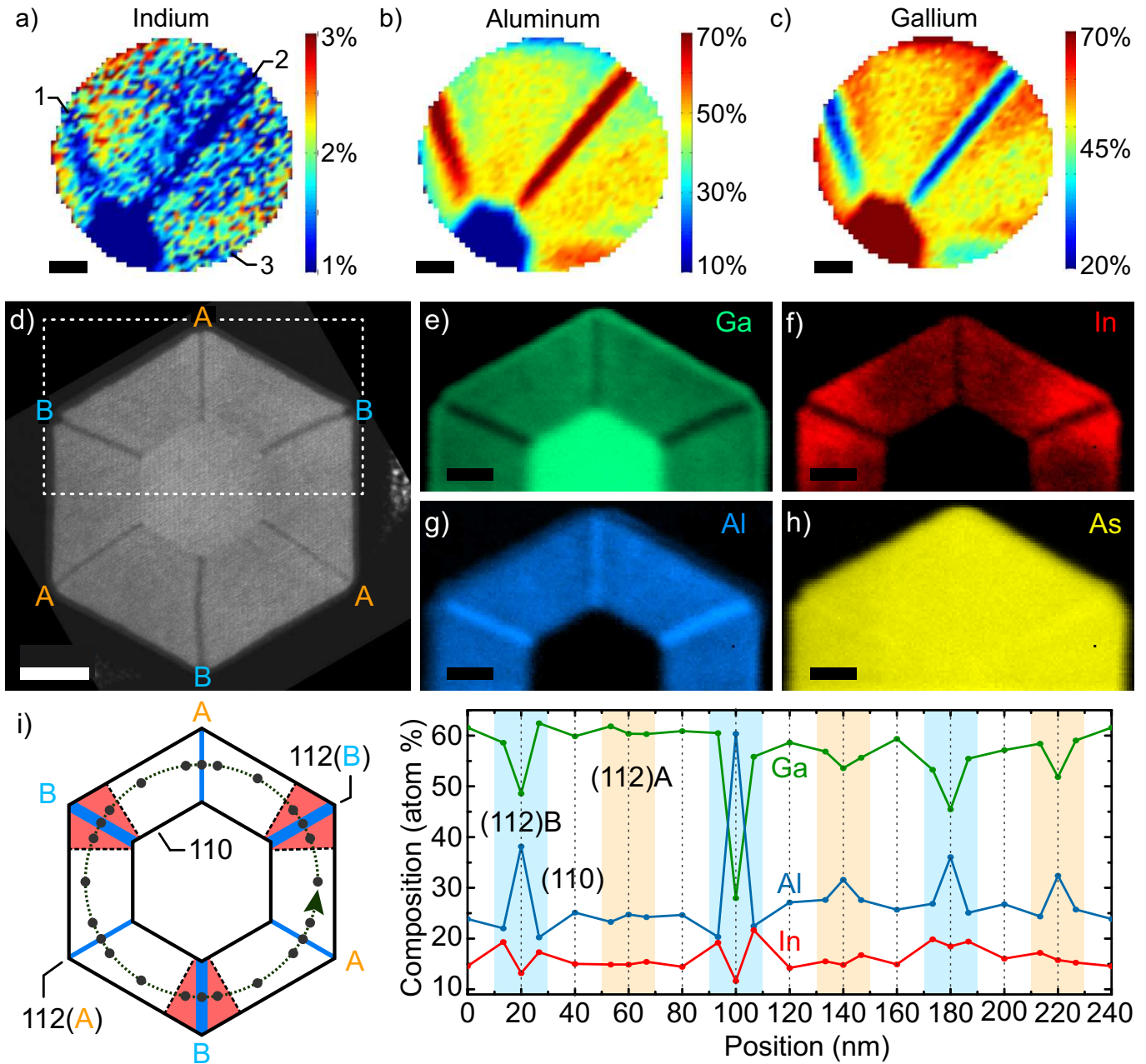


FIG. 2: (a), (b), and (c) APT of In, Al, and Ga fractions (2% In). (d) HAADF STEM micrograph of a NW cross section (15% In) with A and B polarity indicated. (e-h) EELS Ga, In, Al, and As maps of the dashed rectangle in (d). (i) Left: sketch of shell segregation. Right: shell Ga (green), Al (blue), and In (red) atomic group-III percentages vs position along the circumference in the sketch. Scale bars: (a-c) and (e-h) 10 nm, (d) 20 nm.

249 shell (figure 2e). As is randomly distributed through- 250  
 250 out (figure 2f), as opposed to the distributions of the 260  
 251 group-III elements: the presence of Al-rich planes along 261  
 252 the hexagon ridges is confirmed (figure 2g), together with 262  
 253 the Ga and In depletion at the same positions (figure 2e 263  
 254 and f). Reports on GaAs-InGaAs and GaAs-InAlAs core- 264  
 255 shell NWs [34] also show Ga and In depletion along the 265  
 256  $\langle 112 \rangle$  direction. The brightness and thickness of the 266  
 257 Al-rich planes agree with the 3-fold polarity-driven seg- 267  
 258regation previously discussed. 268

In agreement with the localized In enrichment shown 259  
 260 by APT (figure 2a), STEM EELS shows in figure 2f 261  
 262 In-rich features surrounding the Al-rich planes at the 262  
 263  $\langle 112 \rangle$ B directions: interestingly, moving from the core 263  
 264 to the outer shell, the In-rich regions become broader 264  
 265 with a wedge-like shape on each side of the Al-rich 265  
 266  $\{112\}$ B planes. Ga has a more random distribution in 266  
 267 the shell, although a slight increase is visible around the 267  
 268 A-polar  $\{112\}$  nanofacet in figure 2e. This Ga-rich fea- 268  
 268 ture is not as pronounced as the In-rich ones surrounding

the  $\{112\}$ B planes; it also has a relatively constant thickness around the  $\{112\}$ A planes, while the In-rich features develop an unusual wedge-shaped profile. One may also observe that, in proximity of the  $\{112\}$  Al-rich planes of both polarities, the Al fraction slightly decreases (figure 2g). Although the contrast is not sharp, this would agree with the local increase in the Ga and In fractions at the same positions. EELS maps on whole NW cross sections confirm the 3-fold symmetry of the shell segregation and the In enrichment only around the three  $\langle 112 \rangle$ B directions (see the supporting information). On the left of figure 2i, a scheme summarizes the main compositional segregations observed in the NW cross sections (the scheme is aligned with figures 2d - h): the red-shaded areas indicate the wedge-shaped In-rich segregation and blue stripes indicate the alternatively thicker and thinner Al-rich planes. The polarity of the  $\langle 112 \rangle$  nanofacets is labelled.

We acquired STEM-EDS maps of NW cross sections from the same samples as the one used for EELS. We use a FEI Tecnai Osiris TEM operated in STEM mode at 200 kV with a probe current exceeding 1 nA. The X-ray signal is collected by four silicon drift detectors under a solid angle of 0.9 sr. The NW cross sections are prepared by embedding the as-grown sample into epoxy. After hardening, the epoxy with embedded NWs is detached from the growth substrate and is mounted into an ultramicrotome and cut into 80-nm-thick slices transferred on a TEM grid. The EDS maps (figure S3 in the supplementary information) for the Ga, Al, and In distributions in the core-shell cross sections confirm the presence of the same features observed in the EELS maps. Our observations are particularly robust: three independent techniques (APT, EELS and X-ray EDS) confirm the formation of novel 3-fold wedge-shaped In-rich segregations in NWs with different average In fractions; together with the clear symmetry of these features, the agreement of the three techniques excludes artifacts due to sample preparation.

We now turn to a quantitative analysis of the distribution of the different species in the NW cross section: the EDS-based quantification is more accurate than the one based on EELS [35]. Figure 2i shows on the right a plot of the Ga (green curve), Al (blue curve), and In (red curve) atomic percentage extracted from a STEM EDS map as a function of the position along the dashed circumference shown in the sketch on the left (i.e. a circular linescan of 240nm in length). The data points at 0 nm and 240 nm correspond to the same position on the circumference (arrow in the sketch). The shaded areas are colored in orange and blue with the same coding as in figures 2g to distinguish the A- and B-polarity of the  $\{112\}$  directions respectively. The data points acquired in the middle of the flat  $\{110\}$  facets have a white background.

We focus on the most prominent fluctuation around the  $\langle 112 \rangle$ B direction at position 100 nm. Moving from the  $\{110\}$  plane to the  $\{112\}$ B nanofacet, the Al fraction first decreases (from 25% to 20%), then rises to 60% at

the corner. In smaller proportions, the In distribution is opposed to the Al trend: for the same positions, In first increases (from 14% to 19%) and then decreases to 12% at the corner. Ga is almost constant at 61% and 60% and then decreases to 28% at the corner. Very similar trends in Al/Ga/In fractions are reproduced at the six corners of the hexagonal NW cross section and are modulated according to a 3-fold symmetry: the compositional variations are consistently more pronounced in proximity of the  $\langle 112 \rangle$ B directions.

**Strain analysis and growth model** We turn now to the understanding of the element distribution in the shell by taking strain into account. Figure 3a shows the PL spectra of the GaAs core at 12K for core-shell NWs of increasing In fraction. The band-edge GaAs PL redshifts from 1.51 eV down to 1.37 eV and 1.29 eV for shells containing 15% and 25% In, which we attribute to the tensile strain imposed by the shell. Tensile strain in the core is also confirmed by Raman spectroscopy (supplementary information).

To gain insight in the strain distribution imposed by the lattice-mismatched core-shell NW, we simulated the strain given the lattice mismatch, dimensions and geometry of the structure. We use the methodology of Boxberg et al. [36], adapted to hexagonal GaAs-AlGaInAs NWs at low temperature [21] in the software Nextnano [37, 38]. The strain magnitude and distributions in the core and shell are calculated by minimizing the elastic energy due to the lattice mismatch between the two. Figure 3b maps of the hydrostatic strain  $\epsilon_{hydro}$  in NWs with 25% In in the shell.  $\epsilon_{hydro}$  is dominated by the principal component along the NW longitudinal axis, as reported elsewhere [39]. Shear strain components are one order of magnitude smaller than  $\epsilon_{hydro}$ , in agreement with previous reports [39]. As expected from the core-shell lattice mismatch,  $\epsilon_{hydro}$  is on average negative in the shell (compressive) and positive in the core (tensile). We find about 2.2% of tensile strain (position "A" in figure 3b), softly modulated following a 6-fold symmetry. The shell is subject to a maximum compression of -0.35% in the middle of the NW facet (position "C" of figure 3b). The shell strain intensity is modulated by a 6-fold symmetry in a more pronounced way. The outer corners are relaxed (minimum strain of 0% at position "B" in figure 3b). Importantly, the regions of smaller strain in the shell expand around the NW sidewall ridges, becoming larger the farther they are from the core. The experimental (PL) [40] and simulated strain values are plotted in figure 3c: the simulated compressive strain in the shell (position C) and the experimental and simulated tensile strain in the core increase linearly with the In fraction. The theoretical and experimental values agree within 15%.

A growth model that explains the non-random distribution of In and Al in the shell is presented. The expected behavior of Ga and In adatoms at the NW facets is sketched in figure 3d. We start by addressing the Al distribution. Al tends to incorporate more efficiently at the vertexes of the hexagonal cross section

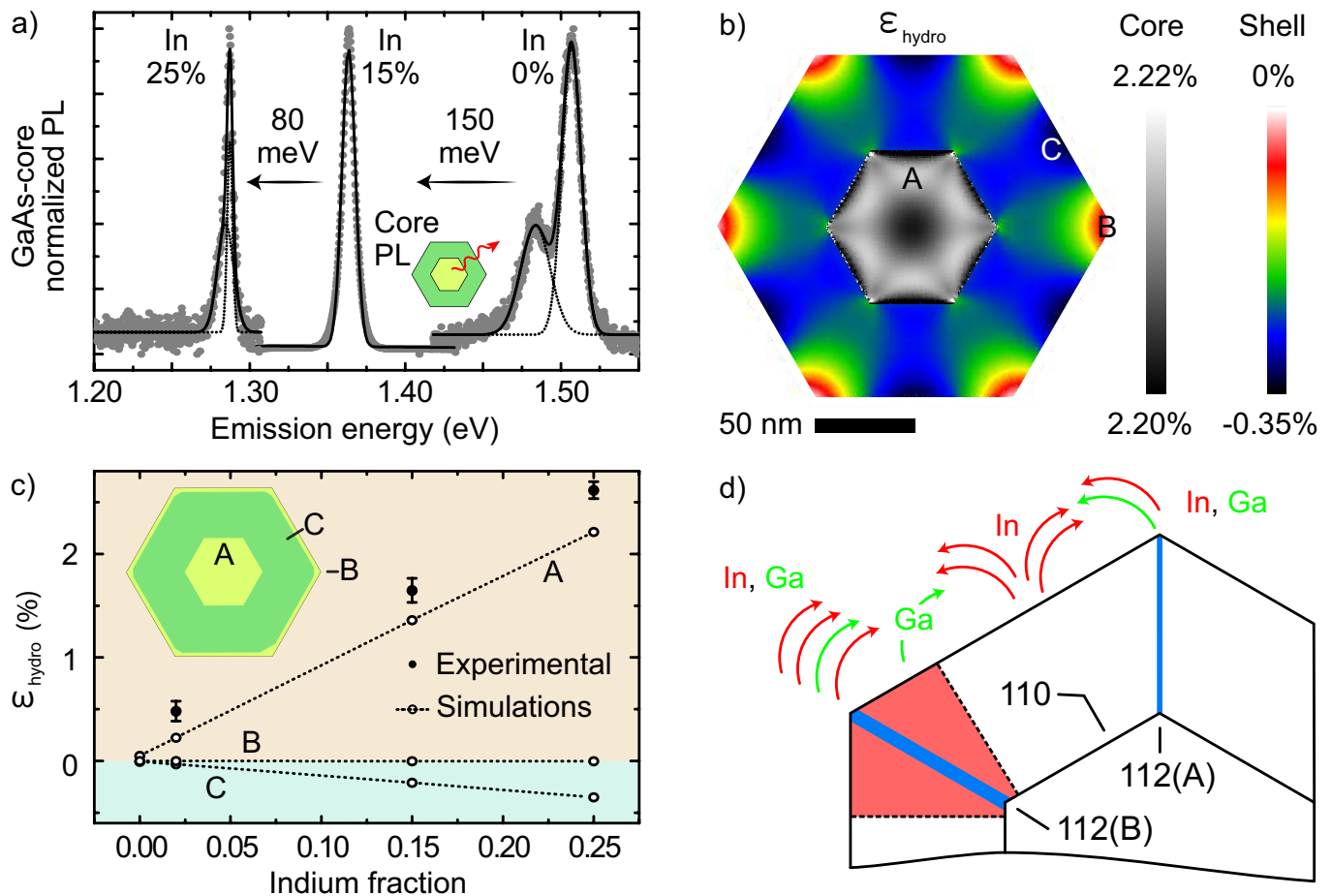


FIG. 3: (a) GaAs PL of NWs with In shell fraction of 0%, 15%, and 25%. (b) Hydrostatic strain simulation in core-shell NWs (25% In) as in ref. [36]. (c) Simulated and average experimental ( $1\sigma$  error bar) hydrostatic strain vs In fraction. (d) Sketch of the Ga and In diffusion on the NW sidewalls.

386 ( $\{112\}$  nanofacets) [15, 18, 19]. This is the consequence  
 387 of a larger sticking coefficient on those facets [41, 42]. For  
 388 a similar reason, it has been shown that the incorpora-  
 389 tion is higher on B- than A-polar facets [18]. This results  
 390 in a three-fold symmetry of the Al distribution.

391 We now turn to the incorporation of In. According  
 392 to our simulations, there is a compressive strain in the  
 393 middle of the facets (position C in the drawings). Since  
 394 AlGaInAs exhibits a larger lattice constant than AlGaAs,  
 395 we expect In to incorporate favorably in the relaxed cor-  
 396 ners. In adatoms exhibit high mobilities, allowing dif-  
 397 fusion to occur at the scale of the NW facets. Strain  
 398 relaxation should support an In flux towards the more  
 399 relaxed  $\{112\}$  corner nanofacets, while the  $\{110\}$  facets  
 400 exhibit a significantly lower In fraction in the centre (ar-  
 401 rows leaving this position in figure 3d). We also observe  
 402 that the In segregation seems to be enhanced by polarity.  
 403 Regions with a markedly higher In fraction coincide with  
 404 the  $\{112\}$ B vertexes. We note that the polarity-driven in-  
 405 corporation of In has not been demonstrated in the past.  
 406 Finally, we should note that In is poorly incorporated  
 407 at the highly Al-rich stripes. We attribute this to the

408 low miscibility of In with Al-containing alloys [43]. At  
 409 the  $\{112\}$  nanofacets In competes with Al for the avail-  
 410 able binding sites. This is particularly relevant for the  
 411  $(112)$ B-polar directions because Al tends to accumulate  
 412 more in these positions.

413 **Conclusion** In conclusion, we have demonstrated a  
 414 significant (300 meV) redshift of the emission energy of  
 415 single-photon emitters embedded in the shell of core-  
 416 shell GaAs-AlGaInAs NWs. The redshifting mechanism  
 417 is based on the In incorporation in the AlGaAs shell alloy  
 418 to form a quaternary semiconductor of lower bandgap.  
 419 The spatial distribution of different species in the shell is  
 420 determined by several high-resolution techniques of com-  
 421 positional analysis. In addition to the well-known Al and  
 422 Ga segregation, we find evidence of novel wedge-shaped  
 423 In-rich segregation. We explain the distribution of the  
 424 different species in the quaternary alloy with their dif-  
 425 fusion on the NW sidewalls driven by crystal-phase, po-  
 426 larity and strain. Finite-element simulations provide in-  
 427 sight in the role of strain to drive the segregation of In.  
 428 These findings advance the understating of the segrega-  
 429 tion phenomena in quaternary alloys, as required to take

430 full advantage of the additional degrees of freedom that  
431 they offer.

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