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A revised (earliest Vallesian) age for the hominoid-bearing locality of Can Mata 1 based on new magnetostratigraphic and biostratigraphic data from Abocador de Can Mata (Vallès-Penedès Basin, NE Iberian Peninsula)

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Abstract

The Abocador de Can Mata (ACM) composite stratigraphic sequence (els Hostalets de Pierola, Vallès-Penedès Basin, NE Iberian Peninsula) has yielded a diverse primate assemblage from the late Aragonian (Middle to Late Miocene). Detailed litho-, bio- and magnetostratigraphic control has enabled an accurate dating of these fossil remains. Comparable data, however, were lacking for the nearby locality of Can Mata 1 (CM1), which yielded a dryopithecine canine of a female individual. Given the lack of hipparionin equids and giraffids, CM1 has been correlated to the latest Aragonian (Mammal Neogene [MN] zone MN7+8). Here we revise the age of CM1 based on fieldwork and associated paleomagnetic samplings undertaken in 2018–2021. Our results extend the ACM composite sequence upward and indicate that CM1 correlates to the earliest Vallesian (MN9). The updated ACM sequence is ~300 m-thick and comprises 12 magnetozones correlated to subchrons C5Ar.1r to C5n.2n (~12.6 to 11.1 Ma; latest MN6 to earliest MN9, late Aragonian to earliest Vallesian). CM1 is correlated to C5r.1r (11.146–11.056 Ma), with an interpolated age of 11.11 Ma, thus postdating the dispersal of hipparionin horses into the Vallès-Penedès Basin—which is correlated to the previous subchron C5r.1n, with an interpolated age of 11.18 Ma, and by definition marks the beginning of the Vallesian. CM1 also minimally postdates the earliest record of giraffids at ACM—representing their earliest well-dated occurrence in the basin—being correlated to C5r.1n with an interpolated age of 11.11 Ma. We conclude that CM1 has an earliest Vallesian (MN9) age of ~11.1 Ma, intermediate

between the Aragonian dryopithecins and the Vallesian hispanopithecins. Ongoing paleontological surveillance at ACM thus offers the prospect to yield additional earliest Vallesian ape remains, which are essential to clarify their taxonomic allocation as well as to confirm whether hispanopithecins evolved locally from dryopithecins rather than immigrating from elsewhere during MN9.

Key words: Fossil apes; Dryopithecinae; Bretxa de Can Mata; Magnetostratigraphy; Paleomagnetism; Vallesian.

1. Introduction

1.1. Fossil primates from els Hostalets de Pierola

Thanks to continued paleontological surveillance of the digging activity during the construction of a dump between 2002 and 2014, a diverse primate assemblage (Alba, 2012; Marigó et al., 2014; Alba et al., 2017; DeMiguel et al., 2021; Fortuny et al., 2021) has been recovered from the Middle to Late Miocene stratigraphic sequence of Abocador de Can Mata (ACM; Alba et al. 2006, 2011a, 2017) in the fossiliferous area of els Hostalets de Pierola (Catalonia, Spain). Both small-bodied catarrhines (Alba et al., 2010, 2012a, 2015) and large-bodied apes—*Pierolapithecus* (Moyà-Solà et al., 2004; Almécija et al., 2009; Pérez de los Ríos et al., 2012; Hammond et al., 2013; Pina et al., 2014), *Dryopithecus* (Moyà-Solà et al., 2009a; Alba and Moyà-Solà, 2012; Pina et al., 2019, 2020), and *Anoiapithecus* (Moyà-Solà et al., 2009b; Alba et al., 2013)—have been recovered there, spanning from ~12.4–12.3 to ~11.6 Ma (Alba et al., 2017). While litho-, bio-, and magnetostratigraphic correlation has enabled an accurate dating of the fossil primate remains from ACM (Moyà-Solà et al.,

2009a; Casanovas-Vilar et al., 2011, 2016a, 2016b; Alba et al., 2017), this does not apply to all the previous hominoid finds from the same area.

A mandibular fragment with M₂–M₃ (IPS1826+1827) found in 1941 at Can Vila (Villalta Comella and Crusafont Pairó, 1941a, 1944) is the holotype of '*Sivapithecus*' *occidentalis*, which is considered a species inquirenda (Alba et al., 2020; Fortuny et al., 2021)—i.e., of uncertain taxonomic validity due to the lack of comparable elements in other species. This specimen comes from minimally older levels than ACM/BCV1 (Barranc de Can Vila 1, the type locality of *Pierolapithecus catalaunicus*) and thus has an age of ~12.0 Ma (Alba et al., 2013, 2017, 2020). However, the age of the two other previous finds is more uncertain. Thus, the stratigraphic provenance of an M² (MGSB48486) from Can Mata s.l. (Van der Made and Ribot, 1999), currently assigned to *Dryopithecus fontani* (Alba et al., 2013, 2017, 2020; Fortuny et al., 2021), is unknown. Therefore, it can only be considered broadly younger than the hominoid finds from ACM (Alba et al., 2013), but it is impossible to determine whether it is latest Aragonian (Mammal Neogene [MN] zone MN7+8) or earliest Vallesian (MN9) in age (i.e., ~11.3–11.0 Ma). The third previous hominoid find from the area is a C₁ of a female individual (IPS1766; Fig. 1) from the locality of Can Mata 1 (CM1) or Bretxa de Can Mata. It was originally assigned to *Hispanopithecus laietanus* (Crusafont-Pairó and Golpe-Posse, 1973; Golpe-Posse, 1993) or *Dryopithecus laietanus* (Begun et al., 1990; Harrison, 1991; Andrews et al., 1996), but most recently it was left unassigned to genus as Hominidae indet. (Casanovas-Vilar et al., 2011; Alba, 2012) or Dryopithecinae indet. (Alba et al., 2013, 2020; Marigó et al., 2014). This canine has been excluded from the hypodigm of *H. laietanus* because this species is not recorded until much later (~10.0 Ma; Alba et al., 2018) and also because the specimen cannot be directly compared with any of the ACM dryopithecine taxa. Although the provenance of IPS1766 from CM1 is well documented

(Crusafont-Pairó and Golpe-Posse, 1973), its age is uncertain owing to the lack of magnetostratigraphic data and poor stratigraphic control of historical finds of hipparionin equids from the area of els Hostalets de Pierola.

1.2. Uncertainties about the age of Can Mata localities

The localities of Can Mata are situated nearby the farmhouse of Can Mata de la Garriga within the fossiliferous area of els Hostalets de Pierola, in the Penedès sector of the Vallès-Penedès Basin (Fig. 2; Casanovas-Vilar et al., 2016a; for details on the geological background, see Casanovas-Vilar et al., 2008; Alba et al., 2009, 2017; Moyà-Solà et al., 2009a). Three Can Mata localities (CM1, CM2 and CM3) have been reported in the literature (Supplementary Online Material [SOM] S1.1), although only CM1 has yielded a rich fossil assemblage. The locality of CM1 (Fig. 3) was discovered in 1941 close to the farmhouse of Can Mata de la Garriga (Fig. 4), being considered pre-Vallesian given the absence of hipparionins and giraffids (Villalta Comella and Crusafont Pairó, 1941b, 1943; Crusafont and Truyols, 1954). In turn, CM2 was initially considered Vallesian in age due to the record of giraffids and the presence of *Hippotherium* molars in the same area (Crusafont Pairó, 1952; Crusafont and Truyols, 1954). Finally, CM3 corresponds to an isolated find of *Hippotherium* ~200 m in NE direction from CM1 (Fig. 4c; Alba et al., 2006; SOM S1.1), which represents the oldest hipparionin find of known provenance from the area of Can Mata. Over the years, and only with a few exceptions (Mein, 1990; Andrews et al., 1996; Pickford, 2013), most authors have correlated CM1 to the latest Aragonian on biostratigraphic grounds (Crusafont-Pairó and Golpe-Posse, 1973; Golpe-Posse, 1974, 1982, 1993; Agustí et al., 1985; de Bruijn et al., 1992; Alba et al., 2006, 2011a, 2017; Moyà-Solà et al., 2009b; Casanovas-Vilar et al., 2011, 2016a, 2016b; Robles et al., 2013), and the same applies to CM2

(Crusafont Pairó and Golpe Posse, 1971; Golpe-Posse, 1974; Agustí et al., 1985, 1997, 2001) with the exception of de Bruijn et al. (1992).

The published faunal list of large mammals from CM1 needs to be updated and revised to confirm the identity of multiple taxa that is confusing according to published sources (see S1.2 for further details). However, it is generally indicative of an age close to the Aragonian/Vallesian boundary. The frequent correlation of CM1 to MN7+8 instead of MN9 almost entirely relies on the lack of *Hippotherium* and giraffids. These large mammals have even been formally considered in the local biozonation of the Vallès-Penedès Basin (Casanovas-Vilar et al., 2016b). Even the claim that CM2 was located above CM1 (e.g., Agustí et al., 1985) might be exclusively based on the fact that Golpe-Posse (1974) roughly correlated CM1 with La Grive-Saint-Alban in France (MN7+8) and CM2 with Castell de Barberà (then considered latest MN7+8, but currently considered earliest MN9; Alba et al., 2019), based on the presence of giraffids in the latter locality. Along the same line of reasoning, Agustí et al. (1985) remarked that not only *Hippotherium*, but also giraffids—traditionally considered to predate the arrival of hipparionins in the Vallès-Penedès Basin (e.g., Golpe-Posse, 1974; Crusafont Pairó, 1975; Agustí et al., 1985, 1997, 2001)—were missing from CM1.

The important role played by hipparionins and giraffids in the local biozonation of the Vallès-Penedès Basin across the Aragonian/Vallesian transition might seem surprising given that most of the local biozones are based on rodents (e.g., Agustí et al., 1997; Casanovas-Vilar et al., 2016b). However, besides the role played by hipparionins in the definition of the Vallesian, there are other historical reasons that explain this situation. In particular, CM1 only yielded very scarce micromammal remains, contrasting with the hundreds of micromammal specimens recovered from the so-called ‘bloc de marga’ (marl block)—

originally attributed to nearby Mas d'Ocata (Schaub, 1944, 1947), but actually coming from Can Flequer (Agustí, 1981; Agustí et al., 1985), which is Vallesian based on the presence of *Hippotherium*. Agustí and colleagues (Agustí, 1981, 1999; Agustí and Gibert, 1982; Agustí et al., 1985, 1997, 2001; Agustí and Moyà-Solà, 1991) relied on a composite micromammal faunal list for all the purportedly pre-Vallesian outcrops from the area, collectively denoted as Hostalets Inferior. These authors noted similarities with micromammals from the unambiguously Vallesian localities (Hostalets Superior) and concluded that their composition remained virtually unchanged through the Aragonian/Vallesian boundary (see also Casanovas-Vilar et al., 2011, 2016a, 2016b). Needless to say, this practice might lead to circular reasoning (particularly if Vallesian levels are mistaken as pre-Vallesian due to the absence of *Hippotherium*). Nevertheless, ever since then, it has been customarily accepted that the latest Aragonian and earliest Vallesian rodent assemblages from the Vallès-Penedès Basin are virtually indistinguishable (Agustí et al., 1997; Alba et al., 2006; Casanovas-Vilar et al., 2011, 2016a, 2016b).

Alba et al. (2006) remained skeptical about whether CM2 was latest Aragonian or earliest Vallesian, but given that CM1 is a well-sampled locality, they assumed that the absence of *Hippotherium* was not a sampling bias and accepted its customary correlation to MN7+8. Under such an assumption, Moyà-Solà et al. (2009b: Fig. 5) and Casanovas-Vilar et al. (2011) correlated CM1 to the upper part of the reverse polarity subchron C5r.2r (postdating the short-duration normal polarity cryptochron C5r.2r-1), implying an age comprised between 11.263 and 11.188 Ma (Ogg, 2020). Casanovas-Vilar et al. (2016a) further correlated CM1 to the *Democricetodon crusafonti*–*Hippotherium* interval subzone of the Vallès-Penedès Basin (Casanovas-Vilar et al., 2016b), whose top boundary in this basin is defined by the first local occurrence (FLO) of *Hippotherium* at 11.18 Ma (Garcés et al., 1996, 1997; Agustí et al., 1997;

Casanovas-Vilar et al., 2016a, 2016b; Alba et al., 2019). In turn, based on the assumption that CM3 was located ~40 m above CM1 (Alba et al., 2006), Moyà-Solà et al. (2009a) tentatively correlated the former to C5r.1r, although such a correlation was revised to C5n.2n by Casanovas-Vilar et al. (2016b: Fig. 3), following some corrections on the then available composite magnetostratigraphic sequence for ACM and Ecoparc de Can Mata (ECM; Alba et al., 2012b).

The tentative magnetostratigraphic correlation of CM1 by Moyà-Solà et al. (2009b) was based on its geographic location and its lithostratigraphic correlation with paleomagnetic samplings in the nearby Riera de Claret—as CM1 was then located outside the area of the Can Mata landfill. Nevertheless, Robles et al. (2013: 999) noted that “paleomagnetic sampling would be required in order to confirm correlation [of CM1] with C5r.2r [...] rather than C5r.1n”. The resumption of the digging activity at ACM in 2018, to enlarge further the landfill, provided the opportunity to extend the sequence upward toward the Vallesian—including the stratigraphic section where CM1 is located, as well as to measure in situ the stratigraphic distance between CM1 and CM3. New paleomagnetic samples were therefore taken in 2020 to update the ACM composite stratigraphic sequence (so as to accurately date the fossil remains recovered between 2018–2020) and provide a revised magnetostratigraphic correlation and interpolated ages for CM1 and CM3. Given their biochronological implications, the oldest well-dated hipparionin and giraffid remains from the area of els Hostalets de Pierola are also reported. Finally, our results are discussed in relation to phylogenetic and paleobiogeographic uncertainties regarding the evolutionary origin of Vallesian dryopithecines from Western Europe.

2. Methods

2.1. Location and paleomagnetic sampling of Can Mata 1

Although no detailed coordinates of CM1 are available from the literature, some of the authors of this paper were able to locate it during a field visit in February 2004 (Fig. 3c), with the help of published pictures (Villalta Comella and Crusafont Pairó, 1941b: Pl. XVI, 1943: Pl. XV; see our Fig. 3a–d). Back in 2004, CM1 was well outside the area of the then planned enlargement of the Can Mata landfill (Fig. 4), but this situation changed when a new enlargement phase started after the publication of Alba et al. (2017). The new outcrops exposed during 2019–2020 allowed us to extend the ACM composite sequence upward toward the Vallesian. The new paleomagnetic samplings were performed in July 2020, including the short section of CM1, which was finally affected by the construction works of the landfill in 2021. With the help of an excavator machine, the outcrop was excavated under paleontological surveillance in May 2021 (Fig. 5). This allowed us to find the original spot of the excavation filled with debris (Fig. 5g), thereby confirming that the CM1 bone bed had been completely excavated by Crusafont and collaborators several decades ago.

Magnetostratigraphic sampling was carried out in the perimeter of the newly excavated Cell 13 and the future Cell 14 of the landfill. A total of 44 sites were sampled along a 78 m-thick composite section, which yielded an average sampling resolution of 1.77 m/site. The locality of CM1 was located at meter 62 of the section. 32 paleomagnetic samples were collected along the ~65 m-thick Cell 13 (C13) section, 6 sites in the Préstec de Terres de l'Abocador (PTA) sector, and another 6 sites in the overlying 15 m-thick Bretxa de Can Mata section (Fig. 6). Fresh unaltered red-brown mudstones and fine-grained sandstones were drilled by means of a battery-powered drill with a water-cooled diamond bit. The samples were oriented in situ with a magnetic compass fixed to an orienting device with a clinometer. Magnetic measurements were carried out at the Paleomagnetic Laboratory of

192 Barcelona (CCiTUB-Geo3Bcn CSIC). Laboratory procedures consisted of stepwise thermal
193 demagnetization and measurement of the natural remanent magnetization of one sample
194 per site. A thermal demagnetizer TD48EU (ASC Scientific) and a 755SRM superconducting
195 rock magnetometer (2G Enterprises) were used. The measurement routine included 12
196 increasing temperature steps, applied to all samples up to 630 °C or complete
197 demagnetization (100, 200, 250, 350, 400, 440, 480, 520, 550, 580, 610, and 630 °C).
198 Magnetic susceptibility was measured with a KLY-2 (Geofyzika Brno) after each
199 demagnetization step with no remarkable changes observed (SOM Table S1), which ensures
200 no significant formation of undesired secondary magnetic minerals during heating. Samples
201 showed a low temperature component that was usually removed below 200 °C, which we
202 interpret as a viscous recent overprint (SOM Fig. S1). Above this temperature, characteristic
203 components were calculated by principal component analysis (Kirschvink, 1980) after visual
204 inspection of the demagnetization diagrams using 'Paleomagnetism.org 2.2.0' interpretation
205 portal (Koymans et al., 2016). Characteristic components were usually defined between 250
206 and 580 °C (SOM Table S2), pointing to magnetite as the main remanence carrier. Only in
207 three samples the characteristic component reached higher temperatures (610 or 630 °C),
208 suggesting a mixture of magnetite and hematite. Characteristic components were ranked
209 into three classes according to their quality. Quality 1 components (42%) showed a straight
210 and complete decay directed towards the origin. Class 2 components (37%) were clearly
211 isolated from the viscous component and were defined by at least three demagnetization
212 steps showing an incomplete demagnetization trend. When only two temperature steps
213 were used to calculate the characteristic components, or polarity was difficult to assess due
214 to unstable behavior and erratic trajectories, they were classified as quality 3 (21%), and
215 were not considered for further magnetostratigraphic analysis.

Mean paleomagnetic directions and associated Fisher parameters (Fisher, 1953) were calculated using the PmagPy software (Tauxe et al., 2016). No fold test could be performed due to the homogeneous gentle dip of the beds, but a significant better agreement with the expected north-directed paleomagnetic direction was obtained after correction for the tilt of beds (SOM Fig. S2; SOM Table S3). The total number of samples of quality 1 and 2 resulted in a good stratigraphic resolution, so that occurrence of missing magnetozones can be considered unlikely. The obtained paleomagnetic directions were used to calculate the latitude of the virtual geomagnetic pole (VGP) for each site. Positive VGP latitudes were interpreted as normal polarity whereas negative VGP latitudes were interpreted as reversed polarity, and a sequence of polarity zones was represented accordingly in the local magnetostratigraphy (black for normal magnetozones and white for the reversed ones).

2.3. Correlation methods

Lithostratigraphic correlation As in Alba et al. (2017), we mostly relied on intensely red-colored horizons interpreted as paleosols for lithostratigraphic correlation in the ACM composite series. The composite lithostratigraphic sequence reported by Alba et al. (2017) was constructed on the basis of 38 local sections ranging from 7 to 150 m in thickness, of which 5 included the magnetostratigraphic data previously reported by Moyà-Solà et al. (2009a). Here we provide an updated version of the ACM composite sequence that incorporates 13 additional local sections, of which the one including CM1 has new magnetostratigraphic data. The new lithostratigraphic sections largely overlapped and were correlated by means of 9 additional correlation levels, resulting in a total of 51 correlation levels in the composite sequence of ACM.

Magnetostratigraphic correlation and interpolated ages The magnetostratigraphic correlation of the ACM composite sequence with the Geomagnetic Polarity Time Scale (GPTS; Ogg, 2020) is based on the 366 paleomagnetic samples reported by Moyà-Solà et al. (2009a) and analyzed by Alba et al. (2017) plus the 44 samples newly reported in the present work. The samples originally analyzed by Moyà-Solà et al. (2009a) were distributed in three different sections from ACM—Barranc de Can Vila, Vial Intern, and Camí de Can Vila—Cells 3–4 (the latter being the composite of three different sections)—plus the nearby section of Riera de Claret. The newly sampled section of Cells 13–14 (including CM1) is added here. The stratigraphic position of each ACM fossil locality was recorded relative to its closest lithostratigraphic section(s), thereby being correlated on lithostratigraphic grounds to the composite magnetostratigraphic sequence.

As in Alba et al. (2017), given short-term variation in sedimentation rates (due to the intermittent nature of sedimentation in alluvial environments), an estimated age was calculated for each locality using linear interpolation based on the sedimentation rate computed for each subchron (Barry et al., 2002; Casanovas-Vilar et al., 2014: Fig. 3); for the top and bottom subchrons, sedimentation rate was approximated by that of the adjacent subchron. Alba et al. (2017) concluded that errors associated with interpolated ages are lower than 100 ka, so that estimated ages for ACM localities can be considered reliable to the nearest 0.1 Ma. However, interpolated ages to the nearest 0.01 Ma are also provided, as they are still informative with regard to the relative age among ACM localities.

Biostratigraphic correlation The local biozonation of the Vallès-Penedès Basin during the late Aragonian and Vallesian is based on Casanovas-Vilar et al.'s (2016b) biostratigraphic scheme, which provides a formal diagnosis and description of each zone and is almost exclusively based on rodents—except for the FLOs of *Hippotherium* and giraffids. The fossils

from els Hostalets de Pierola referred to in this paper are housed in the Institut Català de Paleontologia Miquel Crusafont (ICP) in Sabadell (Catalonia, Spain). The fossils described are referred to by their catalog numbers, which are preceded by the acronym 'IPS' ('Institut de Paleontologia de Sabadell', the former name of the ICP).

3. Results

3.1. Magnetostratigraphy

Our paleomagnetic results for the Cells 13–14 section, where CM1 is located, allowed us to identify the presence of six magnetozones (three normal and three reversed, with the lowermost one having a reversed polarity) defined at least by two consecutive sites of the same polarity (Fig. 6). The reliability of the magnetic zonation could be validated by a cross-correlation with the neighboring overlapping sections of the Riera de Claret and Camí de Can Vila–Cells 3–4 (SOM Fig. S3; Moyà-Solà et al., 2009a; Alba et al., 2017), which can be traced by lithostratigraphic correlation (SOM Fig. S4), supporting a primary origin of the magnetization.

A unique correlation of the local magnetostratigraphy of the Cells 13–14 section with the GPTS (Ogg, 2020) can be established based on its integration with other sections in the ACM area (Fig. 7), and refines earlier results (Alba et al., 2017). Given these constraints, the occurrence of two very short intervals of normal polarity within a period of dominantly reversed polarity points to a correlation of the Cells 13–14 section with the upper end of chron C5r. The lower normal magnetozones would represent the cryptochron C5r.2r-1n, and the second one subchron C5r.1n. The normal magnetozones on top of Cells 13–14 section would then be correlated to the lowermost portion of subchron C5n.2n. Despite not being

listed in the time scale until recently, the cryptochron C5r.2r-1 is a short event of comparable duration to C5r.1n (Abdul Aziz et al., 2003; Ogg, 2020).

The lower portion of C5r.2r was previously the uppermost magnetozone recorded at the ACM composite sequence s.s., which terminated a few meters above locality ACM/CCV1 (Camí de Can Vila 1; Moyà-Solà et al., 2009a; Alba et al., 2017; locality indicated in Fig. 7 and SOM Fig. S3). Therefore, the five polarity reversals documented at the Cells 13–14 section had not previously been recorded within the area of the landfill back in 2014—although they had been documented at the Riera de Claret section, as well as partly (the first two reversals) at the Camí de Can Vila section outside the landfill (Fig. 7; SOM Fig. S3; Moyà-Solà et al., 2009a; Alba et al., 2017). The outcrops exposed subsequently within the area affected by the landfill allowed us to extend the ACM composite section s.s. up to the lowermost portion of the long normal polarity subchron C5n.2n (11.056–9.984 Ma), typical of the early Vallesian. The locality of CM1 is located in the uppermost reversed polarity zone recorded (Figs. 7 and 8), correlated with C5r.1r (11.146–11.056 Ma). The latter subchron approximately extends from meters 278 to 293 of the ACM composite sequence (Table 1), whereas CM1 is located on meter 284, resulting in an interpolated age of 11.11 Ma.

In turn, although no paleomagnetic samples were taken there, computations based on both geographical location and field measurements indicate that CM3 is located ~8 m above CM1 (based on the coordinates reported in Fig. 4 and a strike/dip of 180°/13° W), thus being similarly correlated with C5r.1r with an interpolated age of 11.06 Ma. Finally, these results made us reconsider previous correlations between the ECM and ACM composite sections (Casanovas-Vilar et al., 2016b). Unfortunately, magnetostratigraphic data are only available for the uppermost part of the section, where no polarity reversal is recorded. However, the stratigraphic thickness between the base of the ECM and CM1, based on geographical

location (see coordinates in Fig. 4) and the aforementioned strike/dip figure, is estimated to be 63 m. According to this, the base of the ECM section would be roughly equivalent to ACM/CCV1 (~11.5 Ma), with the oldest localities (e.g., ECM/VCE-A1 [VCE = Variant de la Carretera de l'Ecoparc]) being latest Aragonian (correlated to the upper portion of C5r.2r, with an estimated age of ~11.2 Ma) instead of earliest Vallesian (~10.9 Ma) as previously thought (Alba et al., 2012b). In contrast, the remaining ECM localities would be earliest Vallesian, but older than previously thought (Alba et al., 2012b; Casanovas-Vilar et al., 2016b). In particular, based on the sedimentation rate calculated for the preceding subchron, ECM/VCE-Bb would have an estimated age of ~10.9 Ma, whereas ECM/VCE-B2 would date to ~10.8 Ma and ECM/VCE-C1 to ~10.7 Ma. Only the estimated ages for the ECM localities mentioned later in this paper are provided, as fieldwork during the following years will likely enable a refinement of their dating based on a more secure tie-point of the ECM base to the upper portion of the ACM composite section.

3.2. Biochronology

Small mammals The small mammal collection from the Aragonian levels of els Hostalets de Pierola includes rodent, lagomorph, and insectivoran remains that are mostly labeled as 'Hostalets inferior' or 'Hostalets vindobonien'. Only three specimens are accompanied by the label 'Can Mata 1' and can be unambiguously related to this locality. The remainder of the small mammal sample certainly mixes material from different sites, as already stated by Agustí and Gibert (1982), but probably includes additional specimens from CM1. The assemblage is overwhelmingly dominated by the cricetids *Hispanomys lavocati*, *Hispanomys dispectus*, and *Megacricetodon ibericus*. In addition, the cricetid *Dem. crusafonti*, the terrestrial sciurid *Csakvaromys bredai* and the lagomorph *Prolagus oeningensis* also occur in

lower numbers. The same taxa are also present in the collection from Vallesian levels of the area (Hostalets Superior), largely originating from the ‘bloc de marga’ (Can Flequer).

Given that the cricetids *His. lavocati*, *His. dispectus*, and *M. ibericus* are present in the small mammal collections from both Hostalets Inferior and Superior, it is not surprising that no significant difference was previously noticed across the Aragonian/Vallesian boundary in the area (e.g., Agustí et al., 1985). However, our magnetostratigraphic results—indicating that CM1 is earliest Vallesian and providing an accurate dating for ACM localities—suggest otherwise. At ACM, the FLO of *M. ibericus* is recorded from ACM/C1-E9 (with an interpolated age of 12.29 Ma), whereas the FLO of *His. lavocati* is recorded from various localities dated to ~11.6 Ma (e.g., ACM/C6-A2, with an interpolated age of 11.61 Ma), indicating that both species appeared well before the Aragonian/Vallesian boundary. In contrast, *His. dispectus* has not been reported from the ACM composite section (Casanovas-Vilar et al., 2016b), at least as defined in Alba et al. (2017)—i.e., up to ~11.4 Ma.

Hispanomys dispectus is neither present at the oldest micromammal site from ECM (ECM/VCE-A1), which records *His. lavocati* and, as noted above, is probably latest Aragonian in age (~11.2 Ma, correlated with the upper portion of C5r.2r). Similarly, the roughly coeval locality ACM/PTA-A2, with an interpolated age of 11.21 Ma (Alba et al., 2022), records *His. cf. lavocati* but not *His. dispectus*, thus suggesting that the latter species might have not appeared until the Vallesian.

Large mammals We were unable to find among the ICP collections any *Hippotherium* remains labeled as coming from CM3. However, in all probability, the specimen found by Jordi Agustí in 1976 corresponds to a left mandibular fragment with very worn M₂–M₃ (IPS30897; SOM Fig. S5d–f). This specimen was found in the ICP collections inventoried together with an isolated left M₃ (currently IPS124307) of the same taxon (SOM Fig. S5a–c)

despite belonging to two different individuals, and collectively labeled as coming from “Can Mata, levels superior than those of the breccia” (our translation from Spanish). According to its discoverer, the specimen from CM3 most likely corresponds to IPS30897 (pers. comm. to J.G., December 2021). However, pending a more complete revision of all the *Hippotherium* remains from Hostalets Superior, both specimens are taken into account here (see description in SOM S2.1). The presence of a well-developed pli caballinid in IPS124307 hints at similarities between the Can Mata sample and Pannonian C remains from the Vienna Basin (Austria) assigned to *Hippotherium* sp. (Bernor et al., 2017, 2021).

With regard to giraffids, CM1 slightly postdates the oldest well-dated record of this family at ACM, which corresponds to two M³ germs from a single individual (IPS121859; SOM Fig. S6a–f) found isolated (i.e., not associated with additional fossil remains) in April 2020 at sector ACM/C13/C14-A. The specimen’s provenance is located only 1.3 m below CM1 (i.e., approximately at meter 283 of the ACM sequence), thereby being correlated to the same subchron and having the same interpolated age as CM1. In turn, the giraffid material from ECM consists of a few teeth from locality ECM/VCE-Bb (SOM Fig. S6g–r), which despite the lack of *Hippotherium* has an unambiguous Vallesian age somewhat younger than CM3 (~10.9 Ma, updated from Alba et al., 2012b; see above for further details). A description of IPS121859, together with other giraffid material from the area of els Hostalets de Pierola, is provided in SOM S2.2. Although the scarce remains from ACM and ECM are not informative enough to provide a taxonomic attribution to genus rank, they likely correspond to the same taxon previously identified as *Palaeotragus* sp. by Crusafont Pairó (1952) based on Can Mata (including CM2) remains. Deciphering the taxonomic affinities of these giraffids is hindered by the lack of ossicones and the fact that they display a mix of apparently primitive and derived characters. Some of the latter are similar to those

present in taxa allied to the samotherium-sivatherium lineage, although future systematic work is needed to test this hypothesis. In any case, it is noteworthy that the Can Mata giraffid displays some dental and postcranial differences relative to *Decennatherium pachecoi*, present in the early Vallesian of the Calatayud-Montalbán and Duero basins (Crusafont-Pairó, 1952; Ríos et al., 2016), despite being only slightly older than the material from its type locality (Nombrevilla 1, which has an age of ~10.8 Ma; Van Dam et al., 2014). This supports Crusafont Pairó's (1952) early assessment that a different (even if poorly known) giraffid species is recorded during the earliest Vallesian in the Vallès-Penedès Basin.

4. Discussion

4.1. The age of Can Mata localities based on the updated Abocador de Can Mata sequence

The previous ACM composite sequence, based on the fieldwork undertaken between 2002 and 2014 (Alba et al., 2017), was 234 m in thickness, comprised seven magnetozones that were correlated with C5Ar.1r to C5r.2r, and covered a time span between ~12.6 and ~11.4 Ma. With the addition of the new magnetostratigraphic section of Cells 13–14 and other lithostratigraphic sections excavated since 2017, the ACM composite sequence now has a thickness of 300 m (Fig. 8) and includes up to 12 magnetozones (Table 1), ranging in time from ~12.6 to ~11.1 Ma. The newly added magnetostratigraphic section allows us to refine the interpolated ages of previously known ACM localities from the uppermost portion of the composite sequence. These interpolated ages were formerly based on the stratigraphic situation of the top of C5r.2r in the nearby sequence of Camí de Can Vila (Alba et al., 2017), which has been slightly corrected based on our new results (Table 1). Furthermore, our magnetostratigraphic analyses for the Cells 13–14 section enable an accurate dating of the new fossil localities (and isolated finds) from ACM that have been

discovered since 2017, as well as of CM1. A comprehensive updated list of ACM localities and sectors, based on the subchron chronological and stratigraphic boundaries reported in Table 1, is provided in SOM Tables S4 and S5, respectively. In turn, an updated list of ACM primate-bearing localities as compared with CM1 is also reported in Figure 8 and SOM Table S6.

Our results necessitate us to reconsider the previously published correlations for the hominoid-bearing locality of CM1, as they support an unambiguous correlation with C5r.1r. These results contradict the commonly held notion that CM1 (and CM2) are late Aragonian (MN7+8) in age (see the Introduction) and indicate an estimated age of ~11.1 Ma for both CM1 and CM3. Indeed, CM1 slightly postdates the short normal magnetic polarity subchron C5r.1n (11.188–11.146 Ma), where *Hippotherium* is first recorded in the Vallès-Penedès Basin (Garcés et al., 1996, 1997; Agustí et al., 1997; Casanovas-Vilar et al., 2016a, 2016b; Alba et al., 2019). According to this, CM1 and CM3 would be slightly younger than both Creu de Conill 20 and Castell de Barberà (~11.2 Ma), but older than most of the ECM composite sequence—with the exception of its lowermost portion (e.g., ECM/VCE-A1), which on lithostratigraphic grounds appears to be latest Aragonian in age (contra Alba et al., 2012b; Casanovas-Vilar et al., 2016b). The age of CM2 cannot be estimated with certainty, given that its exact geographic location and stratigraphic position above CM1 are unknown. However, given the geographical proximity to the Can Mata farmhouse, CM2 must be roughly penecontemporaneous with CM1 (~11.1 Ma), in agreement with Crusafont-Pairó's (1952) original dating based on the presence of hipparionin remains in the same area (although not in the same stratigraphic horizon).

4.2. Biostratigraphic and biochronological implications

Small mammals The earliest Vallesian micromammal assemblages have previously been found to be taxonomically identical to the late Aragonian ones (see Section 1.2 as well as SOM S3.1 for further details). Similarly, the local zone H of the Calatayud-Montalbán Basin was originally correlated to the earliest Vallesian (Daams and Freudenthal, 1981) but is currently considered to span the latest Aragonian and the early Vallesian (Álvarez Sierra et al., 2003; Van Dam et al., 2014; García-Paredes et al., 2016). This interpretation is based on the fact that neither Nombrevilla 9 nor 10 (~11.2 Ma) record *Hippotherium*, which is not found there until Nombrevilla 1 (~10.8 Ma). A pre-Vallesian age for Nombrevilla 9 is further supported by its magnetostratigraphic correlation to the uppermost portion of C5r.2r with an interpolated age of 11.20 Ma, whereas Nombrevilla 10 correlates with C5r.1n with an interpolated age of 11.18 Ma (Garcés et al., 2003; Van Dam et al., 2014) has not yielded large mammal remains (Álvarez-Sierra et al., 2003). *Megacricetodon ibericus* is recorded in the latest Aragonian and the earliest Vallesian both in the Vallès-Penedès Basin and the Calatayud-Montalbán Basin (SOM S3.1), being thus irrelevant for determining the Aragonian/Vallesian boundary. Only the later appearance of the cricetid *Cricetulodon hartenbergeri* denotes a significant change in the rodent faunas (SOM S3.2), roughly coinciding with the extinction of the cricetids *Dem. crusafonti* and *M. ibericus*, which were generally dominant elements of the latest Aragonian and earliest Vallesian rodent faunas in both basins (SOM S3.2).

Nevertheless, the revised correlation of CM1 to the earliest Vallesian highlights the potential biostratigraphic significance of *Hispanomys* species during the latest Aragonian and earliest Vallesian, which had been previously neglected due to incorrect correlation of various localities (including CM1) to the latest Aragonian. In the Calatayud-Montalbán and Duero basins (López-Guerrero et al., 2014, 2019; Van Dam et al., 2014; Castillo et al., 2018;

454 SOM S3.3), *His. lavocati* is restricted to the late Aragonian (zone G3), whereas *His.*
455 *nombrevillae* is recorded around the Aragonian/Vallesian transition (zone H), and *His.*
456 *aragonensis* appears later in early Vallesian (zone I). In contrast, in the Vallès-Penedès Basin
457 *His. lavocati* persists into the earliest Vallesian, which is characterized by two different
458 species of this genus: *His. dispectus* and *Hispanomys daamsi*. Neither of these species has
459 been thus far recorded from Aragonian levels from ACM, being restricted to earliest
460 Vallesian localities of the Vallès-Penedès Basin (and also the Empordà basin in the case of
461 *His. dispectus*; Casanovas-Vilar et al., 2016a; SOM S3.3). The only exception is the presence
462 of *His. dispectus* at Hostalets Inferior. However, the fact that the latter assemblage
463 potentially includes Vallesian material from levels previously considered to predate the
464 dispersal of *Hippotherium* (such as CM1) make it uncertain the record of the species before
465 the Vallesian. In contrast, later early Vallesian localities from ECM record *His. aragonensis*,
466 which is first recorded at ~10.8 Ma (Alba et al., 2012b; Casanovas-Vilar et al., 2016b; SOM
467 S3.3). In summary, both *His. dispectus* and *His. daamsi* appear conclusively recorded only
468 from a short time period during the earliest Vallesian (accurately dated localities between
469 ~11.2–11.0 Ma), thus potentially being a biochronological marker of the Vallesian in its type
470 area (see SOM S3.3 for further details).

471 Hipparionins European *Hippotherium* originated from a single dispersal event of
472 *Cormohipparion* from North America sometime after 11.5 Ma (Woodburne, 2007, 2009;
473 Bernor et al., 2017, 2020, 2021). *Cormohipparion* has been recorded in the Old World since
474 10.8 Ma in Asia (Bernor et al., 2003; Wolf et al., 2013) and 10.7 Ma in Africa (Bernor et al.,
475 2004; Bernor and White, 2009), but not in Europe, where *Hippotherium* is recorded earlier
476 and almost simultaneously in the Vallès-Penedès Basin in Spain at ~11.2 Ma (Garcés et al.,
477 1996; Agustí et al., 1997; Casanovas-Vilar et al., 2016a, 2016b; Alba et al., 2019) and in the

478 Pannonian C sites of the Vienna Basin (Gaiselberg, Atzelsdorf, and Mariathal) in Austria at
479 ~11.4–11.0 Ma (Bernor et al., 2017). An earlier age of ~11.5 Ma has been proposed for the
480 first occurrence of hipparionins in China (Fang et al., 2016; Sun et al., 2022), thus being
481 slightly older than the record of *Hippotherium* in the Vienna and Vallès-Penedès basins.
482 However, the corresponding magnetostratigraphy (Fang et al., 2016) is missing many
483 relevant subchrons, such that alternative correlations—including C5r.1r, which could imply a
484 younger age roughly coeval with the European *Hippotherium* datum—cannot be discounted
485 (SOM S4.1).

486 Different species of *Hippotherium* are recorded in the Vallesian of the Vallès-Penedès and
487 Calatayud-Montalbán basins. In the former, *Hippotherium catalaunicum* is customarily
488 recognized as a distinct species (Bernor et al., 1980, 1996, 2021; Woodburne and Bernor,
489 1980), whereas *Hippotherium koenigswaldi* is recorded at Nombrevilla 1 (~10.8 Ma). Both
490 species have been interpreted as derived in facial and dental morphology compared with
491 *Hippotherium primigenium* from Central Europe (Bernor et al., 1996). A third species,
492 *Hippotherium melendezi*, is recorded from MN9 of the Duero Basin (SOM S4.2), although its
493 cranial morphology rather supports its inclusion in *Hipparion* s.s. (Bernor et al., 1996), which
494 is otherwise first recorded at MN10 in Iran (Bernor et al., 2017). Available evidence thus
495 indicates that vicariant speciation events occurred during MN9 after the initial dispersal of
496 *Hippotherium* into the Iberian Peninsula. Nevertheless, the identity of the earliest Vallesian
497 (~11.2–11.0 Ma) hipparionins from the Vallès-Penedès Basin, which are older than either
498 *Hip. catalaunicum* and *Hip. koenigswaldi* and include the remains from els Hostalets de
499 Pierola (this paper) and Castell de Barberà (Rotgers and Alba, 2011; Alba et al., 2019),
500 remains to be clarified (Alba et al., 2019). The dental specimens from Can Mata reported in
501 this paper show a similar morphology to the Pannonian C *Hippotherium* remains from the

Vienna Basin, which are more plesiomorphic than those of *Hipp. primigenium* (Woodburne 2009; Bernor et al., 2017), and may be thus identified as *Hippotherium* sp.

Giraffids and other large mammals The giraffid remains from CM2 described by Crusafont-Pairó (1952), and the whole giraffid sample from earliest Vallesian sites of the Vallès-Penedès Basin, have traditionally been assigned to *Palaeotragus* instead of *Decennatherium* (e.g., Agustí et al., 1985, 2001)—albeit it is noteworthy that Crusafont-Pairó (1952) noted similarities with species currently included in *Schansitherium* sp. (SOM S4.3). Nevertheless, like the new dental remains reported here from ACM and ECM, the remainder of the Vallès-Penedès giraffid sample from the earliest Vallesian only includes scarce dental and postcranial remains, which are not informative enough to justify an attribution to genus rank. Both *Schansitherium* and *Decennatherium* belong to the samotheresivathere clade and the former only appears slightly more basal than *Decennatherium* in cladistic analyses (Ríos et al., 2017). *Decennatherium* has been reported not only from the Vallesian of Spain (Ríos et al., 2016, 2017) but also of Pakistan (Ríos et al., 2019; see SOM S4.3 for further details), where it is first tentatively recorded at 13.5 Ma (Ríos et al., 2019). The combination of more primitive features than in *Decennatherium* with some apparently derived characters raises the possibility that '*Palaeotragus*' sp. from the Vallès-Penedès Basin might belong to a very basal species of *Decennatherium*, or at least to a genus closely related to the latter. Therefore, both '*Palaeotragus*' sp. from the Vallès-Penedès and *Dec. pachecoi* from elsewhere in Spain might have originated from the earliest giraffid immigrants that dispersed from the east.

Despite the lack of giraffids and *Hippotherium* at CM1, our results conclusively indicate that this locality postdates, even if slightly, the earliest well-dated record of giraffids in the area of els Hostalets at ~11.1 Ma (Fig. 8) and elsewhere in the Vallès-Penedès Basin (Castell

de Barberà at ~11.2 Ma; Alba et al., 2019). This indicates that the lack of both giraffids and hipparionins at CM1 is either a sampling bias and/or must be attributed to local paleoenvironmental reasons. Giraffid remains assigned to *Palaeotragus* have also been cited from multiple Vallès-Penedès localities of controversial age (latest Aragonian or earliest Vallesian; see SOM 4.3. for further details). Many of these localities were initially correlated to the Vallesian given that Crusafont Pairó (1950) conceived this land mammal age as characterized by the presence of both hipparionins and giraffids. However, the subsequent more detailed characterization of the Vallesian by Crusafont Pairó and Truyols Santonja (1960) only emphasized the dispersal of hipparionins as a biochronological marker—even though the association with giraffids was still noted. Ultimately, Crusafont Pairó and Golpe Posse (1971) defined a new pre-Vallesian biozone characterized by the lack of hipparionins and the presence of giraffids, to which CM2 and other localities of uncertain age around the Aragonian/Vallesian boundary were correlated (e.g., Golpe-Posse, 1974; Crusafont Pairó, 1975; SOM S4.3). This view was subsequently followed by Agustí and coauthors (Agustí et al., 1985, 1997, 2001; Agustí and Moyà-Solà, 1990), leading to the idea that giraffids of genus *Palaeotragus* would constitute a biochronological marker of MN7+8 and that only the giraffid *Decennatherium* would be indicative of the Vallesian (see SOM S4.3 for additional details). In contrast, during the last decade an earliest Vallesian age has been favored for many of these localities (Casanovas-Vilar et al., 2016a, 2016b; Alba et al., 2019). Our results further indicate that a latest Aragonian age is no longer tenable for CM2, as CM1 correlates to the Vallesian despite the lack of *Hippotherium*. The fact that giraffids, as shown here, are recorded slightly before *Hippotherium* at els Hostalets de Pierola might simply result from insufficient sampling, paleoecological biases, and the lack of stratigraphic control for most historical finds of *Hippotherium* in the area. At ACM, giraffids have yet to

be found in the late Aragonian levels (Alba et al., 2006, 2017), suggesting that giraffids and hipparionins might have dispersed approximately at the same time (as previously supported by Casanovas-Vilar et al., 2016b).

In the Calatayud-Montalbán Basin, the earliest record of giraffids corresponds to *Dec. pachecoi* from Nombrevilla 9, correlated to C5r.2r with an interpolated age of 11.20 Ma and been considered latest Aragonian in age due to the lack of *Hippotherium* (Álvarez-Sierra et al., 2003; Van Dam et al., 2014). If such a correlation is accurate, it implies that giraffids might have immigrated into Iberia during the latest Aragonian immediately before the entry of *Hippotherium* (as recorded in the Vallès-Penedès Basin at 11.18 Ma). On the other hand, Nombrevilla 9 already records several typically Vallesian elements, such as *Dec. pachecoi*, *Deinotherium giganteum*, *Tetralophodon longirostris*, and *His. nombrevillae* (Álvarez Sierra et al., 2003; Van Dam et al., 2014; García-Paredes et al., 2016; Ríos et al., 2016). Given that the dispersal of *Hippotherium* seems to have been roughly synchronous across Europe (Vasiliev et al., 2011; Bernor et al., 2017; this paper), it is conceivable that *Hippotherium* had already dispersed into Iberia when Nombrevilla 9 was deposited, being absent from this locality because of the same (likely paleoenvironmental) reasons as to why this taxon is missing or very rare in many earliest MN9 sites of the Vallès-Penedès Basin.

While the presence of *Hippotherium* indicates a Vallesian age by definition, its absence cannot be taken as unambiguously indicating a pre-Vallesian age (Alba et al., 2019). For example, no hipparionin remains have thus far been identified from the nearby ECM composite series despite the unambiguous Vallesian age of most of the series (Alba et al., 2012b; this paper). The cases of CM2 and Castell de Barberà further highlight the risk of relying almost exclusively on the purported absence of *Hippotherium* when making biostratigraphic correlations for sites located close to the Aragonian/Vallesian boundary—

particularly because other large mammal taxa, such as giraffids, also appear indicative of a Vallesian age.

There are other eastern immigrants that first appear in the Iberian Peninsula and elsewhere in Europe coinciding with the beginning of the Vallesian, such as the machairodontine felid *Machairodus* and the suid *Propotamochoerus palaeochoerus* (see SOM S4.4 for further details). These taxa co-occur with *Hippotherium* at Creu de Conill 20 are also present at presumably Vallesian levels from els Hostalets de Pierola, although it is uncertain when they are first recorded there (SOM S4.4). Further research would be required to discern if the dispersal of *Hippotherium* and other Vallesian faunal elements was synchronous or part of a more protracted faunal turnover across the Aragonian/Vallesian transition. Nevertheless, currently available data indicate that the large mammal assemblages from the earliest Vallesian of the Vallès-Penedès Basin were characterized by an array of new immigrants including at least *Hippotherium*, giraffids, *Propotamochoerus*, and *Machairodus*, even if these taxa do not always co-occur.

4.3. Implications for primate evolution

In the past decades, dryopithecines have been variously interpreted as pongines, hominines, and stem hominids, so that conclusively deciphering their closest phylogenetic affinities would be key for deciding among competing scenarios about the place of origin of crown hominids (see discussion in Almécija et al., 2021; Pugh, 2022). Not so long ago, late Middle and early Late Miocene dryopithecines from Europe were still included in a single genus *Dryopithecus* (e.g., Andrews et al., 1996; Begun, 2002), largely as a result of the limited knowledge of Middle Miocene dryopithecines at the time. This contrasts with the panoply of genera that are currently recognized, following the discovery of more complete

598 Middle Miocene ape remains at ACM in the 2000s, which prompted the realization that
599 dryopithecine diversity could not be accommodated within the single genus (Begun, 2009;
600 Moyà-Solà et al., 2009a). Two alternative proposals were then put forward: Moyà-Solà et al.
601 (2004, 2009a, 2009b) distinguished various Middle Miocene genera (*Pierolapithecus*,
602 *Anoiapithecus*, *Dryopithecus*) and included the Late Miocene material from Spain and
603 Hungary into a single genus *Hispanopithecus*; in contrast, Begun (2009) and Begun et al.
604 (2012) considered that *Pierolapithecus* and *Anoiapithecus* could be synonyms of
605 *Dryopithecus* and distinguished instead *Hispanopithecus* from *Rudapithecus* in the Late
606 Miocene. The distinction between the two latter genera is currently also favored by Alba,
607 Moyà-Solà and colleagues (e.g., Almécija et al., 2021; Fortuny et al., 2021; Urciuoli et al.,
608 2021), while the existence of more than a single genus at ACM is also generally
609 acknowledged (e.g., Böhme et al., 2019; Andrews, 2020). To recognize the fact that the Late
610 Miocene dryopithecines appear more derived and closely related than the Middle Miocene
611 ones, Alba (2012) distinguished the tribes Dryopithecini and Hispanopithecini, although this
612 picture became more complicated after the description of *Danuvius*, which is dated to ~11.6
613 Ma (Kirscher et al., 2016; Böhme et al., 2019) and thus roughly coeval with the youngest
614 dryopithecines from ACM (Alba et al., 2017).

615 On paleobiogeographic grounds, it would be most reasonable to assume that
616 hispanopithecines from the Vallesian evolved locally in central and/or western Europe from
617 Middle Miocene dryopithecine ancestors. However, this cannot be conclusively asserted,
618 owing to phylogenetic uncertainties surrounding the internal phylogeny of dryopithecines
619 coupled with the fragmentary nature of the hominoid fossil record in Europe throughout
620 the Aragonian/Vallesian boundary. Regarding phylogeny, it is generally considered that
621 hispanopithecines are more derived than dryopithecines (e.g., Alba, 2012), but it is currently

622 uncertain whether dryopithecines as a whole constitute a clade or a paraphyletic
623 assemblage (Almécija et al., 2021), because few cladistic analyses have analyzed
624 dryopithecins and hispanopithecins separately. Alba et al.'s (2015) analysis recovered
625 *Hispanopithecus* and *Pierolapithecus* as stem hominids, whereas Pugh's (2022) analyses
626 more conclusively showed that dryopithecines are best interpreted as stem hominids, with
627 hispanopithecins appearing more closely related to crown hominids than dryopithecins.
628 *Danuvius* appears intermediate in the degree of suspensory adaptations and other
629 morphological features between dryopithecins and hispanopithecins (Almécija et al., 2021),
630 thus favoring the regional continuity hypothesis by strengthening the view that the more
631 suspensory hispanopithecins might have evolved from the Middle Miocene dryopithecins.
632 However, *Danuvius* predates the beginning of the Vallesian by more than 0.5 Myr and the
633 earliest record of hispanopithecine genera by about 1.5 Myr. Therefore, *Danuvius* cannot be
634 automatically taken as an evolutionary link between dryopithecins and hispanopithecins—at
635 least, until a formal cladistic analysis shows that this genus is more closely related to the
636 latter.

637 The hypothetical phylogenetic link between dryopithecins and hispanopithecins is further
638 complicated by the fact that Pugh's (2022) analyses sometimes recovered a close link
639 between dryopithecines and kenyapithecines (*Kenyapithecus* and *Griphopithecus*), as
640 previously argued by Moyà-Solà et al. (2009b), whereas '*Lufengpithecus*' *hudienensis* from
641 the Late Miocene (8.2–7.2 Ma) of China (Kelley and Gao, 2012) tended to group with
642 hispanopithecins instead of *Lufengpithecus lufengensis* (the type species of the genus),
643 which clusters with other Asian hominids in a pongine clade (Pugh, 2022). This closer
644 phylogenetic of '*L.*' *hudienensis* with hispanopithecins instead of pongines agrees with the
645 lack of pongine synapomorphies in this species (Kelley and Gao, 2012) and the possession of

646 some dental similarities (in canine and incisor morphology; Begun and Kelley, 2016).
647 Coupled with faunal similarities suggesting “a strong biogeographic link between Europe
648 and China” by the Late Miocene (Begun, 2013: 143; see also Begun and Kelley, 2016),
649 currently available evidence suggests that dryopithecines might have been present in both
650 Europe and Asia during the Late Miocene and at the very least supports more complex
651 paleobiogeographic scenarios than customarily assumed (Gilbert et al., 2020). Although ‘*L.*
652 *hudienensis*’ is younger than the European hispanopithecins, the the oldest and less well-
653 known species included in the same genus—*Lufengpithecus keiyuanensis* from the Middle
654 Miocene of China (12.5–11.6 Ma; Kelley, 2002; Ji et al., 2013)—has not been analyzed from
655 a cladistic viewpoint yet and it is thus uncertain whether it belongs to the pongine radiation.
656 Such a possibility is supported by the fact that closer morphological similarities between the
657 samples of *L. keiyuanensis* from Xiaolongtan and ‘*L.*’ *hudienensis* from Yuanmou (as opposed
658 to *L. lufengensis* from Shihuiba, Lufeng) have been noted by previous authors, who even
659 included the two former within a single species (Gao, 1998; Harrison et al., 2002). Given that
660 the beginning of the Vallesian, as discussed above, is not only marked by the dispersal of
661 hipparionin horses, but also by the more or less simultaneous arrival of other taxa that
662 dispersed from the east—including the machairodontine felid *Machairodus*, giraffids, and
663 the suid *Propotamochoerus*—the possibility that hispanopithecins did not locally evolve
664 from dryopithecins but dispersed from elsewhere roughly coinciding with the
665 Aragonian/Vallesian boundary deserves further consideration.

666 Unfortunately, the identity of earliest Vallesian hominoids from Europe is not well
667 known. In particular, during the time interval between the latest Aragonian (~11.5 Ma, i.e.,
668 after *Danuvius*) and the earliest well-dated occurrences of *Hispanopithecus* (~10.3–10.0 Ma;
669 Alba et al., 2018) and *Rudapithecus* (~10.0–9.8 Ma; Casanovas-Vilar et al., 2011) in the

670 Vallesian, the record of European hominids is very scarce. In the Vallès-Penedès Basin,
671 earliest Vallesian dryopithecines are known from Castell de Barberà (~11.2 Ma; Alba et al.,
672 2019) and CM1 (~11.1 Ma; this paper). The former material was assigned to cf. *D. fontani*
673 based on its inferred body mass when Castell de Barberà was still correlated to the latest
674 Aragonian (Alba et al., 2011b; Alba, 2012; Marigó et al., 2014), but such an allocation is
675 tentative given the lack of craniodental specimens (Almécija et al., 2012; Alba et al., 2019;
676 Pina et al., 2019). With regard to the CM1 canine (Crusafont-Pairó and Golpe-Posse, 1973), a
677 secure taxonomic allocation is not possible (as it cannot be directly compared with ACM
678 dryopithecine taxa). Elsewhere in Europe, when poorly dated localities are excluded, the
679 earliest Vallesian material is restricted to an isolated molar from Mariathal, Austria (Thenius,
680 1982), which has yielded a form of *Hippotherium* (*Hip. aff. primigenium*) slightly more
681 derived than those from Gaiselberg and Atzelsdorf (*Hippotherium* sp.), suggesting a slightly
682 younger age within Pannonian C (but still be older than 11 Ma; Bernor et al., 2017). The
683 Mariathal hominoid molar was formerly assigned to *Dryopithecus brancoi* by several authors
684 (Thenius, 1982; Begun, 2002), or alternatively to *Dryopithecus carinthiacus* by Andrews et al.
685 (1996), but more recently it was considered of uncertain attribution (Casanovas-Vilar et al.,
686 2011; Pickford, 2012).

687 Discovering additional fossil great ape finds from the time span comprised between
688 ~11.5 and ~10.5 Ma is thus required to determine their taxonomic identity as well as to
689 further clarify the phylogenetic relationships between the late Aragonian dryopithecines,
690 *Danuvius*, '*L. hudienensis*' and other species of *Lufengpithecus* from China, and the Vallesian
691 hispanopithecines from Europe. Given that the fossiliferous accumulations from both Castell
692 de Barberà (Almécija et al., 2019) and CM1 (this paper) are exhausted, fieldwork efforts
693 should be redirected to other earliest Vallesian sites. In this regard, the site of Creu de Conill

20 (representing the earliest appearance of *Hippotherium* in western Europe) has some potential, even though it displays marked paleoenvironmental differences as compared with Castell de Barberà (Alba et al., 2019) and has not yielded any hominoid remains yet despite several successive fieldwork campaigns since 2016. Thus far, Vallès-Penedès hominoids dating to the Vallesian were considered to be restricted to the Vallès Sector of the basin. Redating CM1 is thus important because it conclusively shows that these primates were also recorded in the earliest Vallesian of the Penedès Sector, thereby opening new research prospects for the future. In particular, ongoing works of paleontological surveillance during the new phase of enlargement of the Can Mata landfill—mostly focused on deposits that cover the Aragonian/Vallesian transition (~11.5–11.0 Ma)—represent one of the best opportunities to find additional dryopithecine remains that might clarify their taxonomic and phylogenetic status in years to come.

5. Conclusions

We revisit the age of the hominoid-bearing locality of CM1 (or Bretxa de Can Mata), where a dryopithecine canine of a female individual was found in the 1970s. Can Mata 1 is located in the Penedès sector of the Vallès-Penedès Basin (NE Iberian Peninsula), in the fossiliferous area of els Hostalets de Pierola. This locality has traditionally been considered latest Aragonian (pre-Vallesian) in age, supposedly predating the earliest record of giraffids (CM2) and hipparionins (CM3) in the area. Currently, CM1 is included with the area affected by the enlargement works of the Can Mata landfill (i.e., the ACM macrosite), which thanks to paleontological surveillance and rescue excavations has delivered tens of thousands of fossil vertebrate remains from the late Aragonian during the last two decades. Here we report new magnetostratigraphic data that include CM1 and extend the ACM composite

section upward toward the Vallesian. The updated composite section is 300 m-thick and spans from ~12.6 to 11.1 Ma (latest MN6 to earliest MN9), with CM1 being correlated to subchron C5r.1r, with an estimated age of 11.1 Ma. Our results indicate that, despite slightly predating the oldest well-dated record of *Hippotherium* in the area of els Hostalets de Pierola (CM3), CM1 is younger than the first appearance datum of this taxon in the Vallès-Penedès Basin (correlated with C5r.1n, with an estimated age of 11.2 Ma), on which the definition of the Vallesian land mammal age hinges. Our results thus unambiguously indicate that CM1 does not correlate to the latest Aragonian (MN7+8)—as previously assumed based on the absence of *Hippotherium* and giraffids—but to the earliest Vallesian.

We note that *Hippotherium* remains from Can Mata show resemblances to the roughly coeval remains from Pannonian C sites of the Vienna Basin (Austria). We further show, based on new remains from ACM, that the oldest well-dated record of giraffids in the area of Can Mata is not younger (as previously thought) but slightly older than CM1, with a similar estimated age of 11.1 Ma. A critical assessment of Vallès-Penedès localities with giraffids, previously considered latest Aragonian owing to the lack of *Hippotherium*, indicates that this view is no longer tenable. In contrast, hipparionins and giraffids might have dispersed more or less simultaneously across Europe ~11.2 Ma, together with other large mammal immigrants from eastern Eurasia. All in all, our results highlight the need to stop considering the lack of *Hippotherium* as a reliable criterion for correlation with the latest Aragonian without the aid of magnetostratigraphic data or, at least, without taking into consideration other biochronologically important taxa. These include other large mammals and the cricetid *Hispanomys*, since based on currently available data both *His. dispectus* and *His. daamsi* might be characteristic of the earliest Vallesian.

Further research is necessary to ascertain to what extent the Aragonian/Vallesian transition was a relatively protracted faunal turnover or a more punctuated event. From a paleobiogeographic viewpoint, this has potential implications for hominoid evolution, because it is currently uncertain whether Vallesian dryopithecines evolved locally from their late Aragonian counterparts or immigrated from elsewhere at the beginning of the Vallesian (along with other eastern immigrants such as giraffids, *Machairodus*, or *Propotamochoerus*). Testing these competing hypotheses is hampered by the incompleteness and fragmentary nature of the hominoid fossil record, which is particularly meager for the earliest Vallesian (including scarce remains from CM1, Castell de Barberà, and Mariathal). Our study thus highlights the need to strengthen fieldwork efforts focused on this particular time interval, approximately between 11.5 and 10.5 Ma. In this sense, the revised (earliest Vallesian) age for CM1 conclusively shows that dryopithecines were also present in this sector of the basin by this time. Therefore, currently ongoing paleontological surveillance at ACM offers the most promising prospects to recover additional fossil ape remains from this key interval in years to come. This might help disentangle the controversial phylogenetic relationships among dryopithecine taxa, as well as the potential link between hispanopithecins and Late Miocene species from China that appear to be excluded from the pongine radiation.

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1149

1150 **Figure captions**

1151

1152 **Figure 1.** Right C₁ of a female individual from CM1 IPS1766 (old Nos. IPS49, H.223, VP339),
 1153 in occlusal (a), lingual (b), mesial(c), buccal (d), and distal (e) views.

1154

1155 **Figure 2.** Simplified geological map of the Vallès-Penedès Basin: a) location of the Vallès-
 1156 Penedès Basin within the Iberian Peninsula; b) location of Abocador de Can Mata (denoted
 1157 by a black dot, including Can Mata 1) within the basin. Modified from Casanovas-Vilar et al.
 1158 (2016a: Fig. 1).

1159

1160 **Figure 3.** Photographs of Can Mata 1 (CM1) in the early 1940s (a–d) and in 2004 (e): a, b)
 1161 general views; c–e) detail of CM. The location of the fossiliferous level is indicated by the

paleontologists excavating in panels b and d (denoted by arrows); I.C.V. denotes the approximate location of CM1 in February 2004, although the excavation spot is covered by debris. Photographs in panels a and c by Crusafont, reproduced from Villalta Comella and Crusafont Pairó (1943: Pl. XV); photographs in panels b and d by Bataller, reproduced from Villalta Comella and Crusafont Pairó (1941b: Pl. XVI Figs. 1 and 2, respectively); photograph in panel e by D.M.A., © Institut Català de Paleontologia Miquel Crusafont.

Figure 4. Aerial photographs showing the location of CM1 relative to ACM and the nearby farmhouses of Can Mata de la Garriga and Can Vila; the ACM hominoid-bearing locality of BCV1 is also provided for reference. a, b) Orthophotos taken in 1946 (a) and 2019 (b) to the same scale; the extension of the Can Mata landfill is shown in semitransparency in green (up to 2017) and pink (portion enlarged since 2017 and until early 2021), while the extension of the works performed at ECM a decade ago is marked in orange. c) Orthophoto taken in 2021, depicting the inset marked in panel b at a greater scale; the location of CM1 is here denoted by a yellow circle, while that of CM3 is denoted by a white square; the base of the lithostratigraphic composite section of ECM is marked by a red triangle. Universal Transverse Mercator (UTM) coordinates (European Terrestrial Reference System 1989 [ETRS89]): CM1 = 31N 399765E, 4598400N, 285 m a.s.l.; CM3 = 31N 399848E, 4598585N, 312 m a.s.l.; base of ECM = 31 N 400221E, 4598659N, 325 m a.s.l. Base orthophotomaps: © Institut Cartogràfic i Geològic de Catalunya, downloaded from VISSIR v. 3.26 and reproduced with permission by means of a Creative Commons license CC BY 4.0 (ICGC, 2021). Abbreviations: ACM = Abocador de Can Mata; BCV1 = Barranc de Can Vila 1; CM1 = Can Mata 1; ECM = Ecoparc de Can Mata.

Figure 5. Photographs of Can Mata 1 (CM1) in 2001 (b–f) compared with 2004 (a): a) the dashed line indicates the break along which a large block of rock was detached sometime before 2021, while the circle denotes the approximate location of the CM1 fossiliferous accumulation (see Fig. 3); b, c) panoramic (b) and detailed (c) view of CM1 in January 2021, before the area was deforested; d–f) different perspectives on the excavation of CM1 with the help of a digger machine in May 2021; g) detail of the original excavation spot once exposed. Note that, within the circle in panel g, the stratigraphic layer of intense red coloration (denoted by a dashed line) is partly interrupted by a hole filled with debris that corresponds to the original excavation; this spot was excavated until the whole depth of the layer was exposed, and no further fossil finds were found anywhere nearby. The numbers have been added to facilitate the recognition of the same physiographic structures across different panels and despite different perspectives. Photographs by D.M.A. (a–c), J.M.R. (e, f), and V.V. (d, g), © Institut Català de Paleontologia Miquel Crusafont.

Figure 6. New magnetostratigraphic results for the Cells 13–14 section at Abocador de Can Mata, where the locality of Can Mata 1 (CM1, denoted by a four-pointed red star) is located. Sideways triangles denote the stratigraphic position of paleomagnetic samples. Normal polarity (positive VGP latitude) and reversed polarity (negative VGP latitude) magnetozones are denoted in black and white, respectively. Abbreviations: M = mudstones; S = sandstones; CG = conglomerates; VGP = virtual geomagnetic pole.

Figure 7. Magnetostratigraphic correlation of the Abocador de Can Mata (ACM) and Riera de Claret sections with the Geomagnetic Polarity Time Scale (GPTS) after Ogg (2020), as well as European land mammal ages (ELMA) and local biozones of the Vallès-Penedès Basin

(after Casanovas-Vilar et al., 2016b). The location of Can Mata 1 (CM1)—being correlated to the *Hippotherium*–*Cricetulodon hartenbergeri* interval subzone—is denoted by a four-pointed red star. Note that the boundary between this subzone and the preceding *Dem. crusafonti*–*Hippotherium* interval subzone is defined by the first local occurrence of *Hippotherium* in the Vallès-Penedès Basin at the base of C5r.1n (see Fig. 8 for further details). Short white lines in the GPTS represent short geomagnetic excursions within C5n.2n after Evans et al. (2007). Short horizontal lines to the left of the polarity column represent ‘tiny wiggles’ from sea floor magnetic anomaly stacks (Cande and Kent, 1992) that have been interpreted to represent geomagnetic polarity intervals (e.g., Roberts and Lewin-Harris, 2000). Note the inclusion in the GPTS of cryptochron C5r.2r-1 within C5r.2r (Abdul-Aziz et al., 2003; Ogg, 2020) and that the bottom boundary of the lowermost local biozone is unknown. See SOM Figure S3 for detailed magnetostratigraphic results of the three sampled sections that cover the uppermost portion of the ACM sequence. Abbreviations: M = mudstones; S = sandstones; CG = conglomerates.

Figure 8. Synthetic correlation of the composite local magnetostratigraphy of ACM and nearby areas with the Geomagnetic Polarity Time Scale (GPTS; after Ogg, 2020), as well as European land mammal ages (ELMA), Mammal Neogene units (MN9), and local biozones of the Vallès-Penedès Basin after Casanovas-Vilar et al. (2016b)—note that, for the local biozonation, a plus symbol denotes a concurrent range subzone, whereas an en dash denotes an interval subzone. The stratigraphic position of CM1, CM3 and the ACM primate-bearing localities (reported in SOM Table S4) is shown to the right on the composite lithostratigraphic column (black arrows), together with the first local occurrences of some biochronologically relevant taxa commented in the text (red arrows). The first local

1234 occurrence of hipparionins (*Hippotherium datum*) in the Vallès-Penedès Basin as recorded in
1235 Creu de Conill 20 [CCN20]) is denoted by a green discontinuous arrow. Modified after Alba
1236 et al. (2017: Fig. 3) based on the results reported in Figure 7 (see the caption for further
1237 details) and in the main text. The ACM primate-bearing localities were listed by Alba et al.
1238 (2017: Table 2) except for ACM/C8-Au, where a damaged upper premolar of Dryopithecinae
1239 indet. (IPS121187) was recently identified while preparing a deinotheri tooth, with an
1240 interpolated age of 11.77 Ma.

Table 1

Subchrons correlated with the ACM composite section, including their boundaries and duration according to the Geomagnetic Polarity Time Scale (Ogg, 2020), their stratigraphic position and thickness in the composite sequence, and their computed within-subchron sedimentary rate. Updated from Alba et al. (2017) based on fieldwork performed during 2018–2021. CM1 is located in meter 284 of the sequence and correlated with C5r.1r, which postdates the first subchron on the Vallesian (C5r.1n), while CM3 is located 8 m above CM1 (meter 292 of the sequence) and hence approximately correlated to the top of C5r.1r.

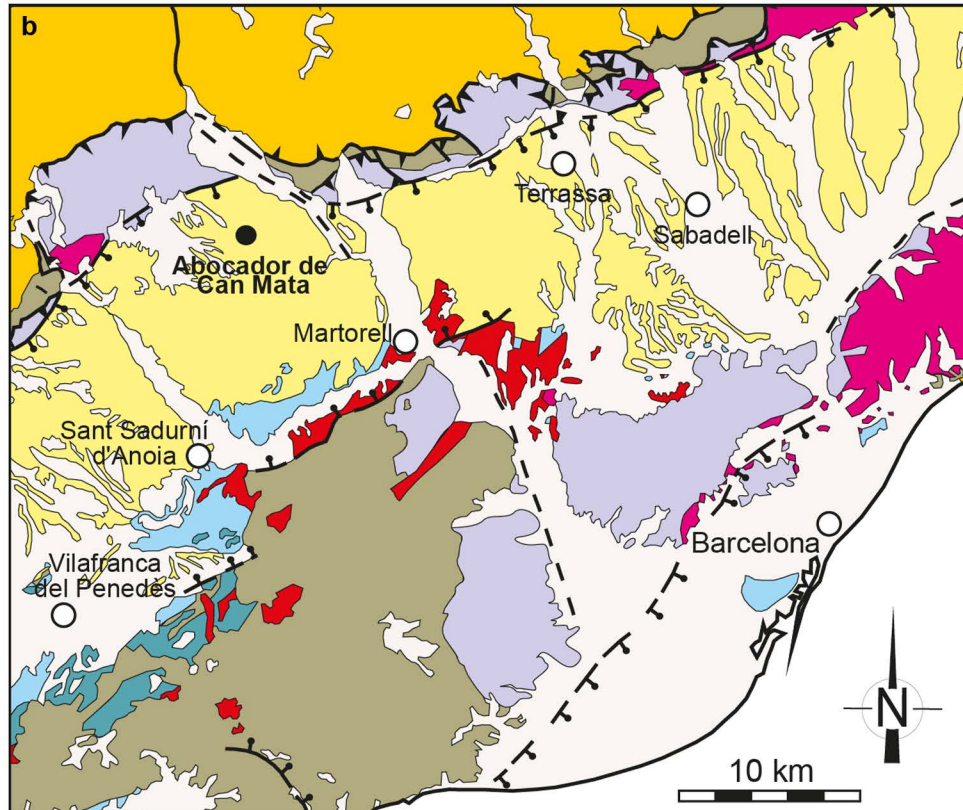
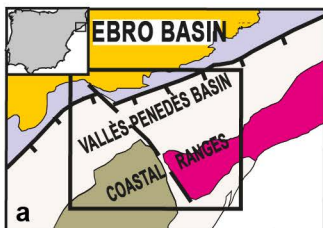
Subchron	Bottom (Ma)	Top (Ma)	Duration (Myr)	Bottom (m)	Top (m)	Thickness (m)	Sedimentation (cm/kyr)
C5n.2n	11.056	9.984	1.072	292.8 ^a	—	—	—
C5r.1r	11.146	11.056	0.090	277.5 ^a	292.8 ^a	15.3 ^a	17.0 ^a
C5r.1n	11.188	11.146	0.042	272.9 ^{a,c}	277.5 ^a	4.6 ^a	11.0 ^a
C5r.2r (2 nd portion)	11.263	11.188	0.075	251.5 ^a	272.9 ^{a,c}	21.4 ^a	28.5 ^a
C5r.2r-1	11.308	11.263	0.045	238.6 ^a	251.5 ^a	12.9 ^a	28.7 ^a
C5r.2r (1 st portion)	11.592	11.308	0.284	204.2 ^b	238.6 ^a	34.4 ^a	12.1 ^a
C5r.2n	11.657	11.592	0.065	177.1 ^b	204.2 ^b	27.1	41.7
C5r.3r	12.049	11.657	0.392	84.1 ^b	177.1 ^b	93.0	23.7
C5An.1n	12.174	12.049	0.125	68.6 ^b	84.1 ^b	15.5	12.4
C5An.1r	12.272	12.174	0.098	47.2 ^b	68.6 ^b	21.4	21.8
C5An.2n	12.474	12.272	0.202	15.9 ^b	47.2 ^b	31.3	15.5
C5Ar.1r	12.735	12.474	0.261	—	15.9 ^b	≥16.0	—

^a New results reported in this paper.

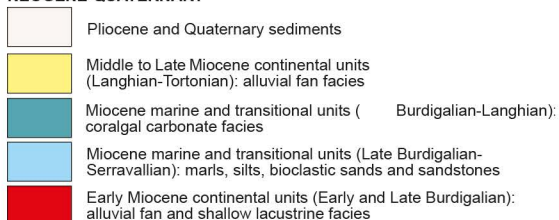
^b Results previously reported by Alba et al. (2017).

^b Alba et al. (2017) noted that the top of C5r.2r was not recorded at ACM but reported its top at 262.2 m based on the nearby sequence of Camí de Can Vila. New samplings within the ACM area sensu stricto enable to refine the subchron top within the ACM composite series.

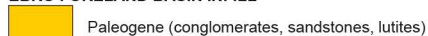




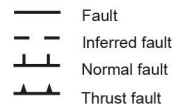
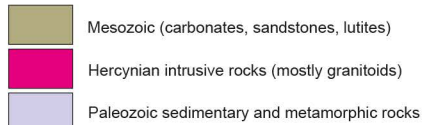
NEOGENE-QUATERNARY

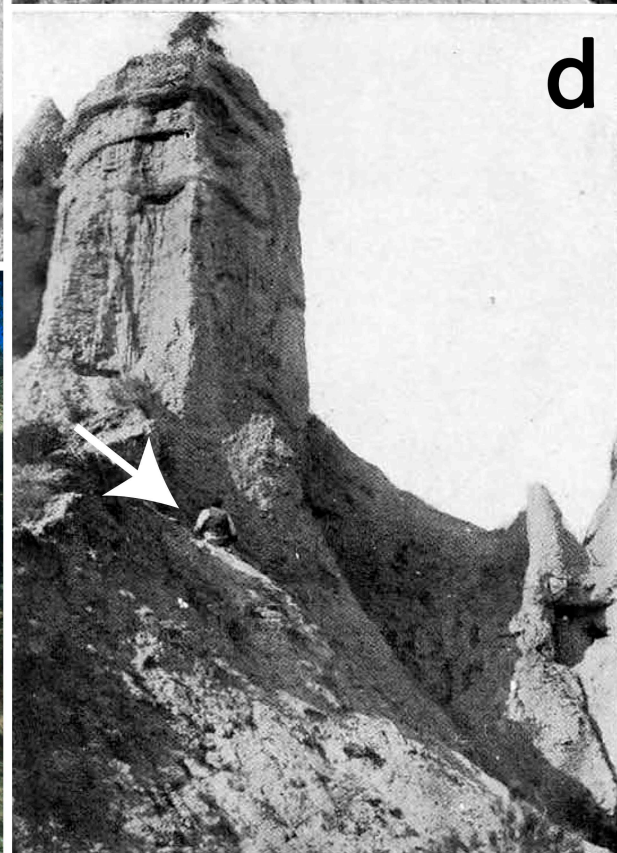
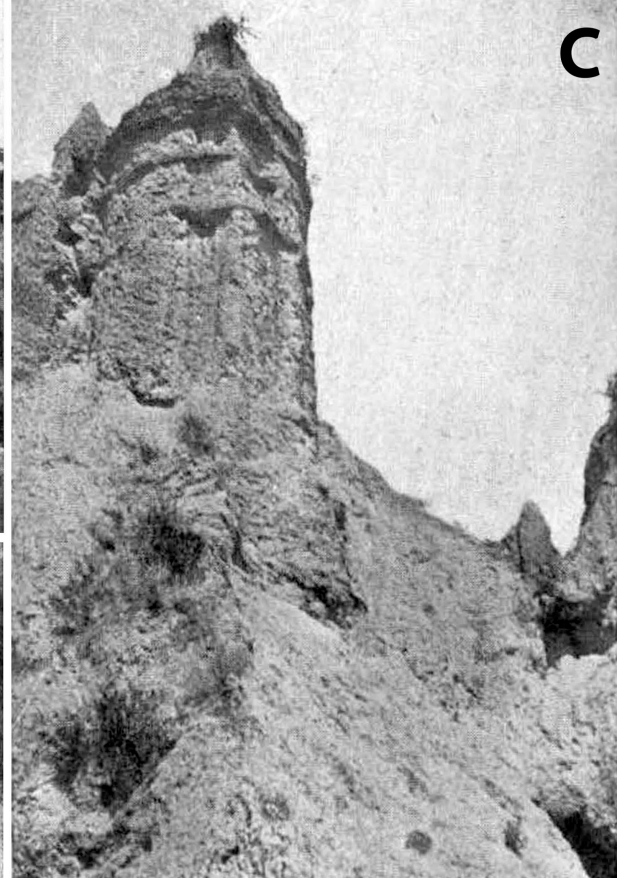
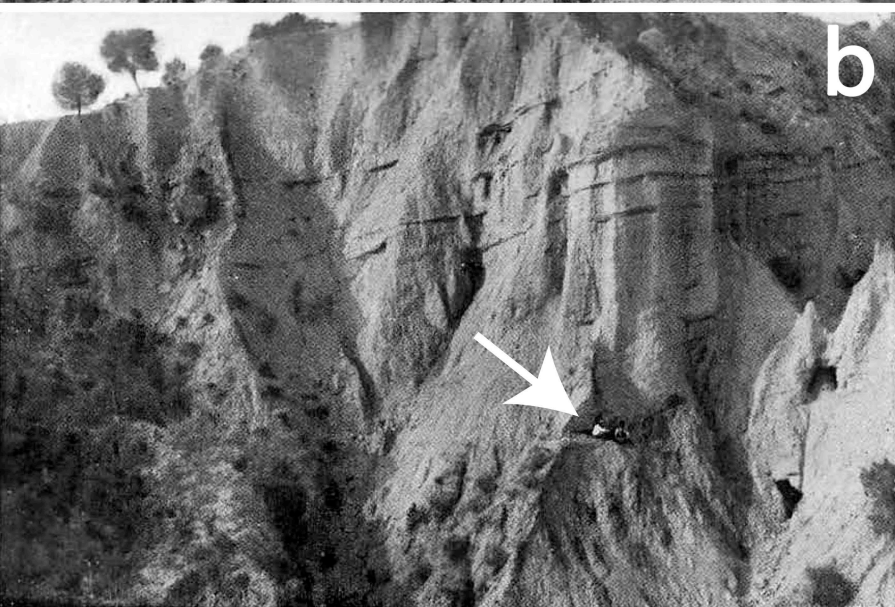


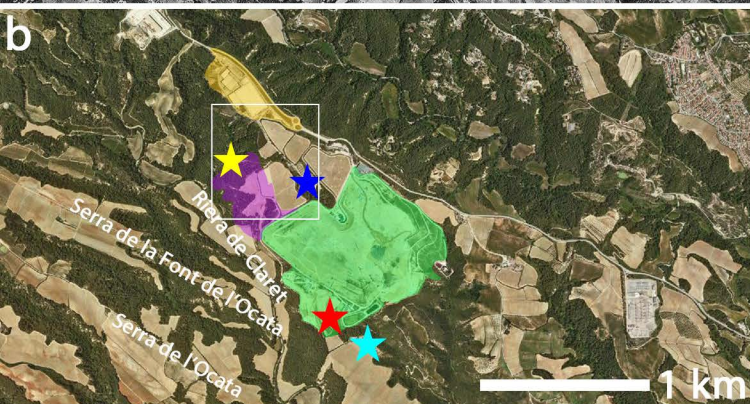
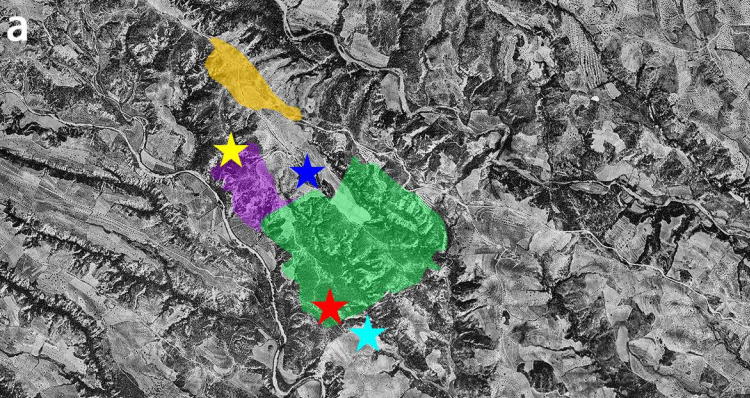
EBRO FORELAND BASIN INFILL



PALEOZOIC BASEMENT AND MESOZOIC COVER



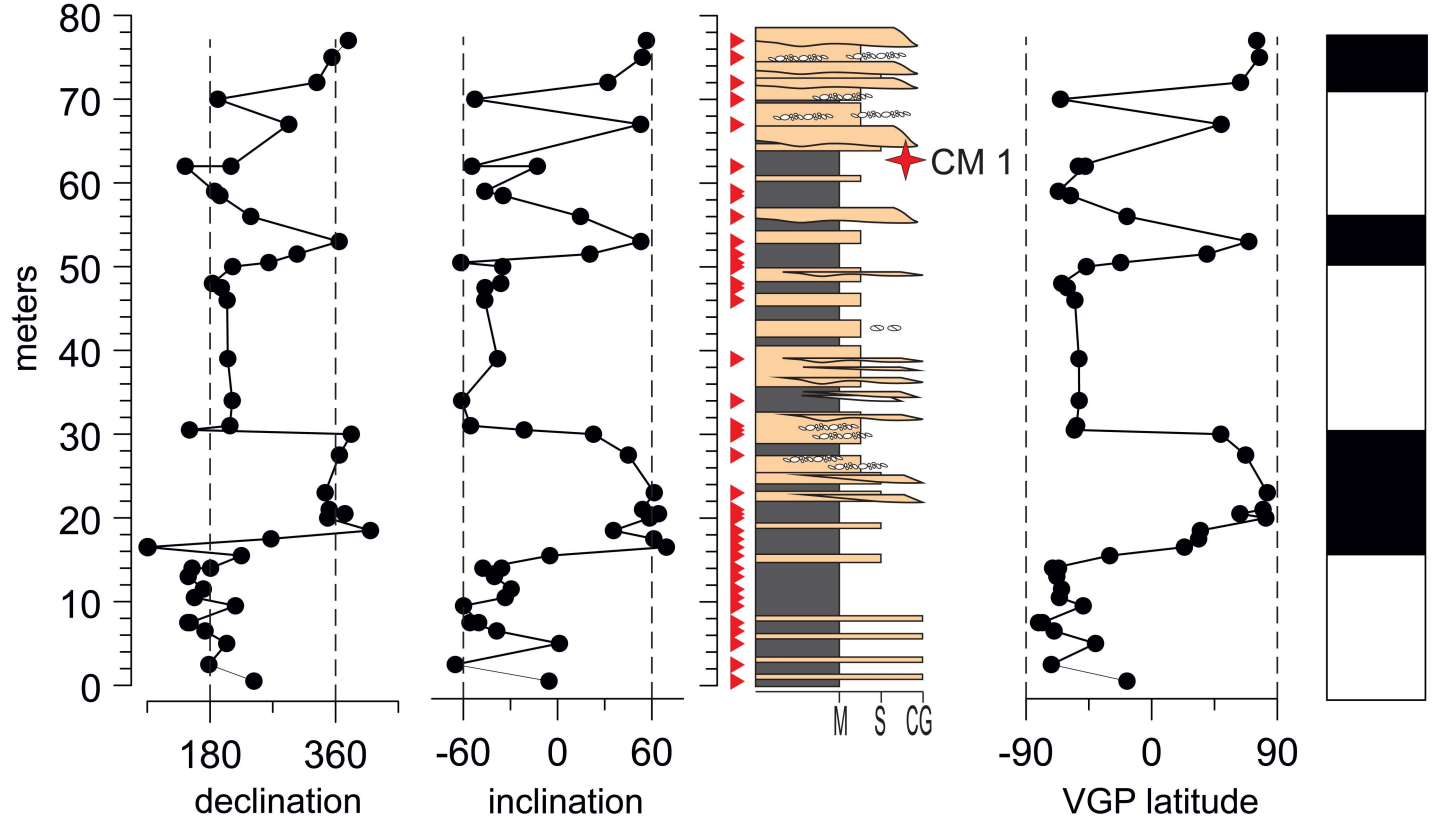


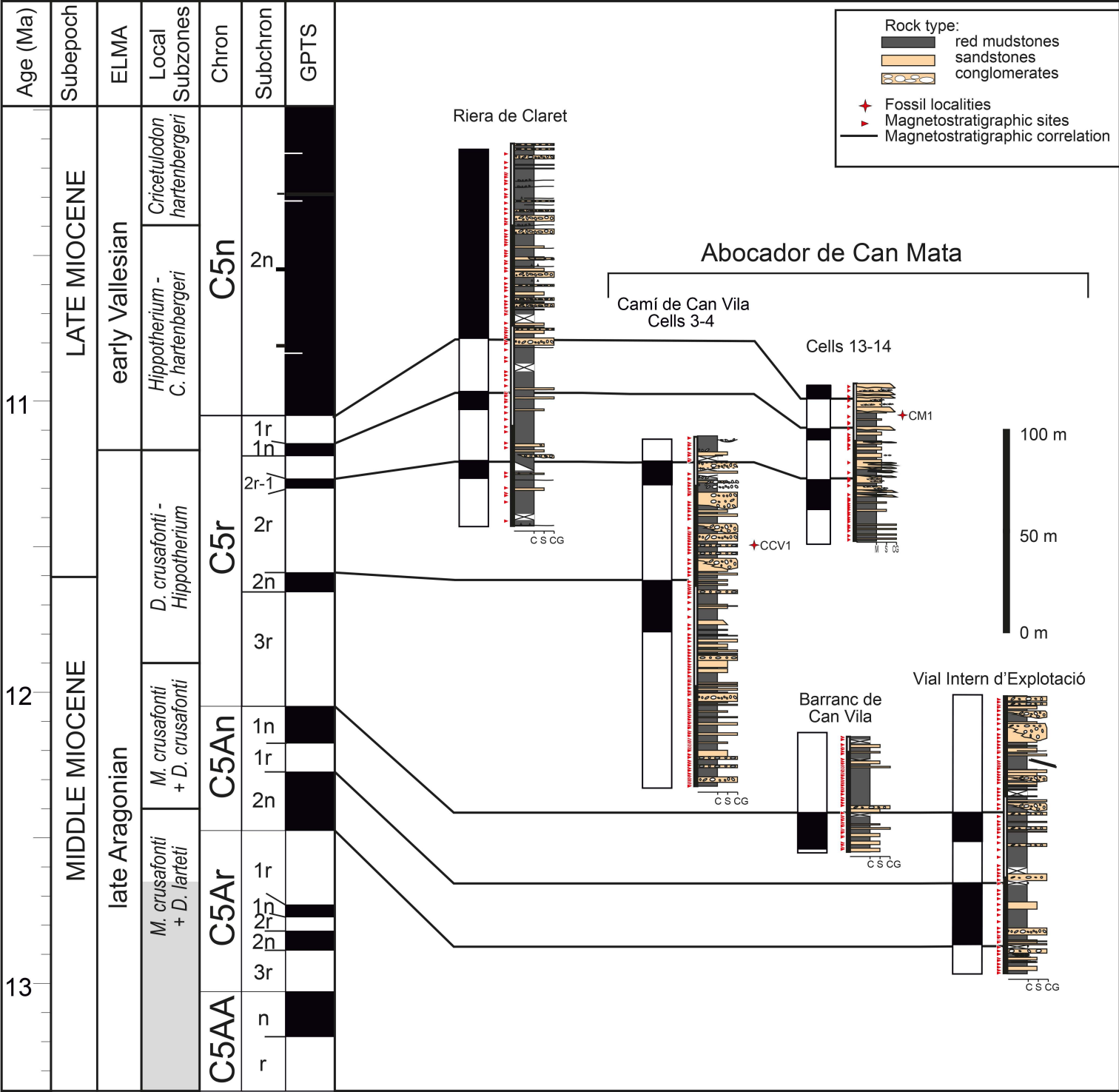


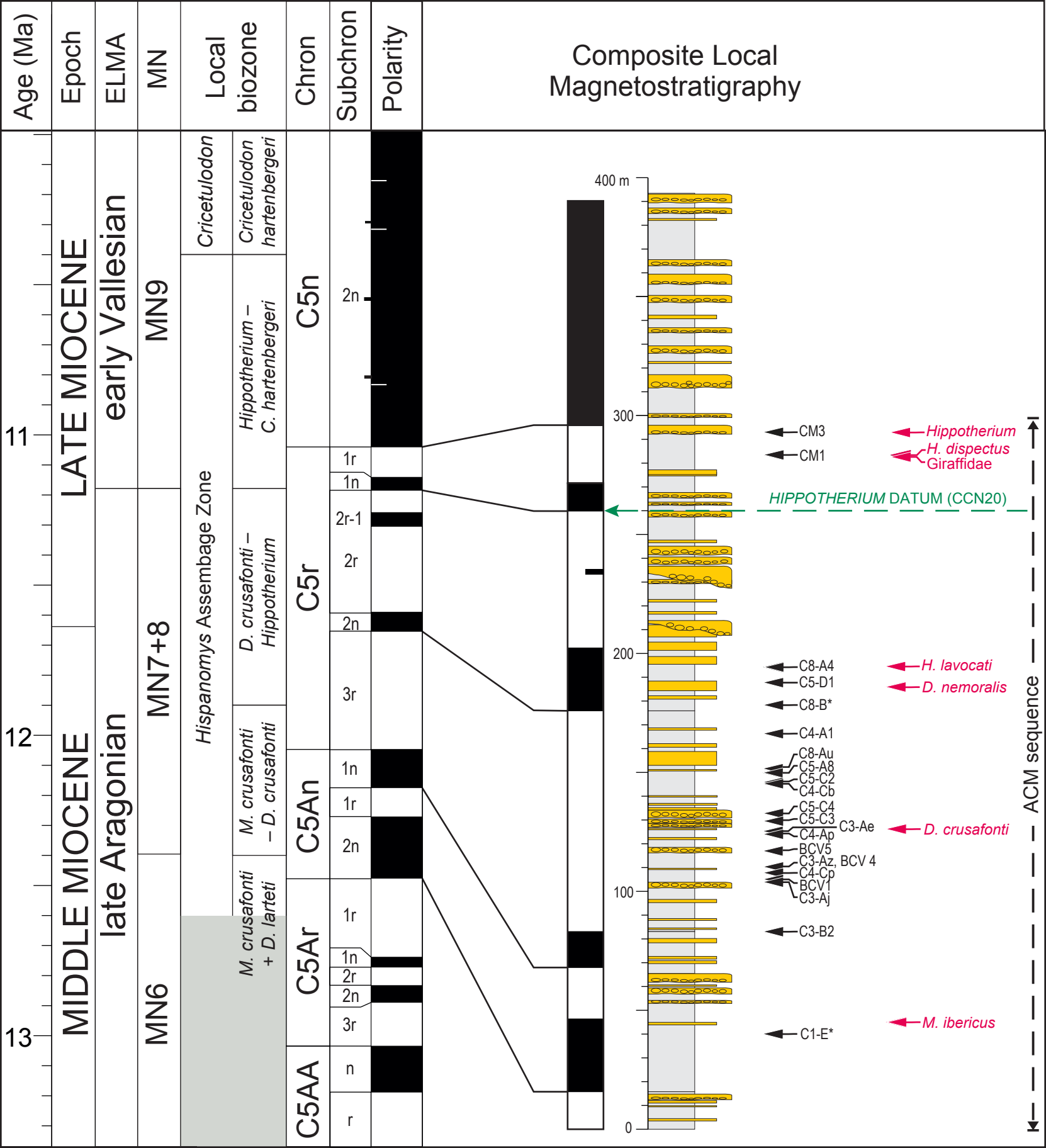
★ CM1
 ★ ACM/BCV1
 ★ Can Vila
 ★ Can Mata de la Garriga

 ACM
 ECM









Supplementary Online Material (SOM):

A revised (earliest Vallesian) age for the hominoid-bearing locality of Can Mata 1 based on new magnetostratigraphic and biostratigraphic data from Abocador de Can Mata (Vallès-Penedès Basin, NE Iberian Peninsula)

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SOM S1

Details about the Can Mata localities

1.1. Historical details about the discovery and definition of Can Mata localities

The locality of CM1 was discovered in 1941 and initially referred to as “an ossiferous breccia” in the surroundings of the farmhouse of Can Mata de la Garriga (Villalta Comella and Crusafont Pairó, 1941: 206, Pl. XVI Figs. 1 and 2) or as “the breccia from Hostalets de Pierola” (Villalta Comella and Crusafont Pairó, 1943: 172, Pl. XV)—our translations from Spanish. The term “ossiferous breccia” used by the authors implied a rich fossil bone bed rather than a breccia in the geological sense of the term. In turn, Crusafont and Truyols (1954) first mentioned the locality of Can Mata 2 (CM2), from which they reported the giraffid *Palaeotragus* sp., in all probability corresponding to the find of scarce dentognathic and postcranial remains previously reported by Crusafont Pairó (1952: Pls. XVIII Fig. 1 and XIX Figs. 1–5). These remains had been recovered from “the inferiormost layers of the Meotian [i.e., Vallesian] in the region of Hostalets de Pierola, nearby «Can Mata» and, thus, almost on the contact with Vindobonian [i.e., Aragonian] levels, [...] in a small breccia located almost on the bottom of a ravine” (Crusafont Pairó, 1952: 71, our translation from the original in Spanish). They were found together with molars of the bovid *Protragocerus chantrei* and considered Vallesian based on the find of *Hippotherium* molars in the same area (Crusafont Pairó, 1952). Finally, the locality of Can Mata 3 (CM3) was first mentioned by Alba et al. (2006), corresponding to an isolated find of *Hippotherium* that thus far represents the oldest hipparionin find of known provenance from the area of Can Mata. Indeed, among the collections of the Institut Català de Paleontologia Miquel Crusafont in Sabadell there are multiple dental or postcranial remains of *Hippotherium* labeled as coming from Vallesian levels of els Hostalets de Pierola, including a few from nearby Can Mata or its surroundings. However, they mostly correspond to finds made by Crusafont and Villalta, whose exact geographic and stratigraphic provenance is unknown. In contrast, according to Alba et al. (2006) CM3 would be located ~200 m in NE direction from CM1, based on a personal communication of Jordi Agustí to J.G. Consulted again by the same author in December 2021, Agustí confirmed that the *Hippotherium* specimen was found during a survey undertaken by him with Josep Gibert in 1976.

1.2. Considerations about the faunal list from Can Mata 1

The published faunal list of large mammals CM1 (Crusafont-Pairó and Golpe-Posse, 1973; Golpe-Posse, 1974; Agustí et al., 1985) needs to be updated and revised, to confirm the

identity of multiple taxa that is confusing according to published sources. In particular, from a biostratigraphic viewpoint it would be interesting to ascertain whether CM1 records typically Aragonian taxa instead of their respective Vallesian counterparts. The former includes the gomphothere *Gomphotherium*, the deinother *Deinotherium levius*, the chalicother *Anisodon grande*, or the cervid *Euprox* aff. *furcatus* (Golpe-Posse, 1974; Agustí et al., 1985; Azanza and Menéndez, 1990)—compared with *Tetralophodon*, *Deinotherium giganteum*, *Chalicotherium goldfussi*, or *Euprox dicranoceros*, which are typically Vallesian (Azanza and Menéndez, 1990; Anquetin et al., 2007; Mazo and van der Made, 2012; Pickford and Pourabrishami, 2013). Nevertheless, despite being mostly Vallesian in age, both *Tetralophodon* (Mazo and Van der Made, 2012) and *Chalicotherium* (Anquetin et al., 2007), are first recorded within MN7+8. In turn, both *Dei. levius* (Konidaris and Koufos, 2019; Alba et al., 2020) and *E. furcatus* (Azanza and Menéndez, 1990; Vislobokova, 2007) apparently persisted since the Aragonian until the earliest Vallesian. Therefore, the usefulness of all these taxa to date localities around the Aragonian/Vallesian boundary is restricted. In contrast, it would probably be more relevant the identification of the large non-listriodontine suid from CM1, customarily attributed to *Propotamochoerus palaeochoerus* (Villalta Comella and Crusafont Pairó, 1943; Golpe-Posse, 1971, 1972, 1974; Crusafont-Pairó and Golpe-Posse, 1973; Agustí et al., 1985), which is currently considered to be conclusively recorded only from the early Vallesian (MN9) across Europe (Pickford, 2016a). Revising the faunal list of CM1 is outside the scope of this paper and will not be an easy task because Crusafont and coauthors seldom provided collection numbers and many of the remains from CM1 housed in the collections of the Institut Català de Paleontologia Miquel Crusafont (ICP) are not labeled as such, but merely referred to Hostalets or Hostalets Inferior.

SOM S2

Descriptions and comparisons of equid and giraffid material from Can Mata

2.1. *Hipparionins*

The best preserved M₃ (IPS124307; SOM Fig. S5a–c) shows a well-developed pli caballinid on the labial side of the hypoconid—a primitive feature that can be found among the Pannonian C *Hippotherium* from Atzeldorf, Gaiselberg, and Mariathal in the Vienna Basin, Austria (Bernor et al., 2017, 2021). Bernor et al. (2017) provided a detailed summary of the upper and lower dental morphology of the Pannonian C hipparionins from the Vienna Basin, identifying the presence of pli caballinids and ectostylids. However, these occlusal and buccal structures of the lower cheek teeth are not always present (Bernor et al., 2017), as their

occurrence may be related to dental wear stage and preservation of the specimens (e.g., pli caballinids may be obliterated in late wear stages and ectostylids may be short, welded to the buccal enamel wall, and buried within the cementum). Although the sample does not include maxillary cheek teeth, the presence of a well-developed pli caballinid in IPS124307 hints at similarities between the Austrian and Iberian samples.

2.2. *Giraffids*

The P² and P³ from CM2 assigned by Crusafont-Pairó (1952: Pl. XVIII Fig. 1) to *Palaeotragus* sp., like the P³ and P⁴ from les Martines assigned by the same author to a different but unnamed species of the same genus (Crusafont-Pairó, 1952: Pl. XVIII Figs. 2 and 3), are buccolingually extended (not only the P⁴). In this regard, the Vallès-Penedès remains differ from those of species such as *Palaeotragus primaevus* that have more or less elongated P² and P³ as well as rounded/subtriangular P⁴. Instead, they more closely resemble, among others, *Giraffokeryx punjabiensis*, *Palaeotragus germaini*, *Decennatherium* spp., and the extant *Giraffa* and *Okapia* (see, e.g., Colbert, 1933; Churcher, 1978; Ríos et al., 2017), which have derived rounded upper premolars that are buccolingually extended. The M³s from ACM (SOM Fig. S6a–f), in turn, are characterized by well-developed cristae, column-like and very rectilinear styles and buccal structures, and presence of protoconal folds. The upper molars from les Martines described by Crusafont-Pairó (1952: Pl. XVIII Figs. 4 and 5) also closely resemble those from ACM in the possession of relatively high and internally flattish cristae, thin and column-like buccal structures, and a strong protoconal fold. A minor difference between the ACM and les Martines upper molars consists in the alignment of the cusps in the buccal wall, which is totally straight in ACM and slightly angled in the latter.

The lower molars from Ecoparc de Can Mata (ECM; SOM Fig. S6m–r) also have extended cristids, higher than those of *Injanatherium* (e.g., Morales et al., 1987), among others. Also, the cusps in the ECM material are flatter and the third lobe of the M₃ is relatively larger than in *Injanatherium*. The M₃ displays a protruding metastylid which projects distally above the pre-entocristid, as in *Giraffokeryx punjabiensis*. The P₄ (SOM Fig. S6j–l) is similar to that of *Palaeotragus*: rectangular, short, and with a complex occlusal surface that has all cusps united by their cristids. As in these forms, the P₄ has a large distal region, as noted by Hamilton (1978). However, the posterior wing (distolingual conid and distolabial stylid) is short, being covered lingually by an extended distolingual cristid that reaches the distolingual corner of the crown. In contrast, both *Giraffokeryx* and *Palaeotragus* have a very large posterior wing that occupies the entire distal lobe. Furthermore, contrary to

Palaeotragus and many other giraffids, in the ECM giraffid the distal lobe is not particularly separated from the rest of the crown by a marked constriction, being more similar to the standard ruminant morphology. On the other hand, *Decennatherium*, a basal member of the samotheres-sivathere clade recorded from the Vallesian of the Iberian Peninsula (Ríos et al., 2016, 2017), has flatter cuspids in the lower molars and more elongated P₄ with clear posterior lobe separation, although the metastylid projection is very similar to the form recorded at ECM.

The transversal section of the metacarpal III–IV from Can Mata described by Crusafont-Pairó (1952: Pl. XIX Fig. 1) shows a medium deep palmar trough (as described by Ríos et al., 2017: character 97, state 0) very similar to that of the early Vallesian *Decennatherium rex*—a character state that can be considered basal for giraffids (Ríos et al., 2017). The (derived) deep condition of the palmar trough is diagnostic of the clade including *Birgerbohlina* + *Bramatherium* (Ríos et al., 2017) and thus the Can Mata form can be excluded from that lineage of very derived giraffids. However, its relationship with more basal samotheres-sivathere (as pointed out by Crusafont-Pairó, 1952) should be subjected to further scrutiny in the future. Similarly, the proximal phalanx of the Can Mata giraffid (Crusafont-Pairó, 1952: Pl. XIX Fig. 3) is somewhat primitive—i.e., elongated and more similar to a regular pecoran phalanx—compared to that of basal members of the samotheres-sivathere clade such as *Dec. rex*, which have proximally broadened and overly blunter proximal phalanges. Also, the proximal surface of the metacarpal III–IV shows a more triangular and laterally elongated unciform facet than that of *Dec. rex*. However, as in the latter, the Can Mata giraffid displays a totally enclosed nonarticular fossette (or synovial fossa) between the facets for the unciform and the magnotrapezoid. Finally, the astragalus (Crusafont-Pairó, 1952: Pl. XIX Fig. 5) has equally deep proximal and distal trochleae, being different from the stouter astragalus of giraffids such as *Decennatherium*, in which the distal trochlea is shorter than the proximal one, and the astragalus is overall short and square.

SOM S3

Biostratigraphic remarks based on rodents

3.1. Similarities between latest Aragonian and earliest Vallesian rodent assemblages

According to the most recent local rodent biozonation of the Vallès-Penedès Basin (Casanovas-Vilar et al., 2016a), the latest Aragonian corresponds to the *Democricetodon crusafonti*–*Hippotherium* interval subzone (from the first local occurrence [FLO] of *Dem. crusafonti* to that of *Hippotherium*), whereas the earliest Vallesian is represented by the *Hippotherium*–*Cricetulodon hartenbergeri* interval subzone (from the FLO of *Hippotherium* to that of *C. hartenbergeri*). Both subzones belong to the *Hispanomys* assemblage zone (Casanovas-Vilar et al., 2016a), and they are exceptionally based on the equid *Hippotherium* because the respective micromammal assemblages have been considered indistinguishable in taxonomic composition.

The presence of *Megacricetodon ibericus* is not informative regarding the placement of the Aragonian/Vallesian boundary in the Vallès-Penedès Basin, because it is not only recorded from the late Aragonian (Casanovas-Vilar et al., 2016a) but also from the earliest Vallesian sites of Creu de Conill 20 and 22 (Casanovas-Vilar et al., 2006, 2016a), CM1 and Hostalets Superior (Agustí and Gibert, 1982; Agustí et al., 1985; Casanovas-Vilar et al., 2016a), and tentatively also Can Missert and ECM (Agustí et al., 2005; Alba et al., 2012; Casanovas-Vilar et al., 2016a). Similarly, despite its later appearance in the Calatayud-Montalbán Basin, *M. ibericus* s.s. is already recorded there from unambiguously late Aragonian (zone G3) sites (Álvarez Sierra et al., 2003; Van Dam et al., 2014; Oliver Pérez, 2015), persisting until the early Vallesian site of Nombrevilla 1 at ~10.8 Ma (Daams and Freudenthal, 1988; Álvarez Sierra et al., 2003; Van der Meulen et al., 2003; Oliver Pérez, 2015).

3.2. The *Cricetulodon* event

The cricetid *Cricetulodon hartenbergeri* is first recorded at ~10.6 Ma (zone I) in the Calatayud-Montalbán Basin (Van Dam et al., 2014) and at ~10.3 Ma in the Vallès-Penedès (Autopista de Rubí-Terresa 8 locality, with an interpolated age of 10.34 Ma; Casanovas-Vilar et al., 2016a). An isolated M¹ from Hostalets Inferior, originally ascribed to *Democricetodon nemoralis* by Agustí (1981) but no longer present in the ICP collections, was subsequently attributed to *Cricetulodon hartenbergeri* by Agustí and Gibert (1982: Pl. II Fig. 5; see also Agustí et al., 1985; Alba et al., 2006). However, this specimen is slightly larger than the M¹s of *C. hartenbergeri* and we concur with Agustí (1981) that, despite the

presence of a double protolophule (very common in *C. hartenbergeri* and much rarer in *Dem. nemoralis*), it better fits with the latter species, which is first recorded in the ACM composite section at ACM/C6-C2, with an interpolated age of 11.60 Ma. On the other hand, no *Cricetulodon* material has been recovered from Can Mata—not even from well-sampled ECM localities (ECM/VCE-B2 and ECM/VCE-C1; Casanovas-Vilar et al., 2016a), which have an estimated age of ~10.8 and ~10.7 Ma, respectively. This supports the hypothesis that a major turnover of small mammal faunas, characterized by the appearance of *C. hartenbergeri* and the disappearance of *Dem. crusafonti* and *M. ibericus*, did not occur until much later than the Aragonian/Vallesian boundary.

3.3. The *Hispanomys* succession across the Aragonian/Vallesian boundary

Hispanomys species offer some prospects to distinguish between the latest Aragonian and the earliest Vallesian based on micromammals. Casanovas-Vilar et al. (2016a: 209) already noted that “*Hispanomys* species could also be of great use in refining the diagnoses for some zones and their boundaries”, although they did not specifically note its potential utility for characterizing the Aragonian/Vallesian transition. The succession of *Hispanomys* species in the Vallès-Penedès Basin is different from that in other Iberian basins. In the Calatayud-Montalbán Basin, *Hispanomys lavocati* co-occurs with *M. ibericus* and *Dem. crusafonti* in MN7+8 (zone G3) sites dating to ~11.9–11.8 Ma (López-Guerrero et al., 2014, 2019; Van Dam et al., 2014), whereas localities dating to zone H (Nombrevilla 9, 10, and 1) are characterized by the appearance of *Hispanomys nombrevillae* (López-Guerrero et al., 2014, 2019; Van Dam et al., 2014; García-Paredes et al., 2016). The latter species is also recorded in early Vallesian sites from the Duero Basin before the appearance of *Hispanomys aragonensis* (see Castillo et al., 2018). In summary, in both the Calatayud-Montalbán and Duero basins, *His. lavocati* is restricted to zone G3 (latest Aragonian), *His. nombrevillae* to zone H (around the Aragonian/Vallesian transition), and *His. aragonensis* to zone I (early Vallesian; Castillo et al., 2018; López-Guerrero et al., 2019).

In the Vallès-Penedès Basin, *His. lavocati* co-occurs with *M. ibericus* and *Dem. crusafonti* since it is first recorded from ACM at ~11.6 Ma, as well as in other MN7+8 sites such as Trinxera del Ferrocarril in Sant Quirze (Casanovas-Vilar et al., 2016a), suggesting that the latter is similar in age to the latest Aragonian localities from ACM. However, unlike in the Calatayud-Montalbán and Duero basins, in the Vallès-Penedès Basin *His. lavocati* persists into the earliest Vallesian and *His. nombrevillae* is not recorded. In contrast, two different species of *Hispanomys* (*Hispanomys dispectus* and *Hispanomys daamsi*) are recorded during

the earliest Vallesian of the Vallès-Penedès Basin. Indeed, in NE Iberia *His. dispectus* is also recorded, together with *M. ibericus*, from various localities of the Empordà Basin (La Bisbal 2, Can Colomer 1, and Abocador de Vacamorta 4; Gibert et al., 1979, 1980; Agustí Ballester, 1980; Llenas et al., 2002; Casanovas-Vilar et al., 2010) that are correlated to the earliest Vallesian (Casanovas-Vilar et al., 2010; Tosal et al., 2022). In the Vallès-Penedès Basin, this species was originally reported from Hostalets Inferior, Can Feliu, Castell de Barberà, and Hostalets Superior (Agustí Ballester, 1980), and subsequently from CM1, Can Valls in Masquefa, and Can Missert (Agustí et al., 1985), Autopista de Rubí-Terrassa km 11 and Can Coromines 2 (Agustí et al., 1997), and Creu de Conill 20 and 22 (Agustí et al., 1997; Casanovas-Vilar et al., 2006). The material from Can Missert was later used by Agustí et al. (2005) to erect a *His. daamsi*, to which the material from Castell de Barberà was also reassigned (Casanovas-Vilar et al., 2016a; Alba et al., 2019). Despite earlier preliminary claims (Alba et al., 2006, 2011; Casanovas-Vilar et al., 2008), ACM localities record *Hispanomys* cf. *aguirrei*, *Hispanomys decedens*, and *His. lavocati* (depending on the locality) but not *His. daamsi* (see Casanovas-Vilar et al., 2016a). Similarly, we have been unable to confirm the citation of *His. dispectus* from the later Vallesian locality Can Coromines 2, which is correlated to C5n.1r (Agustí et al., 1997), indicating an age of ~10.0–9.9 Ma, as the only cricetid specimen from this locality found in the ICP collections (the M₃ IPS18100) cannot be assigned to genus (authors' observation).

All the localities conclusively recording either *His. dispectus* (Hostalets Inferior and Superior, CM1, Can Feliu, Can Valls, Creu de Conill 20 and 22, and Autopista de Rubí-Terrassa km 11) or *His. daamsi* (Castell de Barberà and Can Missert) are currently considered earliest Vallesian in age (Casanovas-Vilar et al., 2016b)—including Autopista de Rubí-Terrassa km 11, which is correlated to C5n.2n (Agustí et al., 1997) with an interpolated age of 10.98 Ma (Casanovas-Vilar et al., 2016a)—with the sole exception of Hostalets Inferior. However, the fact that the micromammal assemblage from Hostalets Inferior includes material from CM1 (presumably in most instances not labeled as such), which is here conclusively dated to the earliest Vallesian, makes it uncertain the presence of either species in the late Aragonian levels of this area (or elsewhere). This is further supported by the absence of both species from ACM, suggesting that they might be characteristic of the early Vallesian or, at most, the Aragonian/Vallesian transition more broadly. Later early Vallesian localities from ECM record *His. aragonensis* instead (Alba et al., 2012; Casanovas-Vilar et al., 2016a), which is first conclusively recorded at ECM/VCE-B2 at ~10.8 Ma (updated from Casanovas-Vilar et al., 2016a), i.e., slightly earlier than in the Calatayud-

Montalbán Basin, where it is first recorded from Nombrevilla 13 at ~10.7 Ma (López-Guerrero et al., 2014; Van Dam et al., 2014). This supports that *His. dispectus* may be restricted to early MN9, thus potentially being—together with *H. daamsi*—a biochronological marker of the Vallesian in its type area. Further refinements will probably be possible once the large sample of *Hispanomys* remains from ACM and ECM (including >4000 specimens) is studied in detail. The study of the micromammal remains recently recovered from the uppermost sectors of the ACM composite section (ACM/C12, ACM/C13, ACM/C14 and ACM/PTA)—jointly extending from 11.4 to 11.1 Ma (from C5r.2r to C5r.1r)—will hopefully show, in the future, whether *His. dispectus* makes its first appearance before the Aragonian/Vallesian boundary or coinciding with the earliest Vallesian.

SOM S4

Biochronological remarks about large mammals

4.1. Age of the earliest hipparionins from China

In the earliest Vallesian, China records *Hippotherium weihoense*, which is characteristic of the Guonigou Fauna of the Linxia Basin and co-occurs with '*Hipparion*' *dongxiangense* (~11.2–11.1 Ma; Deng et al., 2004, 2013; Sun et al., 2022). The Guonigou Fauna is correlated to the early Bahean Asian Land Mammal Age, correlative with the early Vallesian (Deng et al., 2013). Fang et al. (2016) reviewed the tectonosedimentary evolution of the Linxia Basin based on new paleomagnetic analyses, suggesting an age of ~11.5 Ma for the first occurrence of hipparionins in the Guonigou Fauna, thus being slightly older than the record of *Hippotherium* in the Vienna and Vallès-Penedès basins (see also Sun et al., 2022). However, the Guonigou hipparionin record lacks documentation of specimens as well as specific stratigraphic occurrence and magnetostratigraphic correlation within the Linxia Basin, and as such it needs further documentation to be adequately vetted. Indeed, the magnetostratigraphy provided by Fang et al. (2016: Fig. 7) is missing many relevant subchrons, the Chinese *Hippotherium* datum being correlated to a reverse polarity magnetozone interpreted as C5r.1r–C5r.2r (Fang et al., 2016). These subchrons span from 11.59 to 11.15 Ma and include C5r.1n, when the earliest *Hippotherium* is recorded in the Vallès-Penedès Basin. The inability to detect these polarity reversals questions Fang et al.'s (2016) unambiguous magnetostratigraphic correlation of the Chinese *Hipparion* datum with C5r.2r. Alternative correlations cannot be thus entirely discounted, including C5r.1r, which could imply a younger age roughly coeval with the European *Hippotherium* datum.

4.2. Hipparionins from the Duero Basin

In the Duero Basin, hipparionins are recorded together with *Decennatherium* in El Lugarejo, Ávila (Crusafont Pairó et al., 1968; Morales et al., 1981) and in the much better-known site of Los Valles de Fuentidueña, Segovia, which also records *Machairodus* (Crusafont Pairó, 1952; Crusafont-Pairó and Ginsburg, 1973; Alberdi, 1981; Alberdi et al., 1981; Ginsburg et al., 1981; Morales and Soria, 1981; Ríos et al., 2016; Fernández-Monescillo et al., 2019). Both sites are considered similar in age, as supported by similarities in large mammal composition (the giraffid, hipparionin, and *Hispanomeryx* species; Crusafont Pairó et al., 1968; Alberdi, 1974, 1981; Morales et al., 1981; Sánchez and Morales, 2006), being sometimes correlated to zone H (Sánchez and Morales, 2006). However, there are no micromammals from El Lugarejo, while the small mammals from Los Valles de

Fuentidueña (Agustí Ballester, 1978; Alberdi et al., 1981; Sesé Benito and López Martínez, 1981) do not provide a precise dating. The *Hispanomys* species from the latter site, in particular, shows similarities with both *His. aragonensis* and *His. nombrevillae*, whereas a species assignment is neither possible for a member the *Megacricetodon minor*–*Megacricetodon minutus* lineage (Sesé Benito and López Martínez, 1981; Castillo et al., 2018)—the latter species considered here a senior synonym of *Megacricetodon debruijni*. Only the lack of *Cricetulodon* indicates that Los Valles de Fuentidueña should be correlated to zone H (early MN9) instead of zone I (see Van Dam et al., 2014).

The hipparionin recorded in these localities from the Duero Basin was originally considered by Crusafont Pairó et al. (1968) an archaic form probably different from *Hippotherium catalaunicum*. It was subsequently described as *Hippotherium primigenium melendezi* by Alberdi (1974, 1981), who also considered *Hip. catalaunicum* as a subspecies of *Hip. primigenium*. The hipparionin from the Duero Basin is smaller and more gracile than *Hip. catalaunicum*, but based on the cranial evidence from El Lugarejo, Bernor et al. (1980, 1990) first referred it to ‘Group 1’ (i.e., *Hippotherium*). Nevertheless, differences in the morphology of the preorbital fossa noted by Bernor and Hussain (1985), who considered it a distinct species, eventually led to its transferral to *Hipparion* s.s. (Bernor et al., 1996). In contrast, other authors have considered this species to be conspecific with the Turolian species *Hipparion concudense* (Forstén, 1982; Pesquero and Alberdi, 2012), which Zouhri and Bensalmia (2005) referred to *Cremohipparion*. This is not only at odds with the Vallesian age of these sites, but also with the morphology displayed by the cranium from El Lugarejo, which rather resembles those of *Hipparion* s.s. (group 3 of Woodburne and Bernor, 1980), first recorded at MN10 in Iran (Bernor et al., 2017). Detailed comparisons would be required to determine the taxonomic identify of the hipparionin species from El Lugarejo and Los Valles de Fuentidueña, but differences relative to *Hippotherium* from both the Vallès-Penedès and Calatayud-Montalbán basins suggest that an earliest Vallesian age is unlikely.

4.3. Earliest Vallesian giraffids from the Vallès-Penedès Basin

Crusafont Pairó (1952) noted dental similarities between the CM2 giraffid material and the somewhat smaller *Palaeotragus quadricornis* (currently in *Schansitherium*; Solounias, 2007), hypothesizing that the former might belong to a new species, and refrained from including it in *Decennatherium pachecoi*—which he described based on the Vallesian material from other Spanish localities (see also Ríos et al., 2016). Ríos et al. (2016, 2017) considered that this genus is only recorded conclusively from MN9 and MN10 localities of Spain (Catalatayud-

Montalbán, Duero, and Madrid basins; Ríos et al., 2016, 2017). However, more recently it has also been reported from the Siwaliks in Pakistan (Ríos et al., 2019). The latter material is Late Miocene in age (10.1–8.2 Ma, MN10–MN11) and more derived than *Dec. pachecoi*, except for some ossicone fragments from locality Y503 (13.5 Ma) that have been assigned with doubts to cf. *Decennatherium* (Ríos et al., 2019). The latter occurrence supports the view that *Dec. pachecoi* and probably ‘*Palaeotragus*’ sp. from the Vallès-Penedès Basin are local endemisms that originated from a dispersal event from the east.

The earliest well-dated record of giraffids in the Vallès-Penedès Basin corresponds to the earliest Vallesian site of Castell de Barberà (~11.2 Ma), followed by the recent find reported here from ACM/C13/C14-A (~11.1 Ma, probably roughly coeval to the remains from CM2 described by Crusafont-Pairó, 1952) and ECM/VCE-Bb (~10.9 Ma). Besides these specimens, giraffid remains traditionally assigned to *Palaeotragus* have also been cited from the Vallès Sector of the Vallès-Penedès Basin—sites of Can Feliu, Can Missert, Can Poncic 2, Castell de Barberà, les Martines, Poble Nou de Sant Quirze, Sant Miquel de Toudell, and Santiga (Crusafont Pairó, 1952, 1953, 1965; Crusafont and Truyols, 1954; Crusafont-Pairó and Golpe, 1972; Crusafont-Pairó and Golpe-Posse, 1972; Golpe-Posse, 1974; Crusafont-Pairó and Golpe-Posse, 1974a; Crusafont Pairó, 1975, 1982; Agustí et al., 1985, 2005). Some of these sites (Sant Miquel de Toudell and Santiga) are clearly Vallesian (e.g., Casanovas-Vilar et al., 2016a, 2016b), but the age of most of the remaining ones has proven controversial due to the absence of *Hippotherium* coupled with the lack of magnetostratigraphic data and clear differences between the latest Aragonian and earliest Vallesian rodent assemblages (but see SOM S2.3). Such controversies are best exemplified by Castell de Barberà, whose correlation to the latest Aragonian was maintained despite the find of a hipparionin fragment (Crusafont-Pairó and Golpe-Posse, 1974b; see also Rotgers and Alba, 2011), until the find of additional *Hippotherium* remains and magnetostratigraphic data more unambiguously supported a Vallesian age (Alba et al., 2019).

Among the aforementioned localities of ambiguous age with ‘*Paleotragus*’ sp., only the Vallesian dating of les Martines in Terrassa, based on the regional context (Crusafont Pairó, 1952, 1953; Crusafont and Truyols, 1954), has remained uncontroversial. Indeed, the location of the clay quarry where the giraffid remains were recovered is unknown, but the area of les Martines is near that of Creu de Conill, so that an early Vallesian age is plausible. Can Feliu (Crusafont and Truyols, 1954), Can Poncic 2 (Crusafont and Truyols, 1956), and Can Missert (Crusafont-Pairó and Golpe-Posse, 1974a) were originally considered Vallesian as well, whereas Poble Nou de Sant Quirze was first considered (with doubts) pre-Vallesian

(Crusafont and Truyols, 1956), but subsequently dated to the Vallesian by Crusafont Pairó (1962) based on the fact that it was calculated to be 30 m stratigraphically above Trinxera del Ferrocarril (which is likely equivalent in age to the uppermost latest Aragonian portion of the ACM sequence; see SOM 2.3). Following the definition of a new pre-Vallesian biozone characterized by the lack of hipparionins and the presence of giraffids (Crusafont Pairó and Golpe Posse, 1971), most of the aforementioned localities (CM2, Can Poncic 2, Castell de Barberà, Poble Nou de Sant Quirze, and Can Missert) were considered pre-Vallesian (Crusafont-Pairó and Golpe, 1972; Golpe-Posse, 1972, 1974; Crusafont Pairó, 1975; Agustí et al., 1985). Also on this basis, Agustí et al. (1985) considered that CM1 predated CM2, given the lack of giraffids in the former.

The definition of the above-mentioned biozone (Crusafont Pairó and Golpe Posse, 1971) prompted the definitive abandonment of Crusafont Pairó's (1950) original concept of the Vallesian based on the presence of both hipparionins and giraffids, in favor of the view that only the giraffid *Decennatherium* (but not '*Palaeotragus*') would be indicative of the Vallesian (Agustí et al., 1985, 1997, 2001). Agustí and Moyà-Solà (1991: 111) even considered that “no other dispersal event is associated with the appearance of *Hipparion* in Europe, the entry of this equid being an isolated datum”, while Agustí (1999: 103) remarked that “the latest Aragonian and earliest Vallesian faunas in the Vallès-Penedès share the same small and large mammal elements”. These assertions are surprising, given that Agustí et al. (1997) considered the dispersal of *Hippotherium* and *Machairodus* to be synchronous. Notwithstanding the latter, the notion that giraffids dispersed into the Vallès-Penedès Basin before the entry of *Hippotherium* was thus perpetuated by Agustí and colleagues (Agustí and Moyà-Solà, 1990; Agustí et al., 1997, 2005; Agustí, 1999) and further extended to Western Europe as a whole by Agustí et al. (2001), who considered that, unlike *Hippotherium* and the felid *Machairodus*, the giraffid *Palaeotragus* dispersed during MN7+8. More recently, however, an earliest Vallesian age was favored for Can Missert, Can Feliu, Can Poncic 2, and Castell de Barberà (Robles et al., 2011; Casanovas-Vilar et al., 2016a, 2016b; Alba et al., 2019). Only Poble Nou de Sant Quirze and CM2 remained as putative latest Aragonian sites from the Vallès-Penedès Basin where giraffids are recorded. However, the fauna from Poble Nou de Sant Quirze is only poorly known, whereas our results further indicate that a latest Aragonian age is no longer tenable for CM2, as CM1 correlates to the Vallesian despite the lack of *Hippotherium*. Furthermore, our results indicate that CM1 is penecontemporaneous with the oldest well-dated record of giraffids in the area, which in turn slightly postdates the earliest record of *Hippotherium* in the Vallès-Penedès Basin. At ACM, giraffids have yet to

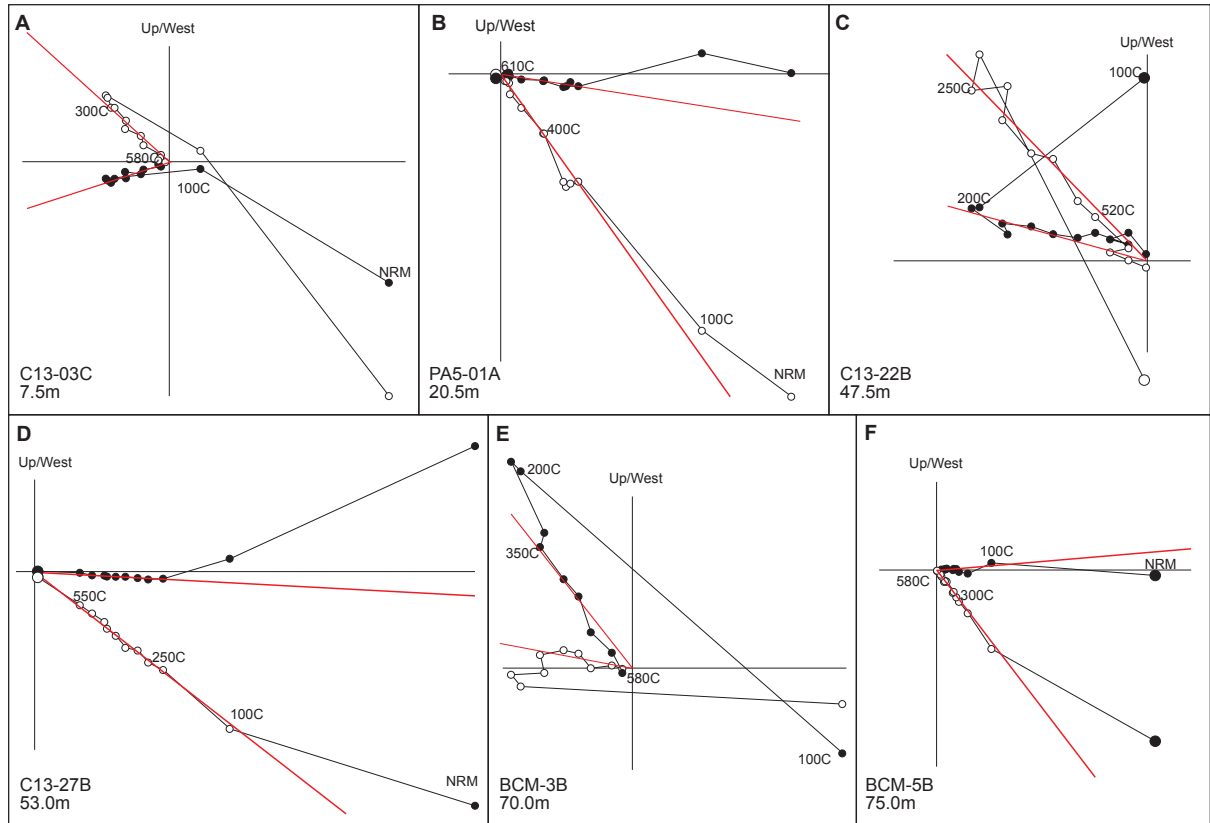
be found in the late Aragonian levels despite the huge sampling effort (Alba et al., 2006, 2017), thus challenging the traditional notion (e.g., Crusafont Pairó and Golpe Posse, 1971; Golpe-Posse, 1974; Crusafont Pairó, 1975; Agustí et al., 1985, 1997, 2001) that giraffids dispersed into the Vallès-Penedès Basin before hipparionin equids. Rather the contrary, our results suggest that giraffids and hipparionins might have dispersed at approximately the same time, as originally considered by Crusafont Pairó (1950).

4.4. Other Vallesian taxa of eastern origin

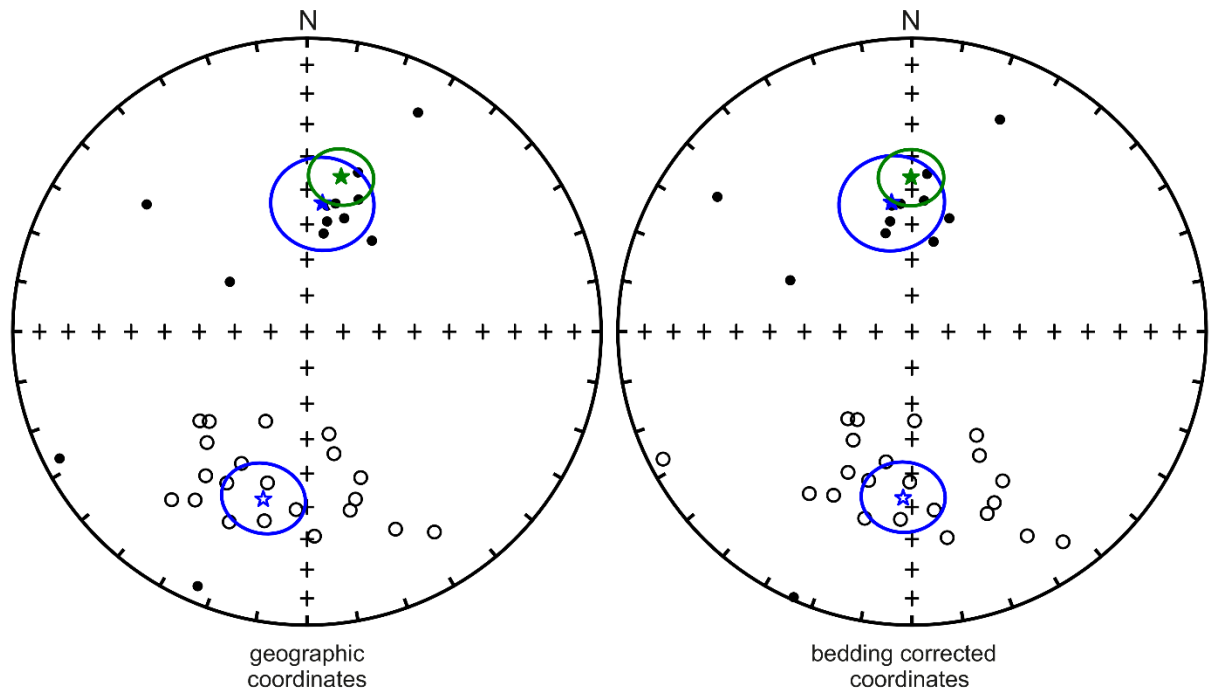
The machairodontine felid *Machairodus aphanistus* is one of the large mammal species that first appears in the Iberian Peninsula and elsewhere in Europe coinciding with the beginning of the Vallesian (Agustí et al., 1997; Agustí, 1999; Agustí and Antón, 2002). The geographic and evolutionary origin of *Machairodus* is somewhat uncertain (Werdelin et al., 2010; Morlo et al., 2020) but has been related to the more plesiomorphic sabertoothed cat ‘*Miomachairodus*’ *pseudailuroides*, which is first recorded in the late Aragonian (MN7+8) of Turkey (Schmidt-Kittler, 1976; de Beaumont, 1978; Werdelin et al., 2010; Morlo et al., 2020). This would support the view that *Machairodus* dispersed into Western Europe from the east at the beginning of the Vallesian (Agustí et al., 1997, 2001), rather than locally evolving from a *Pseudaelurus* ancestor. Another relevant example refers to the suine *P. palaeochoerus*, whose first appearance datum was until recently correlated to the late MN7+8 (Van der Made and Moyà-Solà, 1989; Van der Made et al., 1999; Agustí et al., 2001; Fortelius et al., 2005; Alba et al., 2006). However, this was in part due to the incorrect dating of some sites (e.g., Castell de Barberà, where *P. palaeochoerus* is unambiguously recorded; Pickford, 2016a), and also—as in the Vallès-Penedès sites of Sant Quirze (Van der Made and Moyà-Solà, 1989; Van der Made, 1990a, 1990b) and ACM (Alba et al., 2006, 2011)—owing to the confusion with a large tetraconodontine currently referred to *Parachleuastochoerus valentini* by Pickford (2014, 2016a, 2016b). Although there is disagreement about the taxonomic validity of the latter species (Van der Made, 2020), *P. palaeochoerus* is currently considered to be conclusively recorded only from MN9 localities (Pickford, 2016a). The genus is however already recorded in the Middle Miocene of Asia (Harris and Liu, 2007), from where it would have dispersed into Europe (Agustí, 2015).

In the Vallès-Penedès Basin, the earliest published record of *Machairodus* corresponds to Can Mata (Madurell-Malapeira et al., 2014), but the lack of stratigraphic details precludes confirming whether it predates or postdates the Aragonian/Vallesian boundary. The citation of *Machairodus* from the earliest Vallesian locality of Creu de Conill 20 (Agustí et al., 1997;

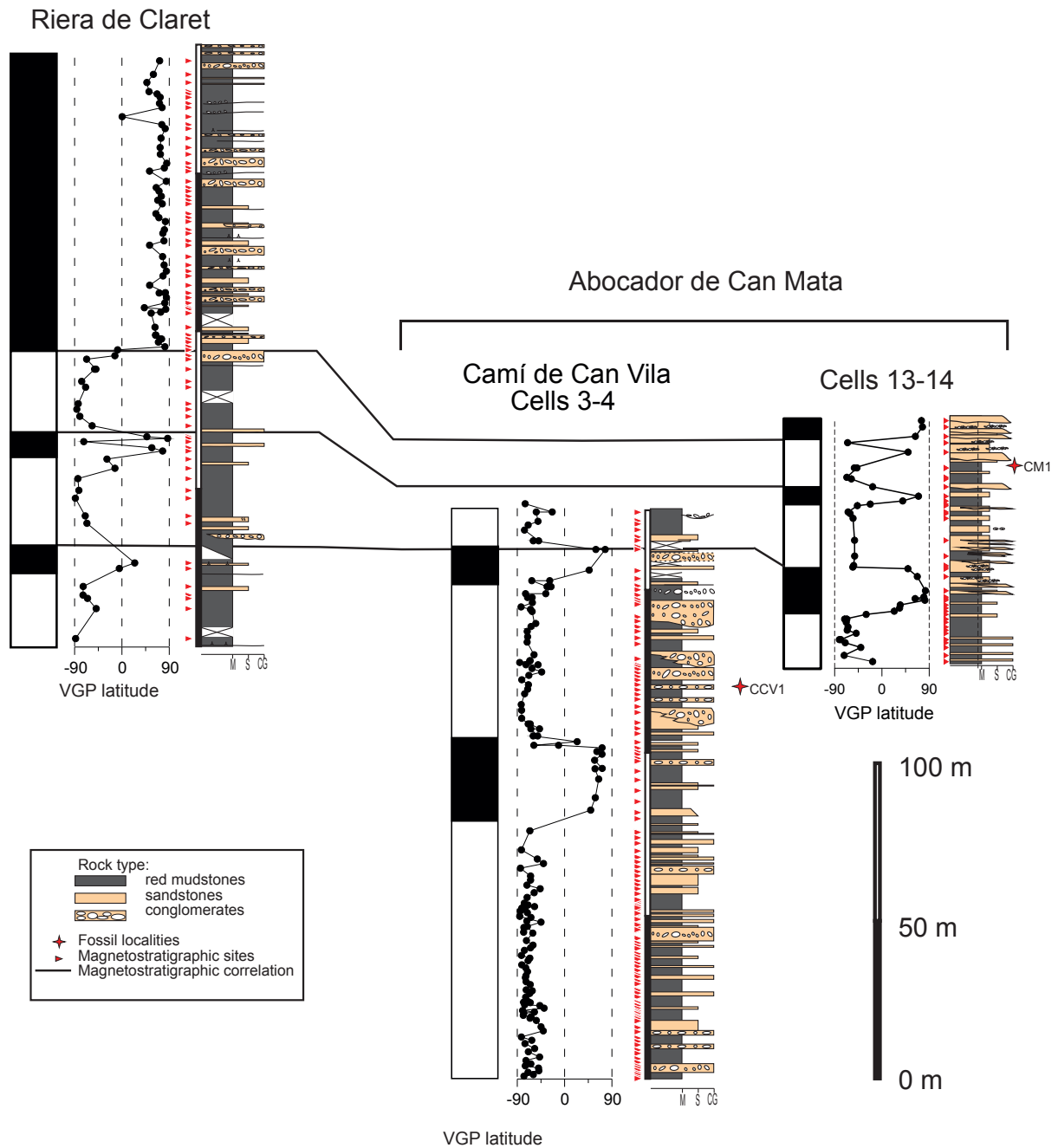
Agustí and Galobart, 1998; Casanovas-Vilar et al., 2006), was dismissed by Madurell-Malapeira et al. (2014). However, recent fieldwork at Creu de Conill 20 (Almécija et al., 2019) has demonstrated the co-occurrence of *Machairodus*, *Propotamochoerus* and *Hippotherium* coinciding with the first appearance datum of the latter taxon in the Vallès-Penedès Basin. Unfortunately, our efforts to locate within the ICP collections the suid remains that led to the citation of *P. palaeochoerus* from CM1 (e.g., Golpe-Posse, 1974; Agustí et al., 1985) have been unsuccessful. Only Golpe-Posse (1971), in her unpublished dissertation, listed catalog numbers for the specimens from Hostalets assigned to this taxon, but she did not note which specimens came from CM1 or even distinguish between Hostalets Inferior and Superior. As a result, it is currently impossible to confirm whether *P. palaeochoerus* is recorded there (as opposed to stratigraphically superior levels customarily included in Hostalets Superior), or whether the previous citations are just another confusion with *Pa. valentini*, as argued by Pickford (2016a, 2016b).



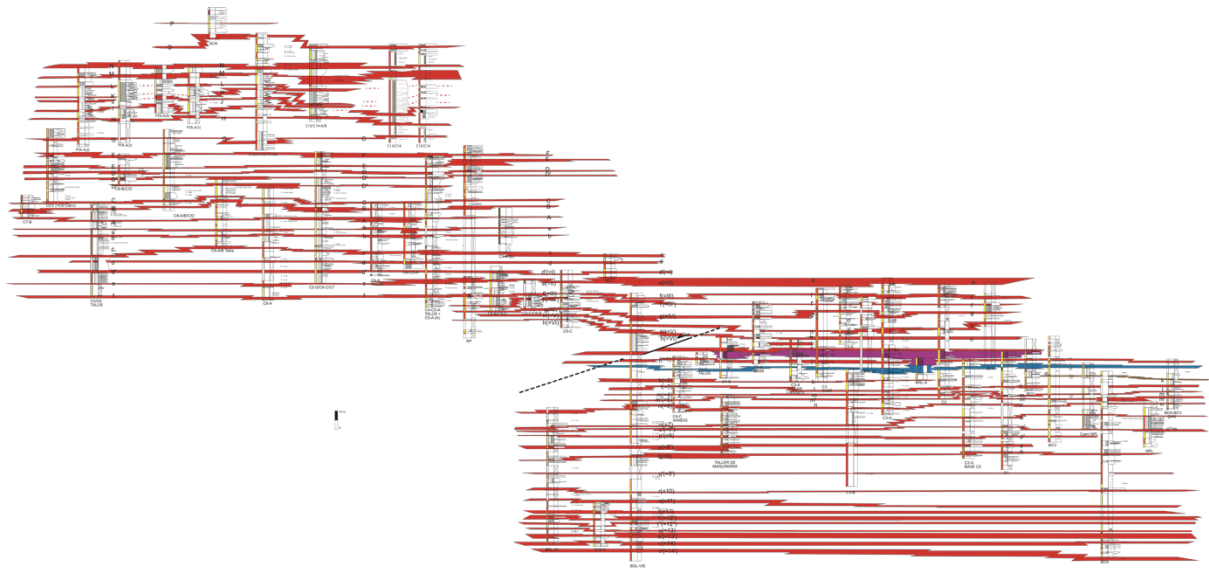
SOM Figure S1. Examples of stepwise thermal demagnetization of samples representative of all recorded magnetozone in the Abocador de Can Mata composite section. Diagrams represent orthogonal projections of vector endpoint demagnetization data; black and white dots represent the projection onto the horizontal and vertical planes, respectively. Red lines represent least-squares fits that define the characteristic paleomagnetic direction. Stratigraphic position in meters is indicated in the lower left corner of each diagram.



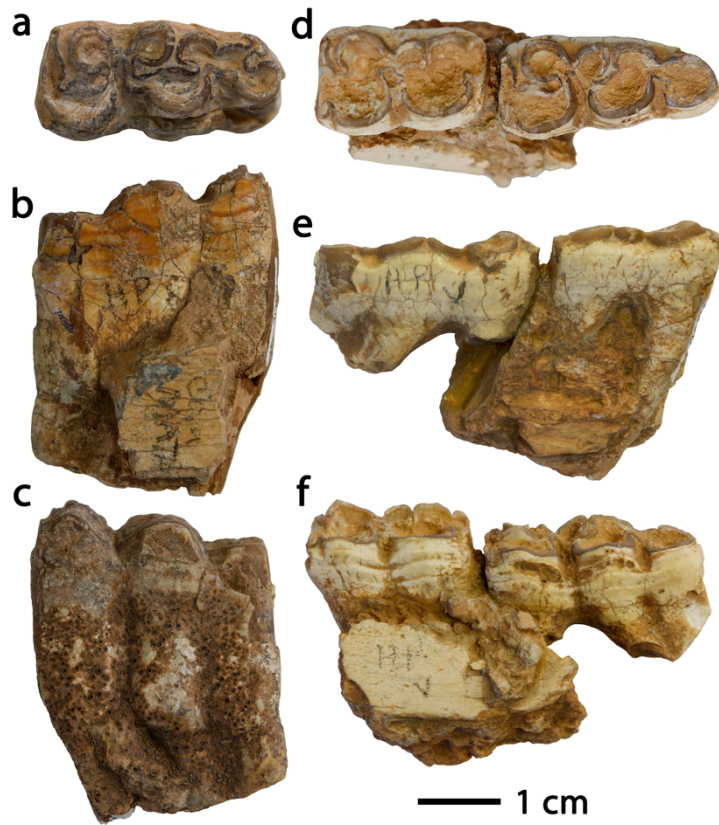
SOM Figure S2. Equal area stereonet projection of paleomagnetic directions and the mean normal and reversed polarity directions in geographic and bedding corrected coordinates (blue stars). Green star represents the mean direction after conversion of all the directions to normal polarity. See also SOM Table S3 for numerical values and Fisher statistics parameters.



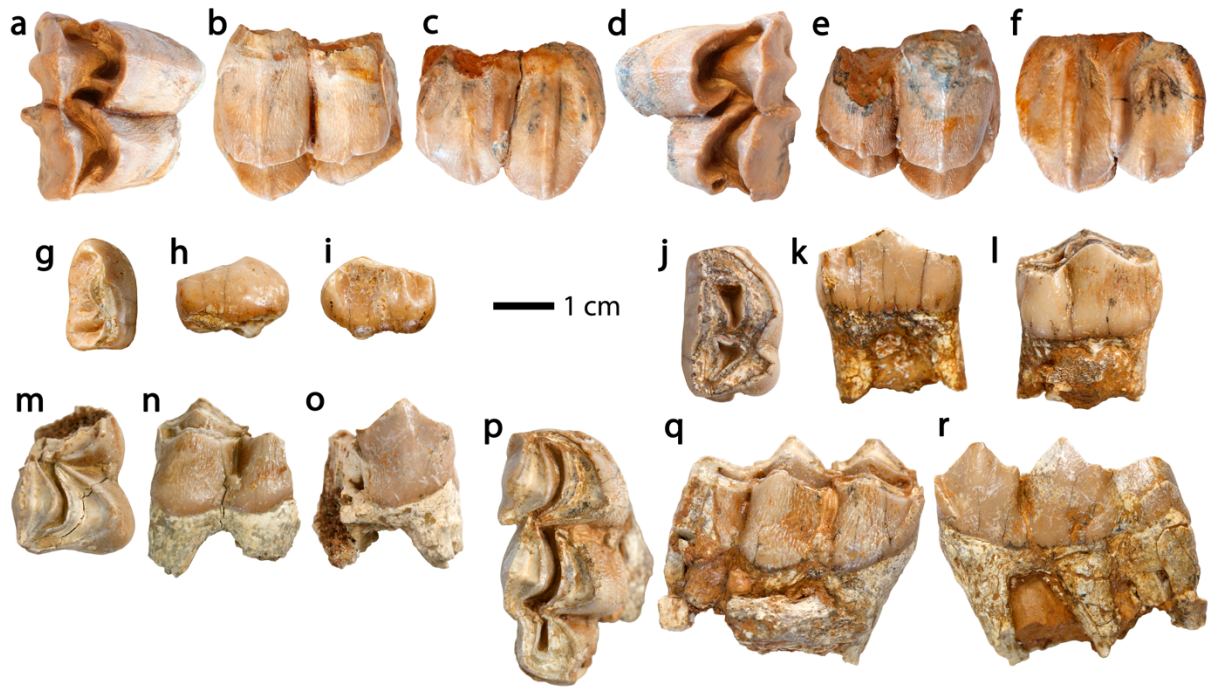
SOM Figure S3. Magnetostigraphic correlation of the new Cells 13–14 section, where Can Mata 1 is located (denoted by a four-pointed red star), with the Camí de Can Vila–Cells 3–4 (Abocador de Can Mata) and the Riera de Claret sections (after Moyà-Solà et al., 2009 and Alba et al., 2017, with some corrections). Sideways triangles denote the stratigraphic position of paleomagnetic samples. Abbreviations: M = mudstones; S = sandstones; CG = conglomerates; VGP = virtual geomagnetic pole.



SOM Figure S4. (The source image is provided separately in much higher resolution as a PDF file). Detailed lithostratigraphic panel showing the correlation between the 51 local sections sampled at Abocador de Can Mata, updated after Alba et al. (2017). The used correlation levels and the defined fossil localities (SOM Table S4) are shown. Correlation levels are denoted by red subhorizontal stripes that cross different sections (except for level α , depicted in blue). Each section is represented by a column depicting the following details (from left to right): stratigraphic thickness (in meters, all to the same scale); approximate coloration (streaked patterns represent mixed colorations); approximate granulometric composition (white = mudstones; light gray = sandstones; darker gray = conglomerates); and rock hardness and type of contact with overlying and underlying strata (see also the symbols' legend within the figure).



SOM Figure S5. Remains of *Hippotherium* sp. from the area of els Hostalets de Pierola: a–c) IPS124307, left M₃ in occlusal (a), lingual (b), and buccal (c) views; d–f) IPS30897, left M₂–M₃ in occlusal (d), lingual (e), and buccal (f) views. IPS30897 comes in all probability from the locality of Can Mata 3.



SOM Figure S6. Dental remains of Giraffidae indet. from els Hostalets de Pierola. a–f) IPS121859, right (a–c) and left (d–f) M³ germs from ACM/C13/C14-A (meter 283 of the ACM composite section; interpolated age of 11.11 Ma), in occlusal (a, d), lingual (b, e), and buccal (c, f) views. g–i) IPS88203, right P₂ from ECM/VCE-Bb (~10.9 Ma), in occlusal (g), buccal (h), and lingual (i) views. j–r) IPS88086, right P₄ (j–l), partial right M₁ (m–o), and right M₃ (p–r) from ECM/VCE-Bb (~10.9 Ma), in occlusal (j, m, p), buccal (k, n, q), and lingual (l, o, r) views.

SOM Table S1Magnetic susceptibility ($\times 10^{-6}$ SI) after each demagnetization step (in °C).

Sample	sus ini	100	200	250	300	350	400	440	480	520	550	580	610	630
BCM-1B	865	870	875	870	880	930	935	1015	955	905	820	—	—	—
BCM-2A	245	240	240	240	295	385	385	480	360	230	240	220	210	200
BCM3B	595	595	600	585	630	690	695	760	725	240	570	550	535	515
BCM4C	255	255	250	255	305	395	400	500	495	475	315	280	265	—
BCM-5B	560	560	560	550	570	620	625	760	775	310	725	625	590	490
BCM-6A	230	230	230	225	240	295	330	400	410	305	395	310	290	275
PTA-1.1A	430	430	430	420	420	430	455	470	455	845	375	—	—	—
PTA-2.1A	265	265	260	260	260	260	300	300	260	260	215	210	205	200
PTA-3.1A	250	250	250	250	245	245	340	355	285	605	220	215	215	205
PTA-4.1A	495	495	495	500	500	490	615	630	550	390	440	430	420	410
PTA-5.1A	330	330	330	330	330	330	435	450	370	765	290	280	275	265
PTA-6.1A	320	315	315	315	320	320	460	475	390	400	280	270	265	255
C13-1A	250	250	255	250	245	255	300	340	370	300	260	250	245	—
C13-2A	365	365	365	365	365	395	465	520	505	385	345	320	305	—
C13-3C	310	305	310	300	310	350	400	450	400	310	285	265	250	250
C13-4A	185	185	185	185	200	260	330	365	270	205	185	—	—	—
C13-5C	235	235	235	235	260	340	415	460	385	290	265	255	255	—
C13-6A	360	360	360	355	365	415	460	485	415	345	325	310	300	—
C13-7C	260	255	260	255	275	345	415	450	365	275	250	230	220	—
C13-8A	220	215	220	215	230	305	375	415	350	245	215	195	185	185
C13-9A	275	275	275	270	270	310	365	415	375	300	280	—	—	—
C13-10B	255	255	255	255	255	295	370	470	455	330	275	—	—	—
C13-11B	270	270	270	270	270	295	370	420	520	440	315	260	240	245
C13-12A	245	240	245	245	240	250	290	300	405	425	330	—	—	—
C13-13A.A	400	400	400	405	395	395	405	435	470	490	445	—	—	—
C13-13B	430	425	430	430	430	420	430	495	510	525	530	380	360	360
C13-14A.B	220	220	220	220	220	220	225	265	295	345	380	230	215	220
C13-14B-A	205	205	205	200	205	200	205	235	255	300	330	225	210	215
C13-15A	430	430	430	430	420	430	465	510	550	485	405	395	375	360
C13-16A	285	285	290	285	285	315	390	465	470	330	285	290	275	240
C13-17B	280	280	280	275	285	330	395	455	400	295	265	270	270	225
C13-18A	290	290	290	285	290	330	365	395	345	270	245	250	250	215
C13-19B	355	355	355	345	360	425	495	545	450	350	315	320	—	—

C13-20A	510	515	515	505	515	565	610	640	580	500	475	460	—	—
C13-21A	400	400	400	395	410	470	530	575	510	400	370	350	340	325
C13-22B	625	625	625	615	625	685	740	785	685	610	580	550	540	525
C13-23B	480	485	485	480	485	535	595	645	640	530	475	460	435	420
C13-24A	260	260	260	255	270	335	420	505	500	360	290	245	—	—
C13-25B	310	310	315	310	315	360	440	515	535	415	325	275	—	—
C13-26B	275	275	275	275	275	305	395	480	530	470	340	275	255	250
C13-27B	370	370	370	370	360	370	400	440	475	475	380	305	290	285
C13-28A	395	395	400	395	385	390	405	430	455	480	445	340	—	—
C13-29C	365	365	365	365	355	355	385	440	485	540	535	350	310	300
C13-30A	450	450	450	455	450	445	455	495	520	570	585	—	—	—
C13-31A	795	795	800	775	785	790	845	850	850	895	915	905	—	—
C13-32A	880	880	895	860	865	885	950	955	955	960	975	995	—	—

Abbreviation: sus ini = initial susceptibility; BCM = Bretxa de Can Mata [Can Mata Breccia]; C1–C13 = Cel·la 13 [Cells 13]; PTA = Préstec de Terres de l'Abocador [Landfill Earth Loan].

SOM Table S2

Directions of the characteristic components in geographic and stratigraphic coordinates calculated for each individual sample.

Sample	Level	Dg	Ig	Ds	Is	VGPlat	int	MAD	Q	T range
PTA1-1A	0.5	242.8	06.1	242.8	-05.5	-17.7	229.6	7.6	2	250–480
C13-1A	2.5	204.7	-62.5	178.4	-65.2	-71.9	260.3	6.9	2	300–520
PTA2-1A	5.0	203.2	06.3	204.0	01.1	-40.2	200.8	9.0	1	350–630
C13-2A	6.5	183.5	-39.1	172.9	-38.7	-69.9	732.4	3.2	2	250–480
C13-3C	7.5	167.9	-60.8	148.2	-55.8	-81.0	509.0	3.7	1	300–580
PTA3-1A	7.5	167.5	-55.0	151.5	-50.3	-78.3	202.0	14.3	2	* 350–550
C13-4A	9.5	230.2	-50.8	216.3	-59.8	-49.1	118.2	11.9	2	* 300–440
C13-5C	10.5	166.4	-37.5	157.6	-33.4	-66.1	297.5	6.4	2	250–520
C13-6C	11.5	177.9	-30.9	170.3	-29.5	-64.6	292.0	9.2	2	250–480
C13-7C	13.0	159.8	-46.0	148.8	-40.3	-68.0	289.7	4.6	1	250–520
C13-8A	14.0	163.9	-40.3	154.4	-35.6	-66.7	641.5	8.0	2	250–520
PTA4-1A	14.0	194.6	-45.8	180.9	-47.5	-71.0	3136.5	2.8	1	300–580
C13-9A	15.5	225.0	04.3	224.9	-04.9	-30.0	61.0	20.5	3	* 250–520
C13-10B	16.5	090.4	56.3	090.5	69.3	23.4	98.5	8.4	3	300–440
C13-11B	17.5	265.3	74.3	267.4	61.3	33.6	180.7	9.9	3	300–480
C13-12A	18.5	054.6	25.2	049.9	35.5	34.9	166.6	20.5	3	250–400
C13-13B	20.0	010.3	58.7	348.8	58.5	81.8	1334.5	4.7	1	250–580
PTA5-1A	20.5	035.5	58.8	013.4	64.2	63.2	1057.3	2.7	1	400–610
C13-14A	21.0	009.0	54.2	350.9	54.1	79.8	444.8	5.1	1	300–550
C13-15A	23.0	009.6	62.2	345.0	61.6	82.8	1808.0	4.4	1	300–580
C13-16A	27.5	017.7	42.5	005.5	45.1	67.2	421.8	6.0	1	300–550
C13-17B	30.0	026.8	17.4	022.5	22.8	49.4	724.7	2.8	1	250–520
PTA6-1A	30.5	155.9	-27.2	150.6	-21.3	-55.4	535.1	6.3	1	300–580
C13-18A	31.0	222.0	-47.9	208.4	-55.4	-53.8	434.3	6.4	2	300–550
C13-19B	34.0	227.3	-52.8	211.8	-61.1	-52.0	951.1	3.3	1	300–550
C13-20A	39.0	213.5	-31.8	205.6	-38.2	-52.1	1191.0	3.7	1	250–480
C13-21A	46.0	215.0	-39.8	204.4	-46.3	-55.0	243.4	7.2	2	250–480
C13-22B	47.5	207.8	-41.3	196.2	-46.2	-60.5	1624.0	4.6	1	250–520
C13-23B	48.0	192.7	-34.3	183.5	-36.1	-64.5	539.9	6.7	2	* 350–610
C13-24A	50.0	218.8	-27.4	212.3	-35.0	-46.8	138.3	21.3	2	250–520
C13-25B	50.5	265.8	-48.7	264.2	-61.6	-22.2	224.1	12.1	3	250–440
C13-26B	51.5	308.5	31.0	304.8	20.5	39.5	305.8	9.2	1	300–580

C13-27B	53.0	021.3	50.0	005.2	53.0	69.6	2583.0	1.4	1	250–550
C13-28a	56.0	235.6	25.4	238.2	14.5	-17.7	105.0	8.1	3	300–440
C13-29C	58.5	202.2	-30.7	194.2	-34.7	-58.2	455.8	8.3	2	400–550
C13-30A	59.0	199.4	-43.4	186.7	-46.3	-66.9	168.8	22.5	3	250–440
BCM-1B	62.0	147.6	-20.3	144.4	-12.9	-47.6	140.0	34.9	2	* 350–520
C13-32A	62.0	222.9	-46.8	210.0	-54.6	-52.6	678.7	18.4	3	250–440
BCM-2A	67.0	303.0	64.5	292.9	52.9	49.8	146.3	16.2	2	520–580
BCM-3B	70.0	206.4	-48.5	191.3	-52.7	-65.3	978.2	5.0	1	350–580
BCM-4C	72.0	341.4	37.2	333.1	32.1	63.6	97.1	11.5	3	250–400
BCM-5B	75.0	012.6	53.1	354.9	53.9	77.1	2162.6	2.8	1	300–580
BCM-6A	77.0	018.1	56.6	018.1	56.6	75.2	370.0	3.8	2	350–580

Abbreviations: level = stratigraphic position (m); Dg = declination in geographic coordinates; Ig = inclination in geographic coordinates; Ds = declination in stratigraphic coordinates; Is = inclination in stratigraphic coordinates; VGPlat = latitude of the virtual geomagnetic pole; int = intensity ($E\ 10^{-6}\ A/m$); MAD = maximum angular deviation; Q = quality; T range = temperature interval selected to calculate the characteristic component forced to origin, unless specified by *, which means not forced to origin; BCM = Bretxa de Can Mata [Can Mata Breccia]; C1–C13 = Cel·la 13 [Cells 13]; PTA = Préstec de Terres de l'Abocador [Landfill Earth Loan].

SOM Table S3

Mean directions and their Fisher statistics parameters. See also SOM Figure S2.

Polarity	<i>n</i>	Geographic coordinates				Bedding corrected coordinates			
		Dg	Ig	k	$\alpha 95$	Dg	Ig	k	$\alpha 95$
Normal	11	006.9	53.6	12.3	13.6	351.0	53.2	11.8	13.9
Reverse	23	194.6	-40.7	9.1	10.6	183.1	-42.7	9.1	10.6
All to normal	34	012.4	45.0	9.6	8.4	359.6	46.3	9.6	8.4

Abbreviations: *n* = number of samples; Dg = declination in geographic coordinates; Ig = inclination in geographic coordinates; k = precision parameter; $\alpha 95$ = alpha 95.

SOM Table S4

List of fossil localities at Abocador de Can Mata (ACM), in chronological order from youngest to oldest (top to bottom), updated from Alba et al. (2017). Their stratigraphic position within the composite series, their correlated geomagnetic subchron, and their interpolated and estimated ages are provided. Each locality is designated by the ACM acronym, followed by a slash, the acronym of each sector (named after the various cells and other structures of the landfill), sometimes a slash and an uppercase letter denoting a division in subsectors, and a few additional letters and/or numbers designating the particular locality. Primate-bearing localities are bolded. Alternative locality names are given within parentheses for the sake of completeness in relation to field reports and museum inventories. See SOM Table S5 for a list of ACM subsectors and their age ranges.

Locality	Stratigraphic position (m)	Subchron	Interpolated age (Ma)	Estimated age (Ma)
CM3 ^a	292	C5r.1r	11.06	11.1
CM1 ^a	284	C5r.1r	11.11	11.1
ACM/C13-A2 ^a	279	C5r.1r	11.14	11.1
ACM/C13-A3 ^a	278	C5r.1r	11.14	11.1
ACM/C13-A5 (=ACM/C13/C14-Ag) ^a	276	C5r.1n	11.16	11.2
ACM/C13/C14-Af ^a	275	C5r.1n	11.17	11.2
ACM/C13/C14-Ad ^a	274	C5r.1n	11.18	11.2
ACM/C13-A6 (=ACM/C13/C14-Ah)	270	C5r.2r [2/2]	11.20	11.2
ACM/C14-Ba	269	C5r.2r [2/2]	11.20	11.2
ACM/PTA-A1 (=ACM/PTA-Ac)	268	C5r.2r [2/2]	11.21	11.2
ACM/PTA-A2	268	C5r.2r [2/2]	11.21	11.2

ACM/C14-Bb	266	C5r.2r [2/2]	11.21	11.2
ACM/C14-B1	266	C5r.2r [2/2]	11.21	11.2
ACM/C13/C14-Aa	264	C5r.2r [2/2]	11.22	11.2
ACM/C13-A1 (=ACM/C13-Ab)	263	C5r.2r [2/2]	11.22	11.2
ACM/C13-A7	263	C5r.2r [2/2]	11.22	11.2
ACM/C13/C14-Ai	263	C5r.2r [2/2]	11.22	11.2
ACM/C13-Aa	262	C5r.2r [2/2]	11.23	11.2
ACM/C13/C14-Ab	262	C5r.2r [2/2]	11.23	11.2
ACM/C13/C14-Aj	260	C5r.2r [2/2]	11.23	11.2
ACM/C13-A4 (=ACM/C13/C14-Ae)	254	C5r.2r [2/2]	11.25	11.3
ACM/C12-A1(=ACM/C12-Ab)	252	C5r.2r [2/2]	11.26	11.3
ACM/PTA-Aa	251	C5r.2r-1	11.26	11.3
ACM/C12-Ba	248	C5r.2r-1	11.28	11.3
ACM/PTA-Ab	245	C5r.2r-1	11.29	11.3
ACM/PTA-Ad	243	C5r.2r-1	11.29	11.3
ACM/C13/C14-Ac	241	C5r.2r-1	11.30	11.3
ACM/PTA-A3 (=ACM/PTA-Ae)	232	C5r.2r [1/2]	11.36	11.4
ACM/C12-Aa	230	C5r.2r [1/2]	11.38	11.4
ACM/CCV1	221	C5r.2r [1/2]	11.45	11.5
ACM/C8-B1	212	C5r.2r [1/2]	11.53	11.5
ACM/C6-Cc	212	C5r.2r [1/2]	11.53	11.5

ACM/C8-Ca	211	C5r.2r [1/2]	11.54	11.5
ACM/C8-Bb	211	C5r.2r [1/2]	11.54	11.5
ACM/C6-Ca	211	C5r.2r [1/2]	11.54	11.5
ACM/C8-Bd	210	C5r.2r [1/2]	11.54	11.5
ACM/C8-Be	210	C5r.2r [1/2]	11.54	11.5
ACM/C6-Aa	209	C5r.2r [1/2]	11.55	11.6
ACM/C8-Ba	209	C5r.2r [1/2]	11.55	11.6
ACM/C8-Bi	209	C5r.2r [1/2]	11.55	11.6
ACM/C8-Br	207	C5r.2r [1/2]	11.57	11.6
ACM/C8-Bc	204	C5r.2n	11.59	11.6
ACM/C8-Ae	204	C5r.2n	11.59	11.6
ACM/C8-Bm	204	C5r.2n	11.59	11.6
ACM/C8-Bq	204	C5r.2n	11.59	11.6
ACM/C5-Ad	203	C5r.2n	11.59	11.6
ACM/C6-Ac	203	C5r.2n	11.59	11.6
ACM/C6-Ab	203	C5r.2n	11.59	11.6
ACM/C6-C1	203	C5r.2n	11.59	11.6
ACM/C8-Bn	203	C5r.2n	11.59	11.6
ACM/C8-Bo	203	C5r.2n	11.59	11.6
ACM/C8-Bh	203	C5r.2n	11.59	11.6
ACM/C8-Bs	203	C5r.2n	11.59	11.6

ACM/C6-C2	202	C5r.2n	11.60	11.6
ACM/C8-B3 (=ACM/C8-Bk)	202	C5r.2n	11.60	11.6
ACM/C8-Bp	202	C5r.2n	11.60	11.6
ACM/C8-Bv	202	C5r.2n	11.60	11.6
ACM/C8-By	202	C5r.2n	11.60	11.6
ACM/C8-Bf	201	C5r.2n	11.60	11.6
ACM/C8-Ac	201	C5r.2n	11.60	11.6
ACM/C8-Ad	201	C5r.2n	11.60	11.6
ACM/C6-C4	201	C5r.2n	11.60	11.6
ACM/C8-Bg	200	C5r.2n	11.60	11.6
ACM/C6-C3 (=ACM/C6-Cd)	200	C5r.2n	11.60	11.6
ACM/C8-B2 (=ACM/C8-Bj)	200	C5r.2n	11.60	11.6
ACM/C8-Ab	198	C5r.2n	11.61	11.6
ACM/C6-Ad	198	C5r.2n	11.61	11.6
ACM/C5-Ac	198	C5r.2n	11.61	11.6
ACM/C6-C5 (=ACM/C6-Ce)	197	C5r.2n	11.61	11.6
ACM/C6-A2	197	C5r.2n	11.61	11.6
ACM/C6-A1	197	C5r.2n	11.61	11.6
ACM/C8-Bt	197	C5r.2n	11.61	11.6
ACM/C8-Bd'	197	C5r.2n	11.61	11.6
ACM/C5-Dc	196	C5r.2n	11.61	11.6

ACM/C8-Aj	195	C5r.2n	11.61	11.6
ACM/C8-A4	194	C5r.2n	11.62	11.6
ACM/C5-A1 (=ACM/C5-Aa)	194	C5r.2n	11.62	11.6
ACM/C8-BI	194	C5r.2n	11.62	11.6
ACM/C8-B4 (=ACM/C8-Bw)	194	C5r.2n	11.62	11.6
ACM/C8-Bx	192	C5r.2n	11.62	11.6
ACM/C6-Cb	192	C5r.2n	11.62	11.6
ACM/C4-Ad	192	C5r.2n	11.62	11.6
ACM/C6-Ah	191	C5r.2n	11.62	11.6
ACM/C7-Ba	190	C5r.2n	11.63	11.6
ACM/C6-A4	190	C5r.2n	11.63	11.6
ACM/C5-Db	189	C5r.2n	11.63	11.6
ACM/C6-Af	189	C5r.2n	11.63	11.6
ACM/C5-Dm	188	C5r.2n	11.63	11.6
ACM/C5-D1 (=ACM/C5-Da)	188	C5r.2n	11.63	11.6
ACM/C4-Aa	188	C5r.2n	11.63	11.6
ACM/C5-Dg	187	C5r.2n	11.63	11.6
ACM/C5-D2 (=ACM/C5-De)	187	C5r.2n	11.63	11.6
ACM/C8-A3 (=ACM/C8-Ah)	187	C5r.2n	11.63	11.6
ACM/C8-Ai	187	C5r.2n	11.63	11.6
ACM/C8-A1	186	C5r.2n	11.64	11.6

ACM/C5-Dl	186	C5r.2n	11.64	11.6
ACM/C5-Dk	186	C5r.2n	11.64	11.6
ACM/C7-Aa	186	C5r.2n	11.64	11.6
ACM/C8-A2 (=ACM/C8-Aa)	185	C5r.2n	11.64	11.6
ACM/C7-Ab	185	C5r.2n	11.64	11.6
ACM/C8-Ak	185	C5r.2n	11.64	11.6
ACM/C8-Af	185	C5r.2n	11.64	11.6
ACM/C8-Ag	185	C5r.2n	11.64	11.6
ACM/C6-A5 (=ACM/C6-Ag)	184	C5r.2n	11.64	11.6
ACM/C5-D3	183	C5r.2n	11.64	11.6
ACM/C5-Dd	183	C5r.2n	11.64	11.6
ACM/C6-Ai	183	C5r.2n	11.64	11.6
ACM/C5-D4 (=ACM/C5-Df)	182	C5r.2n	11.65	11.6
ACM/C5-A4 (=ACM/C5-Ab)	181	C5r.2n	11.65	11.6
ACM/C8-Am	180	C5r.2n	11.65	11.7
ACM/C7-Bc	180	C5r.2n	11.65	11.7
ACM/C6-Ae	180	C5r.2n	11.65	11.7
ACM/C7-Ca	179	C5r.2n	11.65	11.7
ACM/C5-A7	179	C5r.2n	11.65	11.7
ACM/C8-B*	178.5	C5r.2n	11.65	11.7
ACM/C5-D7 (=ACM/C5-Dn)	178	C5r.2n	11.65	11.7

ACM/C5-A3	178	C5r.2n	11.65	11.7
ACM/C8-Bu	176	C5r.3r	11.66	11.7
ACM/C8-Ba'	176	C5r.3r	11.66	11.7
ACM/C8-Bc'	176	C5r.3r	11.66	11.7
ACM/C5-A5 (=ACM/C5-Ae)	174	C5r.3r	11.67	11.7
ACM/C6-A3	173	C5r.3r	11.67	11.7
ACM/C5-A6 (=ACM/C5-Af)	173	C5r.3r	11.67	11.7
ACM/C7-Bb	172	C5r.3r	11.68	11.7
ACM/C7-A1	171	C5r.3r	11.68	11.7
ACM/C7-A2	171	C5r.3r	11.68	11.7
ACM/C8-Bb'	171	C5r.3r	11.68	11.7
ACM/C8-Be'	171	C5r.3r	11.68	11.7
ACM/C8-Ar	170	C5r.3r	11.69	11.7
ACM/C8-AI	170	C5r.3r	11.69	11.7
ACM/C7-Bd	168	C5r.3r	11.70	11.7
ACM/C8-Bz	168	C5r.3r	11.70	11.7
ACM/C6-Aj	167	C5r.3r	11.70	11.7
ACM/C4-A1 (=ACM/C4-Ae)	166	C5r.3r	11.70	11.7
ACM/C5-D5 (=ACM/C5-Dh)	166	C5r.3r	11.70	11.7
ACM/C5-Di	165	C5r.3r	11.71	11.7
ACM/C8-As	164	C5r.3r	11.71	11.7

ACM/C4-Ac	163	C5r.3r	11.72	11.7
ACM/C8-Ap	162	C5r.3r	11.72	11.7
ACM/C8-Aq	160	C5r.3r	11.73	11.7
ACM/C4-A3	159	C5r.3r	11.73	11.7
ACM/C5-Dj	158	C5r.3r	11.74	11.7
ACM/C6-Ak	157	C5r.3r	11.74	11.7
ACM/C4-Ab	157	C5r.3r	11.74	11.7
ACM/C5-Ag	155	C5r.3r	11.75	11.8
ACM/C8-Av	154	C5r.3r	11.75	11.8
ACM/C8-Ao	153	C5r.3r	11.76	11.8
ACM/C4-A4	151	C5r.3r	11.77	11.8
ACM/C8-At	151	C5r.3r	11.77	11.8
ACM/C8-Au	151	C5r.3r	11.77	11.8
ACM/C5-D6	150	C5r.3r	11.77	11.8
ACM/C6-Bc	150	C5r.3r	11.77	11.8
ACM/C5-A8 (=ACM/C5-Ah)	150	C5r.3r	11.77	11.8
ACM/C6-Bb	149	C5r.3r	11.78	11.8
ACM/C6-Ba	148	C5r.3r	11.78	11.8
ACM/C5-A9 (=ACM/C5-Ai)	148	C5r.3r	11.78	11.8
ACM/C8-An	148	C5r.3r	11.78	11.8
ACM/RP2	148	C5r.3r	11.78	11.8

ACM/C6-Bd	146	C5r.3r	11.79	11.8
ACM/C5-C2 (=ACM/C5-Cb)	146	C5r.3r	11.79	11.8
ACM/C5-C1 (=ACM/C5-Ca)	146	C5r.3r	11.79	11.8
ACM/C4-Cb	145	C5r.3r	11.79	11.8
ACM/BPAa	143	C5r.3r	11.80	11.8
ACM/C4-Cc	142	C5r.3r	11.80	11.8
ACM/C4-C1 (=ACM/C4-Ca)	141	C5r.3r	11.81	11.8
ACM/C4-Af	141	C5r.3r	11.81	11.8
ACM/C5-Cd	137	C5r.3r	11.83	11.8
ACM/C3-AT	134	C5r.3r	11.84	11.8
ACM/C3-A1	133	C5r.3r	11.84	11.8
ACM/C2-A1	133	C5r.3r	11.84	11.8
ACM/C5-C4 (=ACM/C5-Ce)	133	C5r.3r	11.84	11.8
ACM/C5-Cf	133	C5r.3r	11.84	11.8
ACM/BNL-Ba	130	C5r.3r	11.86	11.9
ACM/C5-C3 (=ACM/C5-Cc)	129	C5r.3r	11.86	11.9
ACM/C4-Cd	128	C5r.3r	11.86	11.9
ACM/C3-Aa	126	C5r.3r	11.87	11.9
ACM/C3-A2	126	C5r.3r	11.87	11.9
ACM/C3-Ak	125	C5r.3r	11.88	11.9
ACM/C3-Af	125	C5r.3r	11.88	11.9

ACM/C3-Ae	125	C5r.3r	11.88	11.9
ACM/C4-Ap	124	C5r.3r	11.88	11.9
ACM/C3-A4	121	C5r.3r	11.89	11.9
ACM/C4-C2 (=ACM/C4-Ce)	120	C5r.3r	11.90	11.9
ACM/C3-A5 (=ACM/C3-Ac)	120	C5r.3r	11.90	11.9
ACM/C3-Ah	119	C5r.3r	11.90	11.9
ACM/C4-C3	118	C5r.3r	11.91	11.9
ACM/BCV5	117	C5r.3r	11.91	11.9
ACM/C4-Cf	116	C5r.3r	11.91	11.9
ACM/C3-A3	116	C5r.3r	11.91	11.9
ACM/C3-Ab	114	C5r.3r	11.92	11.9
ACM/C4-Cg	110	C5r.3r	11.94	11.9
ACM/C3-Az	110	C5r.3r	11.94	11.9
ACM/C3-A6 (=ACM/C3-Ad)	110	C5r.3r	11.94	11.9
ACM/BCV4	110	C5r.3r	11.94	11.9
ACM/VIE-C1	109	C5r.3r	11.94	11.9
ACM/C4-Cp	108	C5r.3r	11.95	11.9
ACM/VIE-C2	107	C5r.3r	11.95	12.0
ACM/BCV1 (=ACM/BDA-SW2, ACM/BDA-SW3)	105	C5r.3r	11.96	12.0
ACM/C2-A2	105	C5r.3r	11.96	12.0
ACM/C2-Aa	105	C5r.3r	11.96	12.0

ACM/C3-Ag	104	C5r.3r	11.97	12.0
ACM/C3-Aj	104	C5r.3r	11.97	12.0
ACM/NPLa	103	C5r.3r	11.97	12.0
ACM/BDA1	102	C5r.3r	11.97	12.0
ACM/C3-Ap	102	C5r.3r	11.97	12.0
ACM/BCV3 (=ACM/BDA-SW4)	101	C5r.3r	11.98	12.0
ACM/C3-Aq	100	C5r.3r	11.98	12.0
ACM/C2-A3	100	C5r.3r	11.98	12.0
ACM/C3-A1	98	C5r.3r	11.99	12.0
ACM/C3-An	98	C5r.3r	11.99	12.0
ACM/VIE-C3	98	C5r.3r	11.99	12.0
ACM/C3-A7	95	C5r.3r	12.00	12.0
ACM/VIE-C4	95	C5r.3r	12.00	12.0
ACM/BCV2 (=ACM/BDA-SW1)	90	C5r.3r	12.02	12.0
ACM/C3-Ai	89	C5r.3r	12.03	12.0
ACM/BDA2	88	C5r.3r	12.03	12.0
ACM/BDA7 (=ACM/BDAb)	87	C5r.3r	12.04	12.0
ACM/C2-A4	83	C5An.1n	12.06	12.1
ACM/C3-B2 (=ACM/C3-Bb)	83	C5An.1n	12.06	12.1
ACM/C3-Am	82	C5An.1n	12.07	12.1
ACM/C3-Ba	80	C5An.1n	12.08	12.1

ACM/C1-A1	80	C5An.1n	12.08	12.1
ACM/BDA8 (=ACM/BDAA, ACM/BDAC)	77	C5An.1n	12.11	12.1
ACM/BDA3	77	C5An.1n	12.11	12.1
ACM/C1-A2	76	C5An.1n	12.11	12.1
ACM/C2-Ba	75	C5An.1n	12.12	12.1
ACM/C2-B1	75	C5An.1n	12.12	12.1
ACM/C2-Bb	75	C5An.1n	12.12	12.1
ACM/C2-B2	73	C5An.1n	12.14	12.1
ACM/C1-A3	72	C5An.1n	12.15	12.1
ACM/C1-A5	70	C5An.1n	12.16	12.2
ACM/C1-A4	70	C5An.1n	12.16	12.2
ACM/C2-B3	69	C5An.1n	12.17	12.2
ACM/C3-Bc	67	C5An.1r	12.18	12.2
ACM/C2-Bc	66	C5An.1r	12.19	12.2
ACM/BDAd	65	C5An.1r	12.19	12.2
ACM/C3-Bd	58	C5An.1r	12.22	12.2
ACM/C1-E4	51	C5An.1r	12.25	12.3
ACM/C3-B3	47	C5An.2n	12.27	12.3
ACM/C1-E10	46	C5An.2n	12.28	12.3
ACM/C1-E9 (=ACM/C1-Ef)	45	C5An.2n	12.29	12.3
ACM/BDA4	42	C5An.2n	12.31	12.3

ACM/C1-C1	39	C5An.2n	12.32	12.3
ACM/C1-Ed	38	C5An.2n	12.33	12.3
ACM/C1-Ea	38	C5An.2n	12.33	12.3
ACM/C1-E6	38	C5An.2n	12.33	12.3
ACM/C1-Eb	36	C5An.2n	12.34	12.3
ACM/C1-E5	36	C5An.2n	12.30	12.3
ACM/C1-E3	36	C5An.2n	12.30	12.3
ACM/C1-Fa	35	C5An.2n	12.35	12.4
ACM/C1-C3	34	C5An.2n	12.36	12.4
ACM/C1-E*	~34–46	C5An.2n	12.36–12.28	12.4–12.3
ACM/C1-E2	32	C5An.2n	12.37	12.4
ACM/C1-C4	30	C5An.2n	12.38	12.4
ACM/C1-Ee	30	C5An.2n	12.30	12.4
ACM/C1-E7+8 (=ACM/C1-Ec)	29	C5An.2n	12.39	12.4
ACM/VIE-E1 (=ACM/VIE-Ea)	29	C5An.2n	12.39	12.4
ACM/C1-E1	28	C5An.2n	12.40	12.4
ACM/C1-ET2	26	C5An.2n	12.41	12.4
ACM/C1-ET	26	C5An.2n	12.41	12.4
ACM/C10-A1	25	C5An.2n	12.42	12.4
ACM/C1-Ex	25	C5An.2n	12.42	12.4
ACM/C1-D1	20	C5An.2n	12.45	12.4

ACM/C9-A1	19	C5An.2n	12.45	12.5
ACM/BDL3	16	C5An.2n	12.47	12.5
ACM/BDL2	15,5	C5Ar.1r	12.48	12.5
		C5Ar.1r–		
ACM/BDAe	~11–77	C5An.1n	12.51–12.11	12.5–12.1
ACM/BDA6	11	C5Ar.1r	12.51	12.5
ACM/BDL1	9	C5Ar.1r	12.52	12.5
ACM/BDA5	2	C5Ar.1r	12.56	12.6

Abbreviations: ACM = Abocador de Can Mata [Can Mata Landfill]; BCV = Barranc de Can Vila [Can Vila Ravine]; BDA = Bassa de Decantació d'Aigües Pluvials [Settling Pond of Rainwater]; BDA-SW = Bassa de Decantació d'Aigües Pluvials Sud-Oest [Settling Pond of Rainwater South-West]; BDL = Bassa de Lixiviats [Pond of Leachates]; BNL = Bassa Nova de Lixiviats [New Pond of Leachates]; BPA = Base de la Pila d'Acopi [Base of the Stockpile]; C1–C13 = Cel·les 1–13 [Cells 1–13]; CCV = Camí de Can Vila [Can Vila's Trackway]; CM = Can Mata; NPL = Nova Planta de Lixiviats [New Pond of Leachates]; PTA = Préstec de Terres de l'Abocador [Landfill Earth Loan]; RP = Rasa de Desguàs d'Aigües Pluvials [Rainwater Drainage Ditch]; VIE = Vial Intern d'Explotació [Internal Operating Road].

^a Vallesian localities according to the earliest appearance of *Hippotherium* in the Vallès-Penedès Basin (Garcés et al., 1996).

SOM Table S5

List of fossiliferous (sub)sectors at Abocador de Can Mata (ACM), in chronological order from youngest to oldest (top to bottom). Their stratigraphic range within the composite series, their correlated geomagnetic subchron(s), and their interpolated and estimated age ranges are provided. Each sector is designated by the ACM acronym, followed by a slash, the acronym of each sector (named after the various cells and other structures of the landfill), and sometimes a slash and an uppercase letter denoting a division in subsectors. Alternative (sub)sector names are given within parentheses for the sake of completeness in relation to field reports and museum inventories. See SOM Table S4 for an updated list of ACM fossil localities.

(Sub)sector	Stratigraphic position (m)	Subchron	Interpolated age (Ma)	Estimated age (Ma)
ACM/C14-B ^a	268–286	C5r.2r [2/2]–C5r.1r	11.21–11.10	11.2–11.1
ACM/PTA-B ^a	268–286	C5r.2r [2/2]–C5r.1r	11.21–11.10	11.2–11.1
ACM/C12-B ^a	245–286	C5r.2r-1–C5r.1r	11.29–11.10	11.3–11.1
ACM/C13-A ^a	238–286	C5r.2r [2/2]–C5r.1r	11.31–11.10	11.3–11.1
ACM/C13/C14-A (=ACM/C14-A) ^a	238–286	C5r.2r [1/2]–C5r.1r	11.31–11.10	11.3–11.1
ACM/PTA-A	224–271	C5r.2r [1/2]–C5r.2r [2/2]	11.43–11.19	11.4–11.2
ACM/C12-A	223–269	C5r.2r [1/2]–C5r.2r [2/2]	11.44–11.20	11.4–11.2
ACM/CCV	195–234	C5r.2n–C5r.2r [1/2]	11.61–11.35	11.6–11.4
ACM/C6-Camí	195–234	C5r.2n–C5r.2r [1/2]	11.61–11.35	11.6–11.4
ACM/RPA-2	72–234	C5An.1n–C5r.2r [1/2]	12.15–11.35	12.2–11.4
ACM/RP	121–225	C5r.3r–C5r.2r [1/2]	11.89–11.42	11.9–11.4
ACM/C6-C	194–222	C5r.2n–C5r.2r [1/2]	11.62–11.45	11.6–11.5
ACM/C8-C	210–221	C5r.2r [1/2]	11.54–11.45	11.5

ACM/C8-B	166–221	C5r.3r–C5r.2r [1/2]	11.70–11.45	11.7–11.5
ACM/VIE-A	166–221	C5r.3r–C5r.2r [1/2]	11.70–11.45	11.7–11.5
ACM/C5-A	138–220	C5r.3r–C5r.2r [1/2]	11.81–11.46	11.8–11.5
ACM/C6-A	143–209	C5r.3r–C5r.2r [1/2]	11.80–11.55	11.8–11.6
ACM/C8-A	141–207	C5r.3r–C5r.2r [1/2]	11.81–11.57	11.8–11.6
ACM/C11-A	141–207	C5r.3r–C5r.2r [1/2]	11.81–11.57	11.8–11.6
ACM/C11-B	141–207	C5r.3r–C5r.2r [1/2]	11.81–11.57	11.8–11.6
ACM/VIE-B	141–207	C5r.3r–C5r.2r [1/2]	11.81–11.57	11.8–11.6
ACM/C5-D	150–200	C5r.3r–C5r.2n	11.77–11.60	11.8–11.6
ACM/C4-A	107–195	C5r.3r–C5r.2n	11.95–11.61	12.0–11.6
ACM/C7-A	168–190	C5r.3r–C5r.2n	11.70–11.63	11.7–11.6
ACM/C7-B	168–185	C5r.3r–C5r.2n	11.70–11.64	11.7–11.6
ACM/C7-C	175–183	C5r.3r–C5r.2n	11.67–11.64	11.7–11.6
ACM/C6-B	136–160	C5r.3r	11.83–11.73	11.8–11.7
ACM/C6-B	136–160	C5r.3r	11.83–11.73	11.8–11.7
ACM/C5-C	126–158	C5r.3r	11.87–11.74	11.9–11.7
ACM/C3-A	79–153	C5An.1n–C5r.3r	12.08–11.76	12.1–11.8
ACM/BPA	106–151	C5r.3r	11.96–11.77	12.0–11.8
ACM/C2-A	55–150	C5An.1r–C5r.3r	12.24–11.77	12.2–11.8
ACM/BNL-B	110–145	C5r.3r	11.94–11.79	11.9–11.8
ACM/BNL-A/B	98–145	C5r.3r	11.99–11.79	12.0–11.8

ACM/VIE-C	88–145	C5r.3r	12.03–11.79	12.0–11.8
ACM/C4-C	80–145	C5An.1n–C5r.3r	12.08–11.79	12.1–11.8
ACM/C4-B	80–142	C5An.1n–C5r.3r	12.08–11.80	12.1–11.8
ACM/C5-B	135–140	C5r.3r	11.83–11.81	11.8
ACM/BCV (=ACM/BDA-SW)	64–122	C5An.1r–C5r.3r	12.20–11.89	12.2–11.9
ACM/C1-A	70–113	C5An.1n–C5r.3r	12.16–11.93	12.2–11.9
ACM/C1-B	50–113	C5An.1r–C5r.3r	12.26–11.93	12.3–11.9
ACM/BNL-A	98–110	C5r.3r	11.99–11.94	12.0–11.9
ACM/C2-Base	55–109	C5An.1r–C5r.3r	12.24–11.94	12.2–11.9
ACM/NPL	62–107	C5An.1r–C5r.3r	12.20–11.95	12.2–12.0
ACM/C3-A/B	79–104	C5An.1n–C5r.3r	12.08–11.97	12.1–12.0
ACM/C3-B	40–104	C5An.2n–C5r.3r	12.32–11.97	12.3–12.0
ACM/BDA	2–102	C5Ar.1r–C5r.3r	12.56–11.97	12.6–12.0
ACM/C2-B	64–78	C5An.1r–C5An.1n	12.20–12.10	12.2–12.1
ACM/C1-E	23–51	C5An.2n–C5An.1r	12.43–12.25	12.4–12,3
ACM/C1-F	33–46	C5An.2n	12.36–12.28	12.4–12.3
ACM/C1-C	28–43	C5An.2n	12.40–12.30	12.4–12.3
ACM/VIE-E	25–35	C5An.2n	12.42–12.35	12.4
ACM/C10-A	8–33	C5Ar.1r–C5An.2n	12.52–12.36	12.5–12.4
ACM/BDL	0–25	C5Ar.1r–C5An.2n	12.58–12.42	12.6–12.4
ACM/C1-D	18–23	C5An.2n	12.46–12.43	12.5–12.4

ACM/C9-A	17–23	C5An.2n	12.47–12.43	12.5–12.4
ACM/VIE-D	17–23	C5An.2n	12.47–12.43	12.5–12.4
ACM/RBL	10–20	C5Ar.1r–C5An.2n	12.51–12.45	12.5

Abbreviations: ACM = Abocador de Can Mata [Can Mata Landfill]; BCV = Barranc de Can Vila [Can Vila Ravine]; BDA = Bassa de Decantació d'Aigües Pluvials [Settling Pond of Rainwater]; BDA-SW = Bassa de Decantació d'Aigües Pluvials Sud-Oest [Settling Pond of Rainwater South-West]; BDL = Bassa de Lixiviats [Pond of Leachates]; BNL = Bassa Nova de Lixiviats [New Pond of Leachates]; BPA = Base de la Pila d'Acopi [Base of the Stockpile]; C1–C13 = Cel·les 1–13 [Cells 1–13]; CCV = Camí de Can Vila [Can Vila's Trackway]; NPL = Nova Planta de Lixiviats [New Pond of Leachates]; PTA = Préstec de Terres de l'Abocador [Landfill Earth Loan]; RBL = Rasa de la Bassa de Lixiviats [Pond of Leachates Ditch]; RP = Rasa de Desguàs d'Aigües Pluvials [Rainwater Drainage Ditch]; RPA-2 = Rasa Perimetral de L'Abocador 2 [Perimetric Ditch of the Landfill 2]; VIE = Vial Intern d'Explotació [Internal Operating Road].

^a Sectors recording the Aragonian/Vallesian boundary according to the earliest appearance of *Hippotherium* in the Vallès-Penedès Basin (Garcés et al., 1996).

SOM Table S6

List of primate-bearing fossil localities at Abocador de Can Mata (ACM), in chronological order from youngest to oldest (top to bottom), updated from Alba et al. (2017). Their interpolated and estimated ages are provided (for further details, see SOM Table S4).

Locality	Taxon	Interpolated age (Ma)	Estimated age (Ma)
CM1	Dryopithecinae indet.	11.11	11.1
ACM/C8-A4	<i>Pliobates cataloniae</i>	11.62	11.6
ACM/C5-D1	<i>Pliobates cataloniae</i> , Dryopithecinae indet.	11.63	11.6
ACM/C8-B*	Dryopithecinae indet.	11.65	11.7
ACM/C4-A1	<i>Pliopithecus canmatensis</i>	11.70	11.7
ACM/C8-Au	Dryopithecinae indet.	11.77	11.8
ACM/C5-A8	<i>Pliopithecus canmatensis</i>	11.77	11.8
ACM/C5-C2	<i>Pliopithecus canmatensis</i>	11.79	11.8
ACM/C4-Cb	<i>Pliopithecus canmatensis</i>	11.79	11.8
ACM/C5-C4	<i>Pliopithecus canmatensis</i>	11.84	11.8
ACM/C5-C3	<i>Pliopithecus canmatensis</i>	11.86	11.9
ACM/C3-Ae	<i>Dryopithecus fontani</i>	11.88	11.9
ACM/C4-Ap	<i>Dryopithecus fontani</i>	11.88	11.9
ACM/BCV5	<i>Pliopithecus canmatensis</i>	11.91	11.9
ACM/C3-Az	cf. <i>Dryopithecus fontani</i>	11.94	11.9
ACM/BCV4	' <i>Sivapithecus</i> ' <i>occidentalis</i> species inquirenda	11.94	11.9
ACM/C4-Cp	<i>Anoiapithecus brevirostris</i>	11.95	11.9

ACM/BCV1	<i>Pierolapithecus catalaunicus</i>	11.96	12.0
ACM/C3-Aj	<i>Anoiapithecus brevirostris</i>	11.97	12.0
ACM/C3-B2	Pliopithecoida indet.	12.06	12.1
ACM/C1-E*	<i>Anoiapithecus brevirostris</i>	12.36–12.28	12.4–12.3

Abbreviations: See SOM Table S4.

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